



CONFEX: A Database for CONUS Fire EXtent

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Abstract. This article presents the CONUS Fire EXtent (CONFEX) database. This database, based on the VIIRS S-NPP 375m data product, provides wildfire perimeter, centroid, ignition location, start and end date for the period 2012 – 2023, for the CONUS and Alaska regions. The algorithm takes hotspot locations from VIIRS S-NPP, clusters them into actual wildfires based on DBSCAN clustering and calculates the perimeter and centroid of each cluster, attaching a geodata frame to each cluster or fire. When validated for some recent large fires against the CALFIRE database, an F1 score of 85 -96% and a CSI of 74 – 93 % were found, showing the efficiency of the algorithm in aggregating hotspots spatially and temporally accurately. This is the first publicly available high-resolution wildfire extent dataset developed for the CONUS and Alaska regions using VIIRS S-NPP 375m data product. The database provides a valuable resource for researchers to understand the complexities of the fire regimes in the CONUS and Alaska regions.

1. Introduction

Wildfire is nature's self-cleansing ritual and a natural part of the carbon cycle (Santín et al., 2015). Natives living in forested regions understood the importance of wildfires and for centuries implemented cultural burning as a duty towards the land that sustained them (Greenler et al., 2024). Overtaking of the native populations led to the exploitation of the forests (Marks-Block & Tripp, 2021). The implementation of fire suppression policies, curtailment of indigenous fire stewardship, mass deforestation of fire-resistant tall trees for fuel purposes, ever increasing thirst for inhabiting wildlands, and on top of that climate change, has led to the current fire pandemic (Martinez et al., 2023; Copes-Gerbitz et al., 2024). In the last two decades fires have raged wrath in the western fire regime of the Continental United States, and their number and severity are expected to increase (Westerling, 2016; Harvey, 2016; Westerling et al., 2006; Iglesias et al., 2022; Holden et al., 2018; Brown et al., 2004). Wildfires have a massive impact on the biosphere. There is a positive feedback loop between climate change and wildfires, with the latter leading to the release of more greenhouse gases, like CO₂ and NO_x, into the atmosphere from biomass burning further aggravating the already hot planet (Barbero et al., 2014). The permafrost thaw and the wildfire regime in the boreal forests create a positive feedback loop, with the former, caused by a changing climate, leading to drier and warmer near surface conditions, increasing fire propagation capacity, and the latter leading to ground subsidence, thickening of the organic content of surface soil layer, changes in the vegetation from coniferous to broad leafed, all leading to increased permafrost thaw (Kim et al., 2024; Kornei, 2020; Li et al., 2021). Massive fires have also led to the intrusion of aerosol into the stratosphere which is causing a slow but steady ozone depletion (Hirsch & Koren, 2021; Ma et al., 2024). At the surface, wildfires destroy the top layer of the soil leading to enhanced erosion and runoff, increasing the probability of flooding events (Wieting et al., 2017).

In this dynamic and ever worsening cycle, researchers have been trying to understand this force of nature, with a massive increase in studies related to wildfires in the last decade. Three prominent areas in this research are predicting the regions ripe for wildfires, especially at the urban wildland interface and along utility distribution lines (Taylor et al., 2024; Barzani et al., 2024; Syphard et al., 2024), real-time wildfire monitoring (Tong et al., 2024; Zhao & Ban, 2025; Briley & Afghah, 2024), and modelling the propagation of an already ignited fire, to effectively dispose of suppression resources (Hu et al., 2024; Sadrabadi & Innocente, 2024; Pereira et al., 2024). These fields need accurate, detailed and global datasets which can be used to study not only the spatial distribution of these fires but also the temporal and spatial evolution of these fires, given the topographic expanse.



45 Wildfire related data can be acquired through manual or remote sensing methods. Manual methods primarily consist
 46 of firefighting units monitoring fires, and of forest service departments surveying burnt areas post fire. After
 47 monitoring and surveying, wildfire reports are created based on that information and are made available for public
 48 knowledge. The information recorded may include characteristics like probable cause of fire, amount of area burnt,
 49 and the extent of damage inflicted (Porter et al., 2019; Brown et al., 2002; Short, 2022). Manual data has been the
 50 primary mode for centuries, and for most fires around the urban centers can be reliable in terms of its details about the
 51 location of ignition, possible reason, initial mode of spread, and the level and type of discomfort and damage. This
 52 provides valuable knowledge for preparing for future fires, through development of more resilient urban infrastructure.
 53 However, this data is most accurate at urban or urban wildland interfaces, since that has been the priority for centuries
 54 and rightly so. One of such databases is Shorts (2022). This database is one of the most comprehensive ground-based
 55 datasets for CONUS, spanning across 28 years from 1992 -2020. The fire entries are manually submitted by local,
 56 state and federal organizations.

57 Remote sensing methods include satellite and low elevation airborne instruments. Broadband spectral sensing and
 58 hyperspectral sensing are the two competing surveillance techniques in use currently (Barnpoutis et al., 2020;
 59 Dennison & Roberts, 2009; Veraverbeke et al., 2018). Hyperspectral sensing employs a large number (~100) of minute
 60 spectral bands ($< 0.002 \mu\text{m}$) which overlap, while broadband spectral sensing employs a smaller number (3 – 15)
 61 relatively larger ($> 0.002 \mu\text{m}$) spread out spectral bands (Veraverbeke et al., 2018). For active fire detection both reside
 62 in the short-wave near infrared ($1.2 - 2.5 \mu\text{m}$) to mid infrared range ($3 - 5 \mu\text{m}$), where we have a noninterference
 63 window, while the nearest infrared ($0.7 - 1.2 \mu\text{m}$) and visible range ($0.4 - 0.7 \mu\text{m}$) are not useful due to smoke
 64 interference (Veraverbeke et al., 2018). Another spectral range utilized for active fire detection is the thermal infrared
 65 band ($8 - 12 \mu\text{m}$), due to significant fire emissions and the presence of a window in the atmospheric absorption
 66 spectrum, however it is close to the earth's average temperature emission peak ($\sim 9.7 \mu\text{m}$). Burnt area mapping is
 67 usually done by utilizing the differences in the land cover and moisture differences, and for this microwave spectral
 68 region can be utilized especially for negating cloud interferences. Infrared and visible regions are widely employed
 69 and are very efficient except for cloud interference (Leblon et al., 2016).

70 Airborne instruments have been employed for wildfire detection and monitoring since the 1960s. The AAS/5 scanner,
 71 onboard the Beechcraft AT 11 aircraft provided useful fire surveillance imagery as early as 1962 (Warren & Wilson,
 72 1981). In 1998 Airborne Visible/Infrared Imaging Spectrometer (AVIRIS) was developed by the Jet Propulsion
 73 Laboratory, California Institute of Technology, which has been employed for wildfire surveillance through
 74 hyperspectral imaging (Green et al. 1998). In 2008, Airborne Prism experiment (APEX) instrument, another
 75 hyperspectral marvel, was developed and has also been employed to the same end (Itten et al. 2008). Since then,
 76 development in optics and electronics has led to much more sophisticated instrumentation been applied for wildfire
 77 surveillance, techniques like 4π -sr Spectroradiometry, MODIS/ASTER Airborne Simulator (MASTER) scanning
 78 spectrometer, and Differential Absorption Lidar - High Spectral Resolution Lidar (DIAL-HSRL), used in the FIREX
 79 – AQ mission in 2019 (Warneke et al., 2023). The airborne instruments have high spatial resolution but can only be
 80 employed for specific missions because of their limited spatial coverage, the need for a carrier and constant
 81 maintenance, making them unfeasible for long-term or large area surveillance. Tower mounted cameras at optimized
 82 locations based on terrain in forests for early fire detection has also been a very useful tool (Heyns et al., 2019; Bao
 83 et al., 2015). Manual, tower mounted camera-based, and airborne instrument-based monitoring of fire for active fire
 84 management are critical since the revisit time of spaceborne instruments is still too large for such advisory (Leblon et
 85 al., 2016; Allison et al., 2016).

86 Satellite fire surveillance came into use a decade and a half after the beginning of the era of meteorological satellites.
 87 The quality of satellite-based fire data depends on the orbital type, the instrument used, and the accuracy of the
 88 algorithm converting the received signal into actual values. Wildfire-related satellite data has evolved through the
 89 years concurrently with the development of satellite technology and computing power. Dozier (1981) provided the
 90 basis of the earliest method of extracting remote sensing wildfire data from the passive Advanced Very High-
 91 Resolution Radiometer (AVHRR) imagery ($3.55 - 3.83 \mu\text{m}$ and $10.5 - 11.5 \mu\text{m}$ channels) onboard the National Oceanic
 92 and Atmospheric Administration (NOAA) managed, NOAA-6 satellite of the Television Infrared Observation
 93 Satellite-N series (TIROS-N), facilitated by the presence of two channels in the thermal infrared (IR), primarily
 94 intended for sea surface temperature measurements (Nordin & Aziz, 2007). Dozier hypothesized based on the Planck's



radiation law that the temperature difference acquired from the radiation sensed at two different channels at subpixel level is because of varying surface temperature within the pixel rather than interference in the atmosphere and thus can be employed to extract the magnitude and the subpixel spatial fractions of the two temperatures. However, he assumed that there were only two sub-pixel temperatures, a known base (acquired from non-fire pixels), and a target, and this method worked well for unsaturated pixels. However, AVHRR's relatively high resolution meant a lower saturation threshold (450 K over 0.02 Km² field of view (FOV)), hindering large fire monitoring, and the NOAA-6 had a single pass each day, making it difficult to track fire progress (Prins & Menzel, 1992). The same algorithm was applied to the imagery (3.9 and 11.2 μ m channels) from the Visible Infrared Spin Scan Radiometer Atmospheric Sounder (VAS) on board the NOAA's Geostationary operational environmental satellite (GOES). The coarser resolution (13.8 Km for 3.9 μ m and 6.9 km for 11.2 μ m), meant that even though it missed small burns, it performed better for large fires due to higher saturation limit (450 K over 5 km² FOV). Moreover, as early as 1983, it proved to be a useful tool for monitoring the temporal evolution of wildfires thanks to the hourly data availability (Prins & Menzel, 1992; VAS (1980), n.d.). However, the pixels were still manually checked for fire presence. In 1994 the Cooperative Institute for Meteorological Satellite Studies (CIMSS) developed an automated biomass burning algorithm (ABBA) to classify the VAS pixels for fire, essentially for monitoring basin wide biomass burning over South America, while still relying on the Dozier (1981) method (Prins & Menzel, 1994). The temporal resolution was improved to 15 minutes and the IR resolution was improved to 4 km with the 1994 launch of GOES-8 (Kaufman et al., 1998). The Block 5 Defense meteorological Satellite program (DMSP) satellites were used for providing nighttime monitoring of fires at a 2.7 Km resolution in the 0.4 – 1.1 μ m range, as early as 1992 to monitor African savanna burns. However, due to lack of data access the quantification of area was not possible (Cahoon et al., 1992). More detailed information on the early evolution of fire detection techniques using AVHRR onboard NOAA 6 – 14 is provided in the article written by Kaufman et al. (1998). Giglio and Kendall (2001) found that the Dozier algorithm led to $\pm 50\%$ random error in detecting burnt area and ± 100 K random error in detecting temperature, under sensitivity analysis of fires $> 0.05\%$ of the pixel size.

The Moderate Resolution Imaging Spectroradiometer (MODIS) instrument was launched onboard the Earth observing mission (EOM) satellites, terra, in 1999, and aqua, in 2002 (Xiong et al., 2015). The MODIS instrument has 4 special fire channels 2 each for 4 μ m and 11 μ m, with 500 m spatial resolution, and the algorithm used for its data quantification is provided by Giglio et al. (2003, 2016, 2021). The original MODIS fire detection algorithm, the version 3 or Collection 3, was provided by Kaufman et al. (1998) which suffered high false alarm rate in deserts, or regions of sparse vegetation, and omission of smaller fires. In response to which Collection 4 algorithm was developed by Giglio et al. (2003). Collection 5 product was launched in mid-2006, it has been seen that the threshold for fire detection is high when considered over various terrains and regions and omits fires in some regions (Boschetti et al., 2009). The 310 K 4 μ m threshold causes most of the omissions. Even the Collection 5 algorithm struggled with omission of large fires due to thick smoke and falsely generated alarms due to clearing of forests. The Collection 6 and 6.1 algorithm addresses those issues. Offshore gas flaring was also incorporated in the Collection 6 algorithm (Giglio et al. 2016; 2021). The Landsat 7 imagery and Advanced Spaceborne Thermal Emission and Reflection Radiometer (ASTER) imagery are used for validation of the MODIS product. The brightness temperatures from 1 km MODIS channels for 12 μ m, 11 μ m, and 4 μ m are used for deriving the fire product. For sunlight hours, aggregation to 1 km of reflectance of 2.1 μ m, 0.85 μ m and 0.65 μ m are employed. The swath imagery is checked for fire pixels by the algorithm and each pixel is defined as unknown, fire, non-fire, cloud or missing data.

The Visible Infrared Imaging Radiometer Suite (VIIRS) instrument was an upgrade to NASA's AVHRR instrument launched on October 2011 onboard the Suomi National Polar-Orbiting Partnership (S-NPP) and then on NOAA-20 in 2017 and NOAA-21 satellites in November 2022, as part of the EOM (Justice et al., 2013; Cao et al., 2018; Xiong & Butler, 2020; Xiong et al., 2023). The VIIRS has 22 moderate 750 m resolution M bands, and 6 high 375 m resolution I bands. The temporal resolution is 12 h or less, and the threshold for fire detection at nighttime is as low as 5 m² (Schroeder et al., 2014).

VIIRS 750 m channel active fire application related product (ARP) was built on the MODIS Collection 4 algorithm, applying the same tests to the VIIRS 750 m multi-spectrum (Justice et al. 2013). In the subsequent upgrades, a pixelated parameter for fire characterization and 2D image classifier were added to the VIIRS 750 m product (Schroeder et al. 2014). VIIRS 750 m product is based on the 4 μ m and 11 μ m channel data. A new nighttime algorithm



145 uses the Day-Night Band (DNB) and 1.61 μm channel for sub pixel high temperature and gas flare detection at night
146 (Elvidge et al., 2013).

147 The VIIRS 375m channels are also used for fire -detection. The product algorithm processes five I band, and one M
148 band given in figure 1. The five channels of I band extend from the mid to thermal infrared region and are thus used
149 for fire detection. I4 (3.55 – 3.93 μm) is the main channel for fire detection and is used to distinguish between subpixel
150 fires and non-fires, with a 367 K saturation temperature (Schroeder et al. 2014). This channel is 0.3 μm below the
151 M13 channel (3.973 – 4.128 μm) that corresponds to the VIIRS 750 m ARP. Channel I5 (10.5 – 12.4 μm) is the thermal
152 channel against which the I4 channel checks for non-fires. The rest of the three channels are used in the algorithm for
153 water body discrimination, sun glint and clouds. For validation of VIIRS 375 imagery, Landsat- 8 data is used.

Channel	Spatial Resolution (m)	Spectral resolution (μm)	Primary Use
I1	375	0.60 – 0.68	Cloud & water classification
I2	375	0.846 – 0.885	Cloud & water classification
I3	375	1.58 – 1.64	Water classification
I4	375	3.55 – 3.93	Fire detection
I5	375	10.5 – 12.4	Fire detection & cloud classification
M13*	750	3.973 – 4.128	FRP retrieval, fire detection over water and across the South Atlantic magnetic anomaly region

154 * Aggregated (750×750 m nominal) & un-aggregated (250×750 m nominal) data are used

155 **Figure 1: Channels used in the 375 m active fire detection algorithm (adapted from Schroeder & Giglio, 2016)**

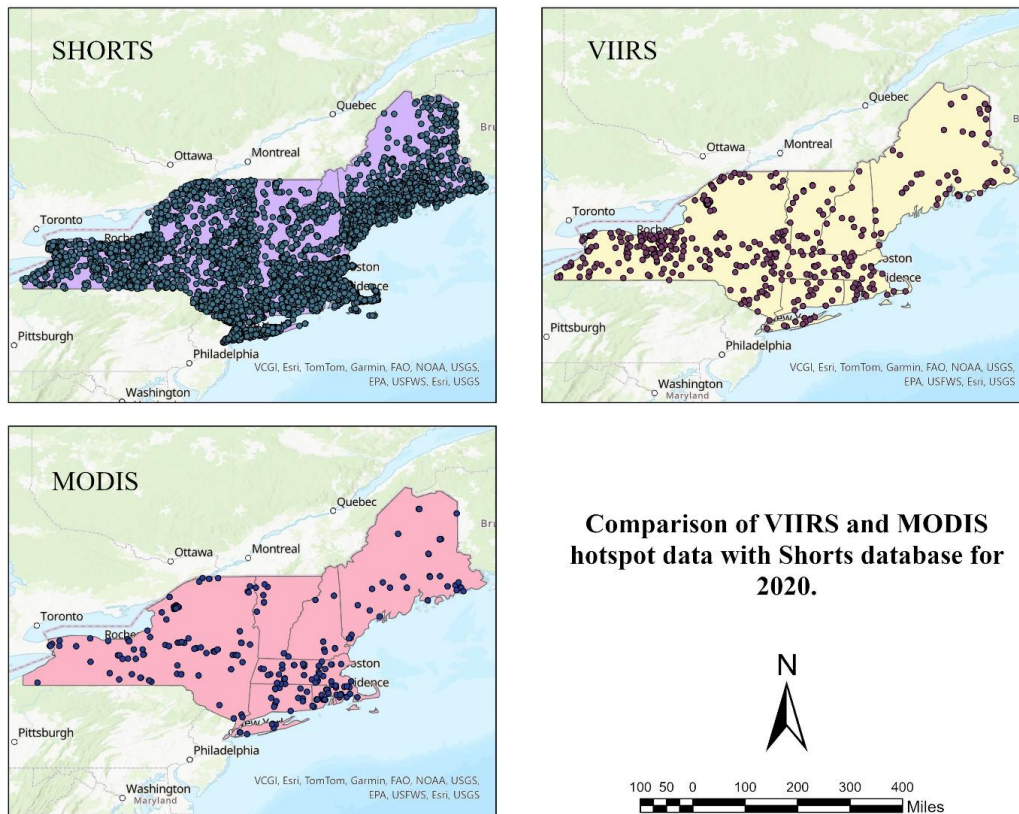
156 In the current article a wildfire dataset developed based on the VIIRS 375 m active fire product is presented (Qadiri
157 and Cerrai, 2025). The dataset was developed to fill the gaps present in the currently available CONUS datasets.
158 Currently available manual datasets do not provide fire extents over large domains, while the satellite derived datasets
159 either have low spatial resolution due to the raw product used or do not provide fire extent. This product uses a very
160 high resolution VIIRS S-NPP product available on a global scale at high temporal resolution (2 - 4 passes per 12
161 hours). In this article, section 2 will explain the motivation and methodology behind the evolution of the dataset.
162 Section 3 will explain data characteristics and some case studies to highlight the dataset in detail with examples of
163 some well-known fires.

164

165 **2. Methodology**

166 **2.1 Motivation**

167 The product described in this manuscript was created due to lack of a dependable wildfire dataset product for the
168 Northeast of CONUS. We analyzed the ground-based dataset by Short et al. (2022), and the MODIS and VIIRS
169 satellite pixel products from NASA fire information for resource management system (NASA | LANCE | FIRMS)
170 (Figure 2). The Short database was created to supplement the National fire program analysis (FPA) system. There are
171 some spatial inconsistencies in the Short database even though significant improvements have been made to the
172 current version compared to previous versions. It further lacks fire extent or perimeter information. The MODIS and
173 VIIRS S-NPP fire product from NASA FIRMS (NASA | LANCE | FIRMS) were just the fire pixels or hotspots and
174 not actual fires. By comparing the MODIS and the VIIRS products, we found that the number of fire pixels in VIIRS
175 is much higher than the one detected by MODIS due to the higher resolution.



176

177 **Figure 2 Comparison of short et al (2016) database and MODIS firms hotspot product for the Northeast CONUS for 2001.**

178 To address the shortcomings related to lack of spatial extent, spatial inconsistency, different resolutions and
179 identification of hotspots instead of actual fires, we developed a novel CONUS Fire Extent (CONFEX) product for
180 CONUS and Alaska, based on the VIIRS hotspot or fire pixel data from NASA FIRMS.

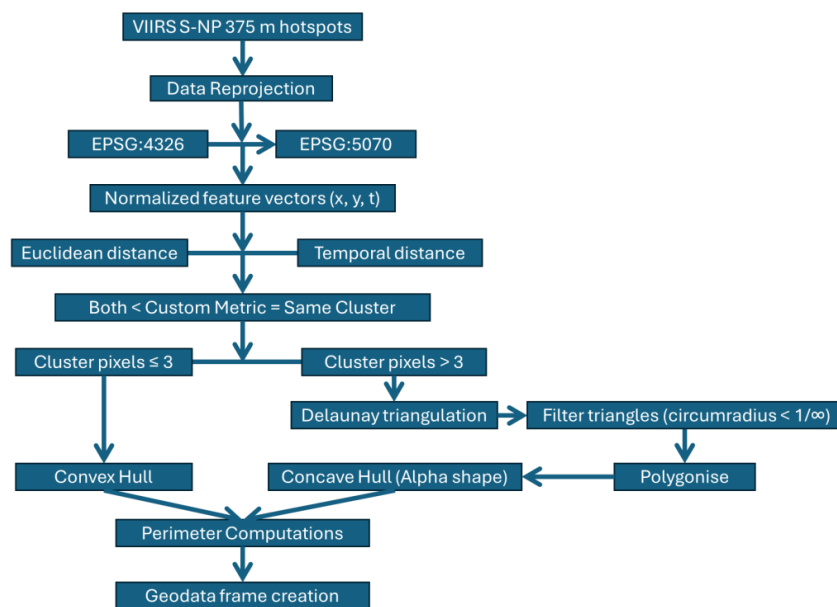
181 The final product was developed for the period 2012- 2023 based on the availability of the NASA FIIRMS data. All
182 the pixels were retained, and the entire CONUS and Alaska were processed. The CONFEX database contains the area,
183 the start date, the end date, the centroid location and perimeter of the fire. The work was aided by some earlier products
184 formed based on MODIS and VIIRS (Artés et al., 2019; Chen et al., 2022) but there is no database of this kind based
185 on VIIRS 375 S-NPP for the CONUS to the authors knowledge.

186 2.2 Algorithm for the generation of the CONFEX product

187 The algorithm uses the VIIRS S-NPP 375 m resolution pixel or hotspot data, available from the NASA FIRMS website,
188 as input (Figure 3). The hotspots are first re-projected into a EPSG: 5070 geographic coordinate system, suited for
189 CONUS, to keep as close to the actual distance as possible accommodating the earth's shape at low latitudes. Then,
190 each hotspot is converted into a normalized feature vector. The algorithm then uses DBSCAN to calculate the
191 Euclidean distance and absolute temporal difference in pairs for hotspots and each pair with both values below custom
192 metrics are in the same cluster. The clustering distance threshold for each pair is 750-meter. This comes from the fact
193 that it is twice as much as 375 m, therefore for two tangent circles of radius 375 m, the distance between two points
194 is 750 m. The clustering temporal distance threshold is 120-hour, which is based on the fact that we assumed absence
195 of a fire, if no fire is detected over multiple overpasses in a period of 5 days. Once the clustering is done, if the cluster



contains more than 3 hotspots, Delaunay triangulation is performed, keeping the triangles in which the circumradius is not more than $1/\infty$. Finally, polygonization of the filtered triangles is performed to assign an alpha shape to the cluster (concave hull) which has been previously used by other researchers for assigning shapes to fire clusters (Chen et al., 2022). For clusters containing less than 3 hotspots, convex hull is used to assign shapes. After assigning the shapes, the areas of the resultant concave or convex polygons are calculated. The resultant polygon is saved as a shape file for each fire cluster. The resultant data set is composed of geodata frames assigned to each cluster, with attributes of ignition time and date, end time and date, area, ignition latitude, ignition longitude, centroid latitude, and centroid longitude. Moreover, a csv file in which each row corresponds to the summary of all the important attributes of one cluster's geodata frame is also created. This summary file also contains the cluster ID which can be used to locate the shapefile with the same ID name.



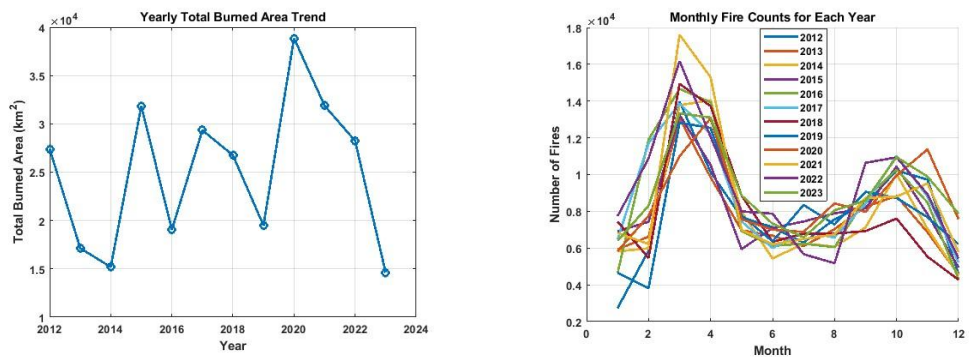
206

207 **Figure 3 Algorithm for developing CONFEX from VIIRS S-NPP 375 m data.**

208

209 **3. Dataset characteristics**

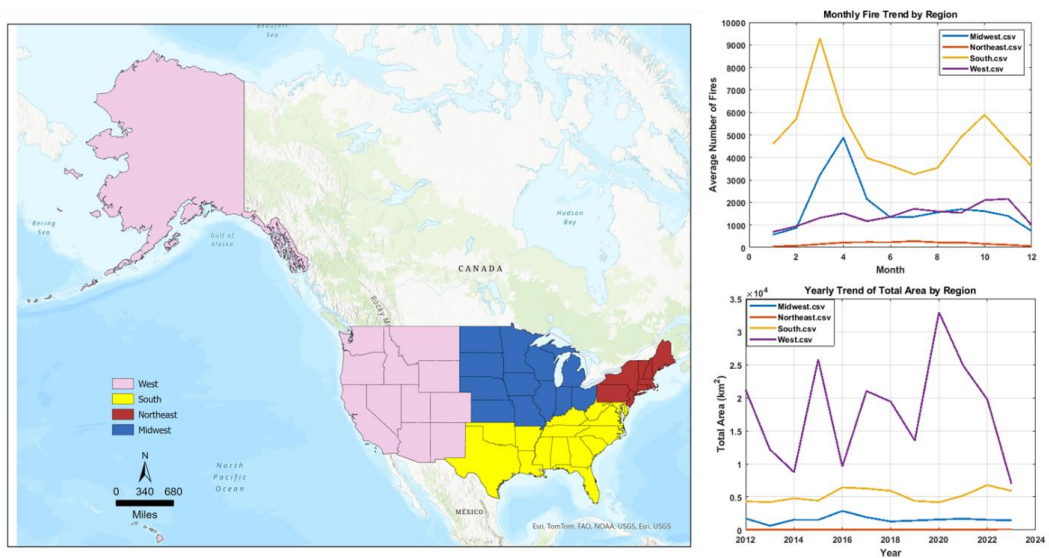
210 In this section, we analyzed the CONFEX database to provide an overview of its spatial and temporal characteristics
211 and highlight its strengths and limitations. From a time series of the newly developed CONFEX summary database
212 computed over the CONUS area and Alaska, it is possible to find that the year characterized by the largest burnt area
213 was 2020, with 38844.96 Km², followed by 2021 (Figure 4 a). The fire season is bimodal with one peak in March and
214 the other in October (Figure 4 b). It is also possible to notice a very high interannual variability. As an example, the
215 burnt area in 2023 has been less than half the activity in 2020.



216

217 **Figure 4 a) Time series for total area burned each year based on the CONFEX; b) Time series for number of fires each**
218 **month across years in CONFEX**

219 The regional trend of the number of fires in CONFEX reveals that the fire season is bimodal for the south with one
220 peak in spring and the other in late fall (Figure 5 b). The west has a unimodal fire activity with peak lasting from late
221 fall to early winter. Midwest has unimodal activity with spring peak, while the activity in northeast is uniform
222 throughout the year with relatively more activity in summer but no clear peaks. The yearly area trend reveals that the
223 west has the most burnt area followed by the south and midwest, respectively (Figure 5 c). The northeast has the
224 lowest fire activity of the four regions.



225

226 **Figure 5 a) Different regions of CONUS and Alaska. b) Regional monthly trend of number of fires for CONUS and Alaska**
227 **regions, based on CONFEX. c) Timeseries for total area burnt for different regions of CONUS and Alaska regions based**
228 **on CONFEX.**

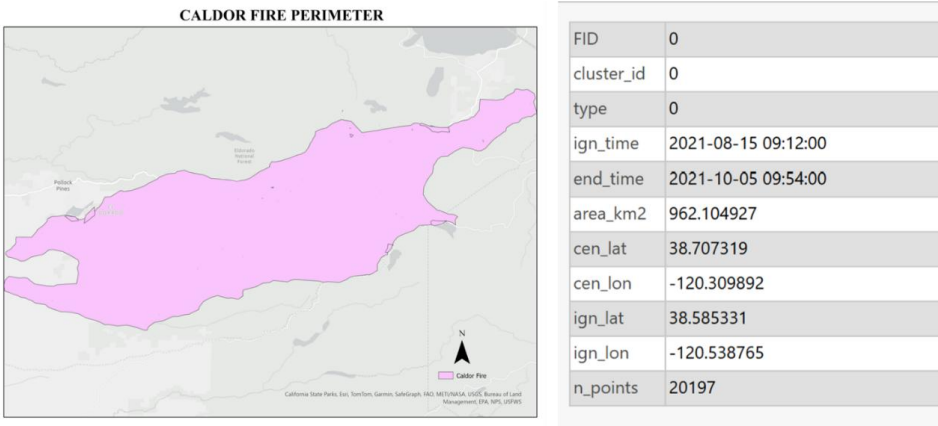
229 In Figure 6, it is possible to see an example of fire perimeter and attributes for the Caldor fire. As per CALFIRE, the
230 area burnt in Caldor was 897.73 Km² which is 6.7% lower than the 962.1 Km² observed through CONFEX ([Caldor](#)
231 [Fire | CAL FIRE](#)). The CALFIRE provides a temporal span of the Caldor fire from 8/14/2021 to 10/21/2021 while
232 CONFEX from 8/15/2021 to 10/05/2021. The alpha shape is supplemented by the various attributes which are



233 available for every shape file or cluster. The number of points in each cluster is also presented, to provide an estimate
234 of the number of hotspots detected by VIIRS in that cluster.

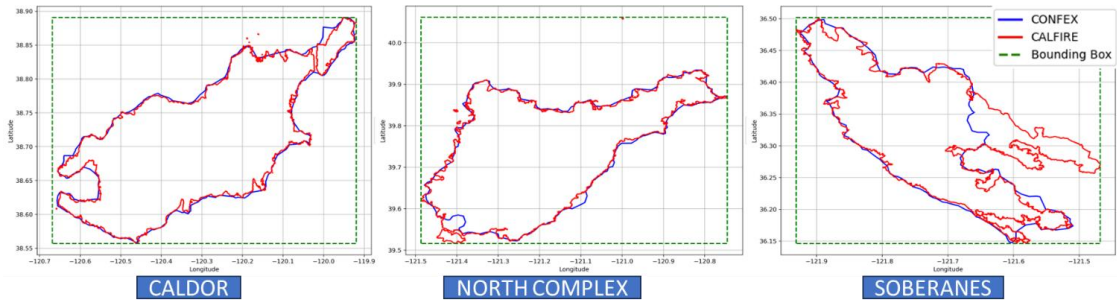
235 The CONFEX contains individual shape files of all the fires from 2012 to 2023, as captured by VIIRS S-NPP 375 m.
236 The final product contains a csv file summary of all the fires, which can be readily used in ignition modelling.

237



238
239 **Figure 6 Database example for Caldor fire California (Aug – Oct 2021).**

240 We compared the fire perimeter of three large fires, Caldor, North Complex, and Soberanes, from CONFEX with the
241 Department of Forestry and Fire Protection's Fire and Resource Assessment Program (FRAP) California database
242 (data.ca.gov/dataset/California-fire-perimeters-1950) (Figure 7). We computed a confusion matrix by creating a
243 bounding box around the two perimeters and we showed the results in Figure 8 and Table 1. The bounding box
244 was divided into 0.01° by 0.01° cells and based on that a confusion matrix was developed. These are just three examples,
245 and the database is enormous (1,202,150) and individual results may vary; however, these examples gave very
246 satisfactory results.



247
248 **Figure 7 Comparison of perimeters from CONFEX and California Fire Department (CALFIRE) for Caldor, North**
249 **complex, and Soberanes fires with respect to bounding boxes for confusion matrix analysis.**

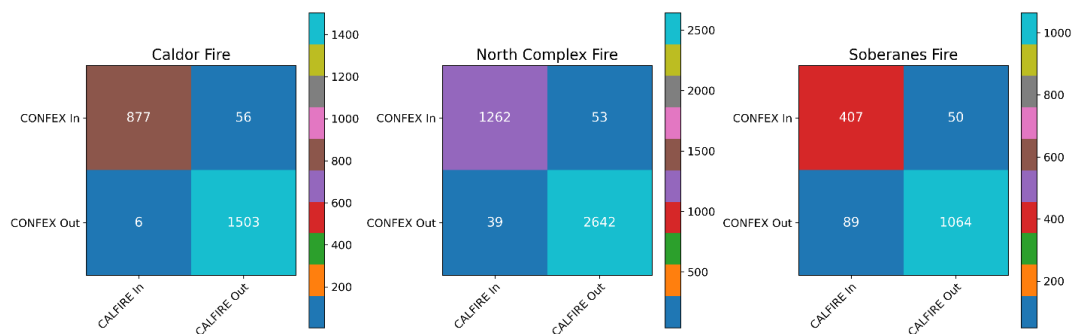


Figure 8 Confusion Matrices for the three fires depicted above.

Table 1 True Positive Rate, False Positive Rate, True Negative Rate, False Negative Rate, F1 score, Critical Success Index for the confusion matrices above, for the three fires.

Fire Name	TPR	FNR	TNR	FPR	F1 Score	CSI
Caldor	99.3%	0.7%	96.4%	3.6%	96.6%	93.4%
North Complex Fire	97.0%	3.0%	98.0%	2.0%	96.5%	93.2%
Soberanes	82.0%	18.0%	95.5%	4.5%	85.4%	74.5%

The TPR values range from 82 – 99% and indicate that the clustering process was able to correctly capture the majority of the three fires. The TNR values from 95 – 98% show that whenever there was no fire, there was no false detection, or the algorithm did not incorporate pixels of a different fire. The FNR value varies from 0.7 – 18%, the high value was in case of Soberanes fire, in which the algorithm was not able to detect a considerable portion of the fire, which is detected as a separate fire cluster, since the clustering process was based on the entire database. However, looking at the F1 score, CSI shows a good overall performance of the algorithm in classifying the pixels in right clusters.

4. Data Availability

The data set presented in this work is available on Mendeley data repository. The DOI for the dataset is 10.17632/sk6jwy7xmg.2 (Qadiri and Cerrai, 2025).

5. Conclusion

In this article we defined a new wildfire database for the CONUS and Alaska regions. The CONFEX database is based on the VIIRS S-NPP 375m satellite data product. Currently available datasets are either spatially inconsistent or lack perimeter information or are based on low resolution satellite products. The CONFEX database was created to address these gaps by using a high resolution VIIRS S-NPP 375m data product. The final product contains the perimeter, start date, end date, ignition and centroid location of the wildfires. The algorithm we developed takes the hotspot data and clusters it into actual fires over the specified region by using DBSCAN clustering, based on custom metrics. When compared with already available fire extents for three well known fires, we attained an F1 score of 85.5 – 96.6 % and a CSI of 73.5 – 93.4 %, showcasing the efficacy of the algorithm in classifying the hotspots. The CONFEX database has the potential to become an important asset for researchers looking for a comprehensive, high-resolution product for the CONUS and Alaska regions.

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