

1 Large synthesis of in situ field measurements of the size distribution of mineral dust aerosols 2 across their lifecycle

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8 9 Abstract

10 Mineral dust aerosol is important in the Earth system and the correct representation of its size
11 distribution is fundamental for shaping the current state and the evolution of climate. Despite many
12 observational dust size data that are available in the literature, using this body of information to properly
13 guide the development and validation of climate models and remote sensing retrievals remains
14 challenging. In this study we collect, evaluate, harmonize, and synthesize 58 size distribution data from
15 the past 50 years of in situ field observations with the aim of providing a consistent dataset to the
16 community to use for constraining the representation of dust size across its lifecycle. Four levels (LEV)
17 of data treatment are defined, going from original data (LEV0), data interpolated and normalized on a
18 standardized diameter grid (LEV1), and data in which original particle diameters are converted into a
19 common geometrical definition under both spherical (LEV2a) and aspherical (LEV2b) assumptions. Size
20 distributions are classified as emission/source (SOURCE, <1 day from emission; number of datasets in
21 this category, N=12), mid-range transport (MRT, 1–4 days of transport; N=36) and long-range transport
22 (LRT, >4 days of transport; N=10). The harmonized dataset shows consistent features suggesting the
23 conservation of airborne particles with time and a decrease of the main coarse mode diameter from a
24 value of the order of 10 μm (in volume) for SOURCE dust to a value of the order of 1-2 μm for LRT
25 conditions. An additional mode becomes evident below 0.4 μm for MRT and LRT dust. Data for the three
26 levels (LEV1, LEV2a, LEV2b) and the three categories (SOURCE, MRT, LRT), together with statistical
27 metrics (mean, median, 25% and 75% percentiles, and standard deviation) are made available as:

28 SOURCE (<https://doi.org/10.57932/58dbe908-9394-4504-9099-74a3e77140e9>; Formenti and Di Biagio, 2023a);

29 MRT (<https://doi.org/10.57932/31f2adf7-74fb-48e8-a3ef-059f663c47f1>; Formenti and Di Biagio, 2023b);

30 LRT (<https://doi.org/10.57932/17dc781c-3e9d-4908-85b5-5c99e68e8f79>; Formenti and Di Biagio, 2023c).

31 32 Introduction

33 Airborne mineral dust aerosols emitted by the aeolian erosion of bare soils contribute in a major way to
34 the Earth's radiative budget and environmental processes, including the human health. Because of their
35 native mineralogical composition and size distribution, they interact with solar and infrared radiation,
36 influence the formation and brightness of liquid and ice clouds, and affect the composition of the
37 atmosphere and the ocean, while also transporting pollutants, viruses and bacteria across the
38 continents and the oceans (Knippertz and Stuut, 2014, and the many references therein).

39 As a consequence, a large effort has started in the last decades to include the representation of those
40 properties in climate and air quality models. Indeed, the complex mineralogy of mineral dust, depending
41 on that of the parent soils (Claquin et al., 1999; Journet et al., 2014; Gonçalves Ageitos et al., 2023a), is
42 now accounted for in models (Scanza et al., 2015; Perlwitz et al., 2015a; 2015b; Menut et al., 2020; Kok
43 et al., 2017; Di Biagio et al., 2020; Gómez Maqueo Anaya et al., 2024) and starts to be retrieved by
44 remote sensing (Green et al., 2020; Zhou et al., 2020; Di Biagio et al., 2023).

45 On the other hand, representing the span and the variability in time and space of the dust aerosol size
46 distribution remains a challenge.

47 The particle size distribution of mineral dust extends over several orders of magnitudes. Iron-rich
48 particles as small as 14 nm in diameter have been observed in the laboratory from deflating soils by
49 Baddock et al. (2013). During sandstorm in Algeria, Gomes et al. (1990) measured an increase of the
50 mass concentration of particles between 100 nm and 1 μm , and attributed this to clays disaggregated
51 by sandblasting. Measurements of the size-resolved vertical dust flux by Gillette et al. (1972; 1974a;
52 1974b) based on microscopy analyses of samples from Texas and Nebraska showed the presence of
53 particles up to several microns in dust emissions.

54 The representation of the accumulation and coarse modes in mineral dust has long been based on the
55 columnar measurements by the sun/sky photometers of the Aerosol Robotic Network (AERONET)
56 network, which provides with normalized size distributions of mineral dust considered as chemically
57 homogeneous particles the 0.1—30 μm optically-equivalent diameter (Dubovik et al., 2002; 2006;
58 Holben et al., 2011), and which, incidentally, serve also the look-up tables of the remote sensing
59 retrievals of dust from space (e.g., Cuesta et al., 2015; Zhou et al., 2020).

60 Nevertheless, in situ observations at ground-based stations and on aircraft in more recent years have
61 shown that particles of several tens, sometimes hundreds, of micron are airborne at emission, and
62 remain so after several days of transport (Reid et al., 2003; Formenti et al., 2003; Rajot et al., 2008;
63 Chou et al., 2008; Kandler et al., 2007; 2009; Wagner et al., 2009; Klaver et al., 2011; Ryder et al., 2013;
64 2015; Rosenberg et al., 2014; Denjean et al., 2016; Wienzerl et al., 2017; van der Does et al., 2018).

65 These observations have been instrumental to a number of advances. Using them as ensemble dataset,
66 to smooth local atmospheric variability, they have served as a basis to a new classification of the dust
67 size distribution in four modes, namely fine dust (diameter $\leq 2.5 \mu\text{m}$), coarse dust ($2.5 < \text{diameter} \leq 10$
68 μm), super coarse dust ($10 < \text{diameter} \leq 62.5 \mu\text{m}$) and giant dust (diameter $> 62.5 \mu\text{m}$), extending above
69 the size range retrieved by AERONET (Adeyemi et al., 2023). Additionally, they have also fostered the
70 revision of the numerical schemes of emissions and deposition, and identified the numerous processes
71 and properties (non-spherical shape of particles, electric forces, atmospheric turbulence), that could
72 counteract the size-selective removal by gravitational settling and keep particles airborne longer than
73 expected (Kok, 2011; Huneus et al., 2011; Mahowald et al., 2011; Kok et al., 2017; Di Biagio et al, 2020;
74 Zhao et al., 2022; Adebiyi and Kok, 2020; Adebiyi et al., 2020; Huang et al., 2021; Meng et al., 2022;
75 Adeyemi et al., 2023).

76 In support of those activities, in this paper we present a large and standardized compilation of *in situ*
77 observations of the particle size distribution of mineral dust conducted during the past 50 years of
78 research. This dataset extends the currently published compilations of measurements (Meng et al.,
79 2022; Adeyemi et al., 2020; 2023) to provide with a state-of-the art of the current knowledge in support
80 to the development of models, and ground-based and satellite remote sensing. Analysis of this dataset
81 may provide with an integrated view of the size distribution of dust particles across their life cycle to
82 evaluate their impacts in the Earth/human system.

83 **2. Methods**

84 **2.1 Constitution of the dataset**

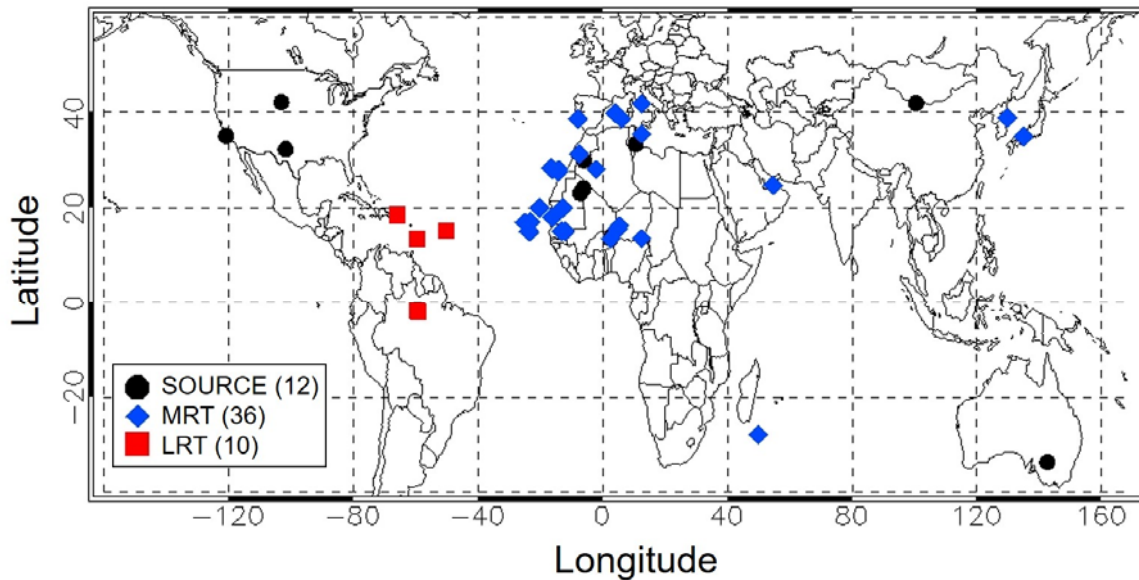
85 Data presented in this paper result from in situ ground-based and aircraft observations of airborne dust
86 conducted during field campaigns during the past 50 years of dust research. Data from deposition
87 samples (e.g., van der Does et al. 2018 or Varga 2021) are not considered in this analysis.

88 Only datasets being published and properly referenced in the open peer-reviewed literature were
89 retained. We also privileged datasets for which the methodology of acquisition, calibration and data

90 treatment was well described so that the data quality can be assessed. Finally, we search for data as
91 much as possible representative of different source and transport regions of the world.

92 The observations contributing to the dataset are listed in **Table S1** and the spelling of the acronyms of
93 the field campaigns is reported in **Text S1** in the supporting material. Data are geo-localized in **Figure 1**,
94 where they are classified with respect to their time after emission. Geographical coordinates are
95 reported in Table S2.

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101 **Figure 1.** Geographical location of the datasets contributing to size distribution observations for the source, the
102 mid-range transport (MRT) and the long-range transport (LRT) categories. The legend indicates the line style
103 used in the plot. The number of data for each category is indicated in the parenthesis in the legend.

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105 Observations obtained at the time of dust emission or within 1 day after emission are classified as
106 SOURCE. Observations corresponding to 1 to 4 days after emission and/or geographically acquired
107 near-source regions (for example, offshore North Africa) are classified as mid-range transport (MRT).
108 Observations at times exceeding 4 days after emission or geographically distant from source regions
109 (for example, observations in the Caribbean) are classified as long-range transport (LRT). To note that
110 potential uncertainties may arise in this classification, in particular for datasets lying at the boundaries
111 of the SOURCE, MRT and LRT categories, and we acknowledge this aspect as a source of error in our
112 analysis. We invite the reader to refer to the Supplementary material (Text S4) for thorough description
113 of the assumptions made in some cases to associate each dataset to a category.

114 The SOURCE dataset (Fig 1, black points) consists in 12 observations in Northern Africa, North America,
115 and Asia, and one data set in Australia. They include works by Gillette et al. (1972, 1974), Gillette (1974),
116 Fratini et al. (2007), Rajot et al. (2008), Sow et al. (2009), Shao et al. (2011), Ryder et al. (2013a, 2013b),
117 Rosenberg et al. (2014), Huang et al. (2019), and Khalfallah et al. (2020), a set of data recently used by
118 Kok et al. (2017), Di Biagio et al. (2020) and Huang et al. (2021) to constrain the shape of dust size
119 distribution at emission in model studies, and the most recent work by Gonzales-Florez et al. (2023).
120 The MRT class (Fig. 1, blue points) is contributed by 36 datasets from field campaigns (ACE2, ACE-Asia,
121 ADRIMED, AER-D, AMMA, DABEX, DARPO, DIAPASON, DODO1-2, FENNEC, GAMARF, GERBILS, INDOEX,
122 NAMMA, RHaMBLe, SALTRACE, SAMUM1-2, TRACE-P, and UAE2) in Western Africa, Capo Verde, the
123 Mediterranean basin, the eastern tropical Atlantic, Saudi Arabia, Japan, and Indian Ocean, downwind
124 sources either over the ocean or over desert areas. Additional datasets from studies performed in the

125 Sahara, the Atlantic Ocean, Canary Islands and Japan (Schütz, 1981; D’Almeida et al., 1987; Maring et
126 al., 2000; Kobayashi et al., 2007) are added to the dataset. The LRT class (Fig. 1, red points) lays on 10
127 datasets of observations across the Atlantic Ocean and South America and is contributed by
128 observations from Bacex, CLAIRE, Dust-Attack, Go-Amazon, PRIDE, and SALTRACE campaigns and
129 intercontinental dust transport data from Schütz (1981).

130 2.2. Instrumentation contributing to the in situ dataset

131 The natural dynamical range of the particle size and concentration of mineral dust can only be
132 represented by a combination of instruments based on different intrinsic particle properties such as
133 density, electrical charge, shape and composition (e.g., Reid et al., 2003a; Formenti et al., 2011;
134 Wendisch and Brenguier, 2013; Mahowald et al., 2014, Adeyemi et al., 2023). As a consequence, the
135 datasets considered in this paper are contributed by different in situ instruments, also described in **Text**
136 **S2** in the supporting material, namely:

- 137 ○ Optical particle counters (OPC) using the dependence of light scattering on particle size and providing
138 with the particle concentration as a function of the optical equivalent diameter (e.g., Reid et al.,
139 2003b; Clarke et al., 2004; Osborne et al., 2008; Formenti et al., 2011; Ryder et al., 2013a, 2018;
140 Khalfallah et al., 2020).
- 141 ○ Particle collection by filtration or impaction followed by individual particle characterization by
142 transmission (TEM) and/or scanning electron microscopy (SEM) sizing particles as function of their
143 equivalent projected-area diameter and coulter geometric sizing methods, (e.g., Gillette et al., 1972,
144 1974a, 1974b; Reid et al., 2003a; Khobayashi et al., 2007; Kandler et al., 2009; Chou et al., 2008).
- 145 ○ Multi-stage filtration or impaction sampling coupled with gravimetric or chemical analysis providing
146 with the mass size distribution as equivalent aerodynamic diameter (e.g., Formenti et al., 2001; Reid
147 et al., 2003b).
- 148 ○ Differential and Scanning Mobility Particle Sizer (DMPS and SMPS) providing the size of particles in
149 the submicron range as the electrical mobility equivalent diameter of a charged particle moving in a
150 static electric field (e.g., Maring et al., 2000, 2003; Bates et al., 2002; Müller et al., 2010; Denjean et
151 al., 2016a, 2016b).
- 152 ○ Aerodynamic particle sizers (APS), measuring the equivalent aerodynamic diameter of a sphere of
153 unit density having the same terminal velocity in an accelerated airflow as the irregularly shaped
154 dust particles (e.g., Maring et al., 2003; Reid et al., 2003b; 2008; Struckmeier et al., 2016)

155 Each of those instrument types size particles on an equivalent diameter (optical, projected-area,
156 aerodynamic, mobility) that depends on their respective working principle. Converting those
157 operational size definitions into a homogenized one is part of the treatment applied in this work, which
158 follows the theory proposed and discussed in the literature and benefits of recent progresses in
159 characterizing/synthetizing dust properties relevant for these treatments (e.g., Hinds, 1999, De Carlo et
160 al., 2004 ; Mahowald et al., 2014; Di Biagio et al., 2019; Huang et al., 2020, 2021). Diameter definitions
161 and formulas to convert each of them into a geometrical diameter, both under the assumption of
162 spherical and aspherical dust, is provided in **Text S3** and summarized in **Table S3**.

163 **Text S4** presents relevant information on each dataset considered in the present analysis. This includes
164 a brief description of the field operations, the experimental conditions, the type of original data
165 (number, volume or mass concentration size distribution, size-resolved emission fluxes), the
166 instrumentation, and the data treatment applied to the measurements (averages, diameter corrections,
167 etc.) in the original publication. Original data were obtained, as much as possible, through a personal
168 contact with the data providers or from the original publications. This is also indicated in **Text S4**.

169 2.3. Data treatment, harmonization, and synthesis

170 The original observations were treated to provide with a harmonized dataset in terms of the definition
171 of particle diameter and data were normalized to remove differences in sampled number
172 concentrations. Four level of data treatment are defined as described below.

173 1/ *Level-0 (LEVO)*: original data, taken at the native resolution or the resolution from digitalization
174 process and converted into volume distribution assuming spherical particles ($\pi/6 \cdot D^3 \cdot dN/d\log D$), where
175 D is the particle diameter used in the publication and $dN/d\log D$ is the particle number concentration.
176 To remove differences in concentration, and in absence of information on original bin width, *LEVO* data
177 are normalized to the maximum of the volume size distribution;

178 2/ *Level-1 (LEV1)*: data from *LEVO* are interpolated over a common size range of equi-logarithmically
179 spaced diameters ($d\log D = 0.05$) encompassing the original diameter range for each dataset and
180 normalized so that the integral is equal to 1 over a common diameter range. The diameter range for
181 integral normalization was set to be the largest as possible and to be covered by more than 90% of the
182 datasets in each category. For SOURCE data it resulted that the diameter range for common integral
183 normalization is within 1.58 and 7.1 μm , and for MRT and LRT it is between 0.71 and 8.9 μm .

184 3/ *Level-2a (LEV2a)*: based on *LEV1*, the *LEV2a* data treatment aims at harmonizing the size distributions
185 by converting the operational original particle diameters, which depend on the physical principle of each
186 instrument, into a common-defined sphere-equivalent geometric diameter. Data from *LEV1* are
187 treated as in the following with respect to their diameter corrections:

- 188 ○ data already provided as geometrical diameters (from coulter counters, i.e., only one dataset in
189 our study) are left unchanged;
- 190 ○ data provided as projected-area diameters (i.e. from microscopy) are left unchanged;
- 191 ○ data provided as aerodynamic diameters (from APS or cascade impactors) are corrected
192 assuming a shape factor (χ) of 1 (under spherical assumption), therefore a size-invariant
193 conversion factor of 1.58 (see Eq. S2) is applied to the dataset assuming dust density of 2.5 g
194 cm^{-3} ($D_{\text{geom}} = D_{\text{aerod}}/1.58$). If original aerodynamic diameter data are already converted into
195 geometrical diameter, we replace the original correction with the conversion factor of 1.58.
196 Since the correction is a multiplicative factor the $d\log D$ of the bins remain unchanged;
- 197 ○ data provided as optical diameters (from OPCs) are converted into sphere-equivalent
198 geometric diameters applying the optical to geometrical correction by assuming homogeneous
199 spherical particles and a value of CRI of 1.53–0.003i. This CRI value is at the average of the dust
200 refractive indices reported in the 370–950 nm spectral range in Di Biagio et al. 2019) for dust of
201 global origin. Data for applying the correction for the different model of OPCs considered were
202 taken from Formenti et al. (unpublished data) and conversion factors were recalculated at the
203 $d\log D$ path of 0.05 assumed in the interpolated sizes. For the GRIMM 1.108 we used the data
204 taken from Formenti et al. (2011) interpolated at the 0.05 $d\log D$ path of our diameters. In order
205 to avoid discontinuities appearing and because of the new $d\log D$ do not significantly differ on
206 average from the value of 0.05 for D_{geom} calculated from D_{opt} interpolated data, we do not
207 update the $d\log D$, so that the conversion only imply a shift of the diameter. More details on the
208 choices applied for corrections in different cases are provided in Text S4. Original datasets
209 already converted into geometrical diameter, are left unchanged. However, it is worth noting
210 that the ensemble of data already applying an optical to geometrical correction uses a CRI
211 varying between 1.53 and 1.55 for the real part and 0.001 and 0.004 for the imaginary part and
212 work under the hypothesis of homogeneous spherical particles (Mie theory), therefore
213 consistent with our treatment. Exceptions are Khalfallah et al. (2020) using a CRI of 1.43–0.00i
214 as for quartz particles, and González-Flórez et al. (2023) using a CRI of 1.49–0.0015i and also
215 applying calculations in ellipsoidal assumption instead of Mie theory. The only dataset not
216 theoretically submitted to the optical to geometric correction is the one provided by Renard et
217 al. (2018) using an OPC built with a specific geometry making the measurements very low
218 sensitive to CRI calibration.

219 4/ *Level-2b (LEV2b)*: based on LEV1, the LEV2b data treatment aims at harmonizing the size distributions
220 by converting the operational original particle diameters into a common-defined geometrical diameter
221 by taking into account that mineral dust is aspherical. Data from LEV1 are treated as in the following
222 with respect to their diameter corrections:

- 223 ○ data already provided as geometrical diameters from coulter counters are left unchanged. This
224 technique is in fact only slightly affected by shape effects, as discussed by Kobayashi et al.
225 (2007);
- 226 ○ data provided as projected-area diameters are corrected using the size-invariant correction
227 factor of 1.56 from Huang et al. (2021) ($D_{\text{geom}}=D_{\text{area}}/1.56$) (see Eq. S1);
- 228 ○ data provided as aerodynamic diameter are corrected assuming a size-invariant conversion
229 factor of 1.45 following Huang et al. (2021) ($D_{\text{geom}}=D_{\text{aerod}}/1.45$) (see Eq. S2);
- 230 ○ data provided as optical diameters and already treated as for LEV2a data, are further corrected
231 by applying a size-dependent aspherical to spherical ratio ($\text{ASR}(D_{\text{geom}})$) correction function,
232 $\text{ASR}(D_{\text{geom}})=(D_{\text{geom}})_{\text{aspherical}}/(D_{\text{geom}})_{\text{spherical}}$, to take into account non-sphericity effects in optical to
233 geometrical conversion. The ASR function (Fig. S1) is obtained by combining the optical to
234 geometrical diameter conversion factors for different OPCs calculated by Formenti et al.
235 (unpublished data) and Huang et al. (2021) both in the assumption of spherical homogeneous
236 particles ($(D_{\text{geom}})_{\text{spherical}}$) and tri-ellipsoids dust ($(D_{\text{geom}})_{\text{aspherical}}$). More details are provided in Text S3.
237 Original datasets derived from OPC measurements already provided as geometrical diameter
238 but under assumption of sphericity are also corrected by applying the $\text{ASR}(D_{\text{geom}})$ converting
239 function. The only exception are González-Flórez et al. (2023), that already apply tri-axial
240 ellipsoids calculations in their optical to geometric conversion, and Renard et al. (2018), not
241 requiring optical to geometrical conversion.

242 As for LEV1, the LEV2a and LEV2b data, for which a known interpolation path is used, are normalized so
243 that the integral of the volume size distribution is 1 over a common diameter range (1.58 – 7.1 μm for
244 SOURCE, 0.71 – 8.9 μm for MRT, LRT).

245 For each category (SOURCE, MRT, LRT) and for each data level (LEV1, LEV2a, LEV2b), the mean, median,
246 and standard deviation of the particle volume concentration per size class are calculated where at least
247 2 datasets are available in the diameter range. Additionally, the 25% and 75% percentiles are also
248 calculated, despite keeping in mind their limited representativeness given the reduced number of
249 samples in the datasets, especially for SOURCE and LRT classes.

250 2.4. Limitations of the chosen approach

251 Some precisions should be given when considering the LEV2a and LEV2b treatment reported in this
252 work. First, the implicit assumption when applying LEV2a and LEV2b dataset corrections is that dust is
253 the dominant aerosol species and possible effects due to internal or external mixing of dust with other
254 aerosol types are not taken into considerations (i.e., in the complex refractive index or shape factor
255 assumptions). Second, for those datasets that are obtained from the combination of different
256 techniques, namely DMPS+APS (Bates et al., 2002; Maring et al., 2000, 2003; Müller et al., 2010),
257 OPC+APS (Chen et al., 2011), SMPS + OPC (de Reus et al., 2000; Otto et al., 2007; Denjean et al., 2016a,
258 2016b), DMPS + APS + microscopy (Kandler et al., 2011), or multiple OPC instruments (Reid et al., 2003b;
259 McConnell et al., 2008; Johnson and Osborne, 2011; Ryder et al., 2013a, 2013b, 2018; Rosenberg et al.,
260 2014; Weinzierl et al., 2009, 2011, 2017), the choice is that of applying artefact corrections for the
261 dominant instrument, often the one in the extended coarse mode range, and consider this correction
262 applicable to the whole diameter range. This is because when multiples instruments are used to build a
263 size distribution it is then not easy to reconstruct the steps of data analysis and merging from the original
264 work. It follows the subsequent considerations:

- 265 1/ the corrections applied for the aerodynamic and projected-area diameter apply a constant
266 size-invariant scaling factor to the ensemble of the size distribution data. In this approximation, if
267 the SMPS/DMPS is combined with aerodynamic or microscopy data, a correction factor between

268 1.45 and 1.58, depending on the level and the technique as detailed in the previous section, is
269 applied in place of the factor 1 (spherical assumption) or 1.19 (aspherical assumption) (see Eq. S3)
270 expected to convert the mobility diameter to geometrical diameter in LEV2a and LEV2b data. As a
271 consequence, the submicron size is 20 to 58% finer than expected only due to mobility to geometrical
272 conversion.

273 2/ A similar approach is used to correct datasets where OPC is the main used technique to size dust
274 particles together with the SMPS. For LEV2a data the Mie correction is applied to the full size
275 distribution, but being the size-dependent correction mostly inactive for submicron particles (i.e.
276 $D_{\text{geom}} \sim D_{\text{opt}}$ for most OPCs), the approach is mostly equivalent at considering a mobility diameter
277 correction with a shape factor of 1. For LEV2b data, using OPC corrections induce a limited right
278 shifting of the size distribution compared to the one that would be obtained from mobility
279 conversion because of the magnitude of the ASR function (Fig. S1) compared to the shape factor of
280 1.19 assumed for aspherical dust.

281 3/ When datasets relying on multiple OPCs measurements, the assumption is that the “dominant” OPC
282 that is the OPC covering the largest range and the coarsest sizes in particular, is considered. Given
283 that optical to geometrical corrections are not relevant for submicron particles and that the
284 magnitude of the correction typically increases for increasing sizes, this assumption is not expected
285 to determine significant biases in the data. To mention additionally a general ambiguity of the optical
286 to geometrical correction around the diameter of 1 μm where a plateau in the scattering calibration
287 function for several OPCs models can be found (i.e. Formenti et al., unpublished data).

288 More details on the specific assumptions and choices done for each dataset are provided in **Text S4**.

289 Further, for LEV2a and LEV2b data for which corrections are applied on the data, caution is taken at the
290 boundary of the size distribution and when the first and/or the last bin of the corrected size showed
291 significant divergence, these data are removed from the dataset.

292 An additional source of error is the individual measurement uncertainty, which varies with the specific
293 setup, instrument and spatial and temporal extent of the measurement.

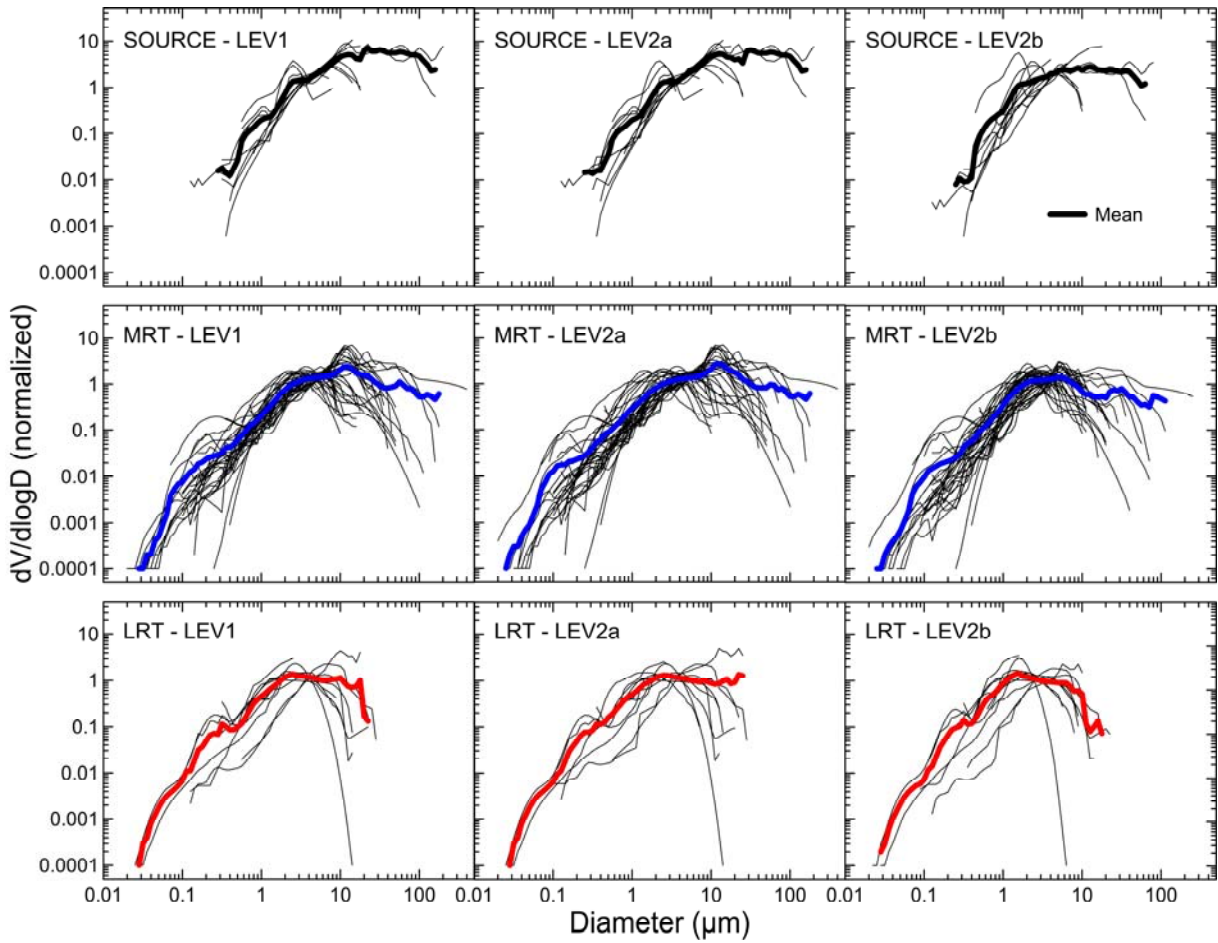
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295 **3. Presentation and discussion of the dataset**

296 Illustration of the data for different levels is provided in Figure 2. Figure 3 presents the synthesis of the
297 LEV2b data and the comparison of SOURCE, MRT and LRT distributions. The contribution of different
298 size classes to the total particle number, surface and volume is summarised in Table 1. Size classes have
299 been defined according to the classification of Adeyemi et al. (2023) defining fine dust ($D \leq 2.5 \mu\text{m}$),
300 coarse dust ($2.5 < D \leq 10 \mu\text{m}$), super coarse dust ($10 < D \leq 62.5 \mu\text{m}$) and giant dust ($D > 62.5 \mu\text{m}$). Within
301 the fine dust class, we further calculate the fractions of particles smaller than 0.4 μm .

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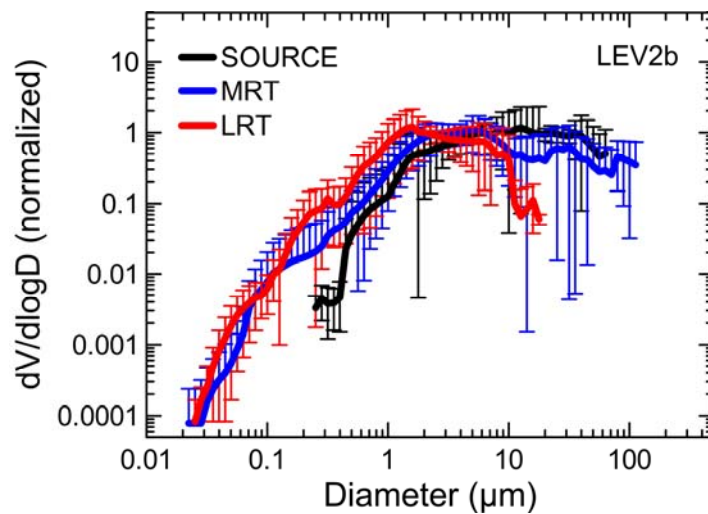
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305 **Figure 2.** Data for SOURCE, MRT, and LRT dust at level 1, 2a, and 2b as described in Sect. 2.3 (labelled as LEV1,
306 LEV2a, LEV2b, respectively). Single datasets, all normalized at the integral of 1, are plotted as black lines. The mean
307 (thick black, blue, and red line for SOURCE, MRT, and LRT, respectively) are shown at all levels. Note that the mean
308 is calculated only where at least 2 datasets are available in the diameter range.

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311 **Figure 3.** Comparison of normalized mean volume size distribution for the SOURCE, MRT, and LRT categories in our study
312 reported as LEV2b data (mean \pm standard deviation). For the sake of comparison, and differently from data in Fig. 2, the
313 SOURCE, MRT, and LRT synthesis datasets reported here are normalized at the integral equal to 1 over a common diameter
314 range corresponding to 0.35–17.8 μm . This is done to remove differences linked to different integration range for SOURCE data
315 compared to MRT and LRT.

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Dataset		$D \leq 2.5 \mu\text{m}$ ($D \leq 0.4 \mu\text{m}$)	$2.5 < D \leq 10 \mu\text{m}$	$10 < D \leq 62.5 \mu\text{m}$	$D > 62.5 \mu\text{m}$
Number	SOURCE	95.4% (20.4%)	4.5%	0.1%	0.4%
	MRT	99.8% (96.1%)	0.2%	0.0%	0.0%
	LRT	99.9% (94.5%)	0.1%	0.0%	0.0%
Surface	SOURCE	45.0% (1.1%)	39.4%	15.5%	0.14%
	MRT	65.4% (16.8%)	30.7%	3.6%	0.29%
	LRT	84.6% (23.1%)	15.1%	0.2%	0.00%
Volume	SOURCE	10.8% (0.1%)	34.9%	52.7%	1.6%
	MRT	22.1% (1.1%)	44.3%	25.7%	8.0%
	LRT	53.4% (3.6%)	44.5%	2.0%	0.0%

318 **Table 1.** Percentages of number, surface and volume size distribution in the diameter ranges $D \leq 0.4 \mu\text{m}$, $D \leq 2.5 \mu\text{m}$, $2.5 < D \leq$
319 $10 \mu\text{m}$, $10 < D \leq 62.5 \mu\text{m}$, and $D > 62.5 \mu\text{m}$ for the mean of the size obtained for the SOURCE, MRT, and LRT LEV2b datasets.

320

321 As shown in Fig. 2 and 3 the shape of the dust size distribution at emission and along transport shows
322 main consistent features. A main mode located at $\sim 10 \mu\text{m}$ (in volume) is observed for dust at emission
323 and close to sources, as based from the few studies allowing to measure up to the coarse fraction. The
324 main dust mode decreases to $\sim 5 \mu\text{m}$ and $\sim 2 \mu\text{m}$ for MRT and LRT conditions, respectively. Below $0.4 \mu\text{m}$
325 the dust volume size shows an additional mode, particularly visible for MRT and LRT. As a matter of fact,
326 the sparse datasets measuring very fine particles at the SOURCE show that particles with diameters
327 below $0.4 \mu\text{m}$ (however measured only down to $0.2 \mu\text{m}$, as shown in Fig. 2) represent approximately
328 20% of the total particles' number, increasing to more than 90% in MRT and LRT. Instruments such as
329 SMPS and DMPS used in MRT and LRT studies measure particles as small as $0.02 \mu\text{m}$ in diameter.
330 Previous single-particle compositional observations showing that the particle number concentration in
331 the size range between 0.1 and $0.4 \mu\text{m}$ is largely contributed by aluminosilicate dust particles at
332 emission, while internal or external mixing with aerosols other than dust gains importance with time
333 and altitude of transport (Chou et al., 2008; Kandler et al., 2007, 2009; Weinzierl et al., 2009; 2017;
334 Klaver et al., 2011; Denjean et al., 2016a; 2016b).

335 The normalized size distribution of dust particles between 0.4 and $10 \mu\text{m}$ is rather consistent and
336 invariant along the dust cycle. This is true in particular when restricting to the 2.5 to $10 \mu\text{m}$ size range
337 when differences are minimal and contribution to total volume is in between 34.9% and 44.5% . Below
338 that range, which is between 0.4 and $2.5 \mu\text{m}$, the contribution of particles for LRT is significantly higher
339 (53.4% in volume) than for SOURCE (10.8%) and MRT (22.1%), likely as, because of the normalization, it
340 compensates the decrease of particles larger than $10 \mu\text{m}$.

341 The magnitude of the particle volume above $10 \mu\text{m}$ remains unchanged almost up to $100 \mu\text{m}$ for both
342 the SOURCE and the MRT conditions, which also present similar particle volume. This mode decreases
343 very strongly for LRT conditions, when it represents only 2% of the total volume, compared to almost
344 55% and 34% for SOURCE and MRT, respectively.

345 The dataset presented in this work, synthetizing available *in situ* observations, allows evaluation of the
346 natural variability of dust size distribution along its lifecycle. To be emphasized, however, that while
347 consistent differences in the mean size distribution curves are obtained going from SOURCE to LRT, as
348 shown in Fig. 3, the inherent range of variability for each category, represented by the standard
349 deviation of the data, is also non-negligible and reflects the large range of documented size
350 distributions, together with the limited statistics available. This is particularly true for both super-coarse
351 and giant dust at MRT and LRT. Lower variability is identified below $0.4 \mu\text{m}$ because of the restricted
352 number of dataset available for MRT and LRT conditions, and there is an absence of data for SOURCE
353 dust below this size range.

354 Finally, to facilitate the use of these data within models and remote sensing schemes, Table 2 provides
 355 the parameters of lognormal size distributions fitting the LEV2a and LEV2b mean values of the three
 356 dust categories. Lognormal functions are set to reproduce the main shape of the dust distribution above
 357 0.4 μm , neglecting the specific features below this diameter where information is lower and the size
 358 affected by particle mixing with other compounds, especially for MRT and LRT. We found that a single
 359 broad mode can be employed to represent the main features of the volume size distributions above 0.4
 360 μm . Plots of the fitting functions are provided in supplementary Fig. S4. Because there is an inherent
 361 level of subjectivity in the choice of the number of modes and their parameters, we invite the individual
 362 researchers using the data to implement the parameterizations in accordance to their scientific needs.

363
 364

Dataset	Lognormal mode				
	$N_{\text{tot}} (\# \text{ cm}^{-3})$	NMD (μm)	$V_{\text{tot}} (\text{nm}^3 \text{ cm}^{-3})$	VMD (μm)	σ
SOURCE – LEV2a	$5.08 \cdot 10^{-10}$	0.355	7.76	26.69	3.32
SOURCE – LEV 2b	$9.8 \cdot 10^{-10}$	0.300	3.38	11.71	3.02
MRT – LEV 2a	$2.11 \cdot 10^{-9}$	0.150	2.55	11.64	3.33
MRT – LEV 2b	$6.82 \cdot 10^{-9}$	0.100	1.57	5.79	3.20
LRT – LEV 2a	$2.35 \cdot 10^{-9}$	0.280	1.39	3.88	2.55
LRT – LEV 2b	$2.96 \cdot 10^{-9}$	0.350	1.15	2.34	2.22

365 **Table 2.** Parameters (total number and volume concentration, $N_{\text{tot}} (\# \text{ cm}^{-3}), V_{\text{tot}} (\text{nm}^3 \text{ cm}^{-3})$, number and volume median
 366 diameter, NMD and VMD (μm), geometric standard deviation, σ) for the log-normal modes used to parameterize the LEV2b
 367 volume size distributions of the SOURCE, MRT, and LRT categories. Parameters refers to the following equations: $\frac{dV}{d\log D} =$

368 $\frac{\pi}{6} D^3 \frac{N_{\text{tot}}}{\sqrt{2\pi} \log \sigma} \exp\left(-\frac{(\log D - \log \text{NMD})^2}{2(\log \sigma)^2}\right)$ and $\frac{dV}{d\log D} = \frac{V_{\text{tot}}}{\sqrt{2\pi} \log \sigma} \exp\left(-\frac{(\log D - \log \text{VMD})^2}{2(\log \sigma)^2}\right)$

369

370 4. Conclusive remark

371 In this paper we present the most possible comprehensive synthesis of *in situ* observations of the
 372 particle size distribution of atmospheric dust aerosols. This compilation reflects the current state-of-
 373 the-art and represents a standardized and synthetic benchmark to constrain and evaluate models and
 374 satellite retrievals. We highlight differences and commonalities of the dust volume distribution as a
 375 function of time in the atmosphere, both in terms of main identified modes and relative contribution of
 376 dust in different size ranges.

377 We did this based on a large statistics of data and permit to retrieve robust information between 0.4
 378 and 10 μm where most of observations exist, while above and below this size range, observations are
 379 rare. Dust particles below 0.4 μm in diameter are seldom measured close to source regions, but are
 380 found in observations at mid- and long-range transport conditions. Their presence at emission, their
 381 size-segregated composition and state of mixing should be better documented and understood. The
 382 dynamics of the coarse mode above 10 μm , its invariance from source to mid-range transport, and
 383 decline afterwards is reported, and can challenge models.

384 We acknowledge the evidence that the compilation of a reference dataset is, almost by definition, a
 385 subjective and incomplete exercise which must revised continuously with the emergence of new
 386 datasets, new field campaigns, and the improvement of sampling techniques. We henceforth encourage
 387 colleagues to provide us with new or revised datasets to feed and update the dataset in the future.

388 Data availability

389 The LEV1, LEV2a and LEV2b datasets discussed in this paper are available on the EaSy Data, the Earth
 390 System Data repository (<https://www.easydata.earth/#/public/home>, last access: 01 June 2024)

391 maintained by the National French DATA TERRA research Infrastructure. Their respective DOIs are
392 summarized here below:

393 SOURCE_LEV1.dat, SOURCE_LEV2a.dat, SOURCE_LEV2b.dat : <https://doi.org/10.57932/58dbe908-9394-4504-9099-74a3e77140e9> (Formenti and Di Biagio, 2023a);

395 MRT_LEV1.dat, MRT_LEV2a.dat, MRT_LEV2b.dat: <https://doi.org/10.57932/31f2adf7-74fb-48e8-a3ef-059f663c47f1> (Formenti and Di Biagio, 2023b);

397 LRT_LEV1.dat, LRT_LEV2a.dat, LRT_LEV2b.dat : <https://doi.org/10.57932/17dc781c-3e9d-4908-85b5-5c99e68e8f79> (Formenti and Di Biagio, 2023c).

399 Figures of the individual datasets (including LEV0) are provided upon request.

400 **Code availability.** Data from images on published papers were digitalized with the online
401 WebplotDigitizer software available at <https://automeris.io/WebPlotDigitizer/>

402 **Author contributions.** PF and CDB designed the research, compiled and analysed the dataset, and wrote
403 the manuscript.

404 **Competing interests.** The authors declare that they have no competing interests.

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