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Radiative sensitivity quantified by a new set of radiation flux kernels based on the ERA5 reanalysis

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19 **Abstract**

20

21 Radiative sensitivity, i.e., the response of the radiative flux to climate perturbations, is essential
22 to understanding climate change and variability. The sensitivity kernels computed by radiative
23 transfer models have been broadly used for assessing the climate forcing and feedbacks for
24 global warming. As these assessments are largely focused on the top of atmosphere (TOA)
25 radiation budget, less attention has been paid to the surface radiation budget or the associated
26 surface radiative sensitivity kernels. Based on the fifth generation European Center for Medium-
27 Range Weather Forecasts (ERA5) atmospheric reanalysis, we produce a new set of radiative
28 kernels for both the TOA and surface radiative fluxes, which is made available at
29 <http://dx.doi.org/10.17632/vmg3s67568.3> (Huang and Huang, 2023). By comparing with other
30 published radiative kernels, we find that the TOA kernels are generally in agreement in terms of
31 global mean radiative sensitivity and analyzed overall feedback strength. The unexplained
32 residual in the radiation closure tests is found to be generally within 10% of the total feedback,
33 no matter which kernel dataset is used. The uncertainty in the TOA feedbacks caused by inter-
34 kernel differences, as measured by the standard deviation of the global mean feedback parameter
35 value is much smaller than the inter-climate model spread of the feedback values. However, we
36 find relatively larger discrepancies in the surface kernels. The newly generated ERA5 kernel
37 outperforms many other datasets in closing the surface energy budget, achieving a radiation
38 closure comparable to the TOA feedback decomposition, which affirms the validity of kernel
39 method for the surface radiation budget analysis. In addition, by investigating the ERA5 kernel
40 values computed from the atmospheric states of different years, we notice some apparent
41 interannual differences, which demonstrates the dependence of radiative sensitivities on the
42 mean climate state and partly explains the inter-dataset kernel value differences. In this paper, we
43 provide a detailed description on how ERA5 kernels are generated and considerations to ensure
44 proper use of them in feedback quantifications.

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48 1. Introduction

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50 Radiative kernels measure the sensitivity of radiative fluxes to the perturbation of feedback
51 variables, such as temperature, water vapor, albedo and cloud (e.g., Soden and Held, 2006;
52 Huang et al., 2007; Shell et al., 2008; Previdi, 2010; Zelinka et al., 2012; Block and Mauritsen,
53 2013; Yue et al., 2016; Huang et al., 2017; Pendergrass et al., 2018; Thorsen et al., 2018; Kramer
54 et al., 2019b; Smith et al., 2020). Compared to the partial radiative perturbation method (e.g.,
55 Wetherald and Manabe, 1988), which is precise but computationally expensive, the kernel
56 method deploys a set of precalculated radiative kernels with simple arithmetic multiplications in
57 feedback quantification and thus is computationally highly efficient, which has greatly facilitated
58 the analysis of radiative feedbacks in global climate models (GCM) (e.g., Soden and Held, 2006;
59 Soden et al., 2008; Jonko et al., 2012; Vial et al., 2013; Zhang and Huang, 2014; Dong et al.,
60 2020; Zelinka et al., 2020; Chao and Dessler, 2021), as well as in observations (e.g., Dessler,
61 2010; Kolly and Huang, 2018; Zhang et al., 2019; Huang, H. et al., 2021). These analyses have
62 helped dissect and understand the climate sensitivity differences among the GCMs, such as those
63 in Coupled Model Intercomparison Projects, CMIP5 (Taylor et al., 2012) and CMIP6 (Eyring et
64 al., 2016). For example, Zelinka et al. (2020) attributed the higher climate sensitivity in the
65 CMIP6 models to their more positive extratropical cloud feedback. The kernel-enabled feedback
66 analyses have also provided insights in the energetics of the climate variations such as the El
67 Niño and Southern Oscillation (ENSO, e.g., Dessler et al., 2010; Kolly & Huang 2018; Huang et
68 al. 2021a), the Madden-Julian Oscillation (MJO, e.g., Zhang et al. 2019) and the Arctic sea ice
69 interannual variability (e.g., Huang et al., 2019), despite the approximate nature of the kernel
70 method and the known limits of its accuracy (e.g., Colman and Mcavaney, 1997; Huang and
71 Huang, 2021).

72 Multiple sets of radiative kernels have been developed to date, using different radiation
73 codes and based on different atmospheric state datasets ranging from GCMs to global reanalysis
74 and satellite datasets, for both non-cloud variables (e.g., Soden and Held, 2006; Shell et al.,
75 2008; Huang et al., 2017; Thorsen et al., 2018; Bright and O'halloran, 2019; Donohoe et al.,
76 2020) and cloud properties (e.g., Zelinka et al., 2012; Zhou et al., 2013; Yue et al., 2016; Zhang
77 et al., 2021; Zhou et al., 2022). As the conventional feedback analyses are mostly concerned with
78 the radiation energy budget change at the TOA, most existing kernels have been developed and
79 tested to address that need, i.e., to measure the feedback contributions to the TOA radiation
80 changes. Although the radiative sensitivity depends on the atmospheric states as well as the
81 radiative transfer codes used to compute the kernel values (e.g., Collins et al., 2006; Huang and
82 Wang, 2019; Pincus et al., 2020), it has been noted that the global mean TOA feedback
83 quantification is insensitive to which kernel dataset is used, as the diagnosed feedback values are
84 close to each other when measured by different kernel datasets (e.g., Soden et al., 2008; Jonko et
85 al., 2012; Vial et al., 2013; Zelinka et al., 2020). However, as there is increasing interest in
86 regional climate change and associated feedback (e.g., Kolly and Huang, 2018; Huang et al.,
87 2019; Zhang et al. 2019), it becomes important to know how the kernels (dis)agree at regional
88 scales. The generation of the global radiative kernels usually requires radiative transfer
89 computation based on a large number of instantaneous atmospheric profiles. Due to this
90 computational cost, many kernel datasets are generated based on the atmospheric data from an
91 arbitrary calendar year. Given the known interannual climate differences, e.g., between El Niño
92 to La Niña years, this warrants investigations to ascertain whether the kernels may differ in
93 important ways for regional feedback assessments.

94 On the other hand, fewer feedback studies have addressed the surface radiation budget,
95 although its importance has been recognized for such problems as the precipitation change
96 (Previdi, 2010; Pendergrass and Hartmann, 2014; Myhre et al., 2018) and oceanic energy
97 transport (e.g., Zhang and Huang, 2014; Huang et al., 2017). The surface budget analysis
98 requires the use of surface kernels, which are not always available from the published kernel
99 datasets. Few of them have been subject to inter-comparisons or rigorous validation. As
100 explained below in this paper, the computation and use of them require different care than the
101 TOA kernels. Possibly due to the lack of such recognition, there exist considerable discrepancies
102 between the existing surface kernels and some surface budget-centered analyses reported
103 alarmingly large non-closure in their radiation budget analyses (e.g., Vargas Zeppetello et al.,
104 2019), calling into question the validity of kernel method for surface radiation budget analysis.
105 Hence, we are motivated to examine the radiative sensitivity quantified by different kernels,
106 especially for the surface budget.

107 In this work, we produce a new set of radiative kernels for both the TOA and surface
108 radiation fluxes based on the fifth generation European Center for Medium-Range Weather
109 Forecasts atmospheric reanalysis (ERA5, Hersbach et al., 2020), which demonstrates superior
110 accuracy in the quantification of various atmospheric states (e.g., Graham et al., 2019; Wright et
111 al., 2020), and document the key considerations in the kernel computation procedure. We
112 intercompare the kernels computed from ERA5 to the other previously generated ones, and
113 investigate the interannual variation of the kernel values due to their atmospheric state
114 dependency. In addition, applying a selected set of kernels to analyzing the feedback in the
115 CMIP6 models, we intercompare the discrepancies in quantified feedbacks across the GCMs and
116 across different kernels.

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119 **2. Construction of ERA5 radiative kernels**

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121 **2.1 Radiative transfer model and atmospheric dataset**

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123 We use the GCM version of the rapid radiative transfer model (RRTMG) (Mlawer et al.,
124 1997) to calculate the radiative kernels. RRTMG conducts radiative transfer calculations in 16
125 longwave (LW) spectral bands and 14 shortwave (SW) bands. The accuracy of this model has
126 been extensively validated against the line-by-line calculations (e.g., Collins et al, 2006).

127 Input data required by RRTMG, including surface pressure, skin temperature, air
128 temperature, water vapor, albedo, ozone concentration, cloud fraction, cloud liquid water content
129 and cloud ice content, are taken from the instantaneous (as opposed to monthly mean) data of the
130 ERA5 reanalysis, with a horizontal resolution of 2.5 degree by 2.5 degree and 37 vertical
131 pressure levels between 1 hPa and 1000 hPa. To ensure the accuracy of radiative kernels in upper
132 atmosphere (Smith et al., 2020), we patch five layers of the U.S. standard profile above 1 hPa in
133 the LW calculations. Other required input variables, such as the effective radii of cloud liquid
134 droplet and ice crystal are taken from the 3-hourly synoptic TOA and surface fluxes and cloud
135 product of the Clouds and Earth's Radiant Energy System (CERES) (Doelling et al., 2013) with
136 a horizontal resolution of 1 degree and then interpolated to the same resolution as the ERA5 data
137 (2.5 degree). A random cloud overlapping scheme is used in our all-sky calculation. Sensitivity
138 tests have been conducted to determine the necessary temporal sampling for a proper
139 representation of the diurnal cycle and 6-hourly and 3-hourly instantaneous profiles are adopted

140 for LW and SW radiative transfer calculations, respectively, to limit the root mean squared error
141 of the computed diurnal mean flux biases to less than one percent.

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144 **2.2 Radiative kernel computation**

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146 Radiative kernels in essence measure the change of radiative flux to unit perturbation of
147 atmospheric variables, i.e., $\frac{\partial R}{\partial X}$, where R is either the upwelling irradiance flux at the TOA or
148 upwelling/downwelling irradiance flux at the surface; X represents the aforementioned feedback
149 variables; K_X is the radiative kernel of variable X . Note that for each radiative flux, K_X varies
150 with the time, geographic and vertical locations of the perturbed variable and is in general a 4-
151 dimensional (4-D) data array. Note also that all radiative fluxes and kernel values are defined as
152 downward positive.

153 Following the previous studies, we compute non-cloud radiative kernels including the LW
154 kernels of surface temperature (T_s), air temperature (T_a), and water vapor ($WV\ LW$), and the SW
155 kernels of surface albedo (ALB) and water vapor ($WV\ SW$). To calculate the kernels, we use the
156 partial radiative perturbation experiments, conducting two radiative transfer computations, one
157 without perturbation (control run) and the other with a perturbation of one atmospheric variable;
158 the difference between these two computations is used to calculate the radiative kernel value. In
159 both experiments, the upward, downward and net radiative fluxes at the TOA and surface are
160 saved at each time instance and location. Then ΔR_0 can be obtained by differencing the saved
161 radiative fluxes between the perturbed and unperturbed experiments. Dividing ΔR_0 with the
162 perturbation of variable X (ΔX_0), the instantaneous radiative kernel K_X is calculated as

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$$164 \quad K_X = \frac{\Delta R_0}{\Delta X_0} \quad (1)$$

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166 Applying such perturbation computations to all the relevant variables (see Appendix for a
167 detailed discussion of the procedure), we obtain instantaneous radiative kernels of these
168 dimensionalities: the surface temperature and albedo kernels are 3-D arrays (time, latitude|73,
169 longitude|144), and the air temperature and water vapor kernels are 4-D arrays (time, level|37,
170 latitude|73, longitude|144).

171 To account for possible interannual variability of the radiative kernel values, we compute
172 the kernels using atmospheric data of five calendar years: from year 2011 to 2015. Among these
173 years, 2011 is a strong La Niña year, 2015 is a strong El Niño year. Monthly or annual mean
174 kernels are then averaged from the instantaneous computations. For example, the LW annual

175 mean kernel of 2011 is obtained as $K = \frac{1}{365*4} \sum_{i=1}^{365*4} K_i$ (365 is the number of days of a year and
176 4 is because 6-hourly data are used for LW calculations) and the SW kernels, $K =$

177 $\frac{1}{365*8} \sum_{i=1}^{365*8} K_i$ (8 is because 3-hourly data are used for SW calculations), where the index i
178 represents the time slices included in the averaging. Similar averaging procedure is applied to
179 monthly mean kernels. The analyses in this work are based on multi-year averaged monthly
180 mean kernels if not otherwise stated.

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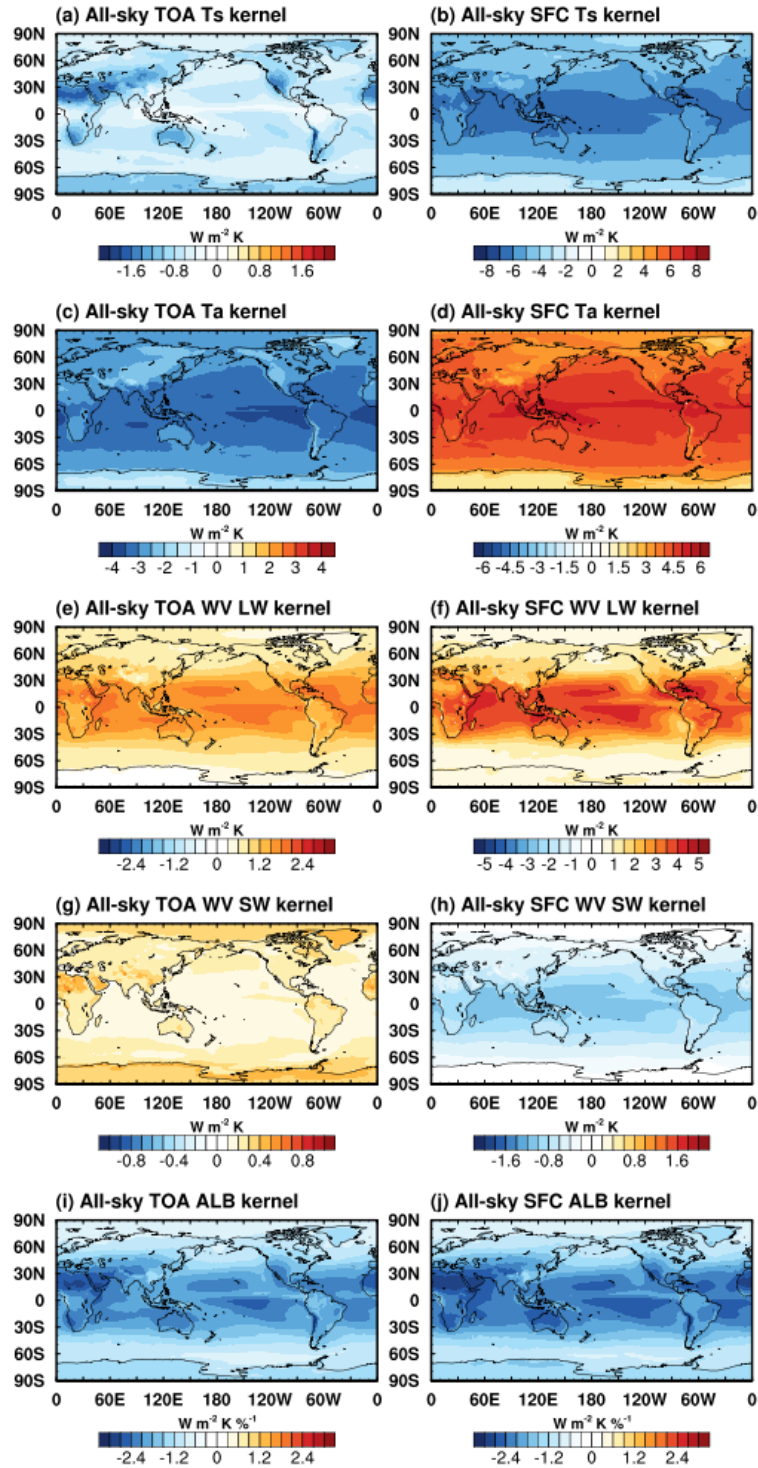
182 **3. Characterization of ERA5 kernels**

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In this section, we first present the all-sky TOA and surface radiative sensitivity kernels quantified from the ERA5 in Figure 1 to 4 (see the clear-sky kernels in Figure S1 and S2). The atmospheric radiation flux kernels, i.e., the change in radiation flux convergence in the atmosphere due to the perturbation of feedback variables and measured by differencing TOA and surface kernels, are shown in Figure S3 and S4 for interested readers. Then, we compare ERA5 kernels with the other kernel datasets and we examine the interannual variability of the ERA5 kernel values, due to the dependency of radiative sensitivity on the background atmospheric state.

3.1 Distribution of radiative sensitivity



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Figure 1. All-sky (left) TOA and (right) surface ERA5 kernels of (a, b) surface temperature (T_s), (c, d) air temperature (T_a), (e, f) water vapor longwave ($WV\ LW$), (g, h) water vapor shortwave ($WV\ SW$) and (i, j) surface albedo (ALB). Note that for T_a , $WV\ LW$, and $WV\ SW$ kernels, vertically integrated values are shown, which represents the sensitivity of radiative flux to a whole-column atmospheric perturbation. Note that the colorbar ranges differ among panels.

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Figure 1 summarizes the spatial distribution of all-sky ERA5 kernels for TOA and surface and Figure 2 illustrates the vertical cross-sections of zonal mean air temperature, water vapor LW and water vapor SW kernels in all-sky (see Figure S1 and S2 for results in clear-sky). For surface temperature kernel, an increase of surface temperature leads to more upwelling longwave radiation (i.e., OLR) both at the surface and TOA, therefore the kernel is negative. The TOA flux sensitivity in clear-sky (Figure S1a) is stronger than that in all-sky (Figure 1a) due to the absence of cloud, and the value increases with latitude, due to the decreasing concentration of water vapor from the tropics to the poles. The all-sky TOA sensitivity is strongly influenced by clouds, showing, for example, the fingerprint of the ITCZ in the tropical oceans (Figure 1a). The locations with less atmospheric absorption due to less water vapor or cloud, e.g., in the Tibetan Plateau and Sahara Desert regions, show relatively stronger sensitivity (Figure 1a). For the surface flux kernels, the increase of surface temperature enhances the upward emission according to the Planck function and thus the distribution follows that of surface temperature in both clear-sky and all-sky (Figure 1b).

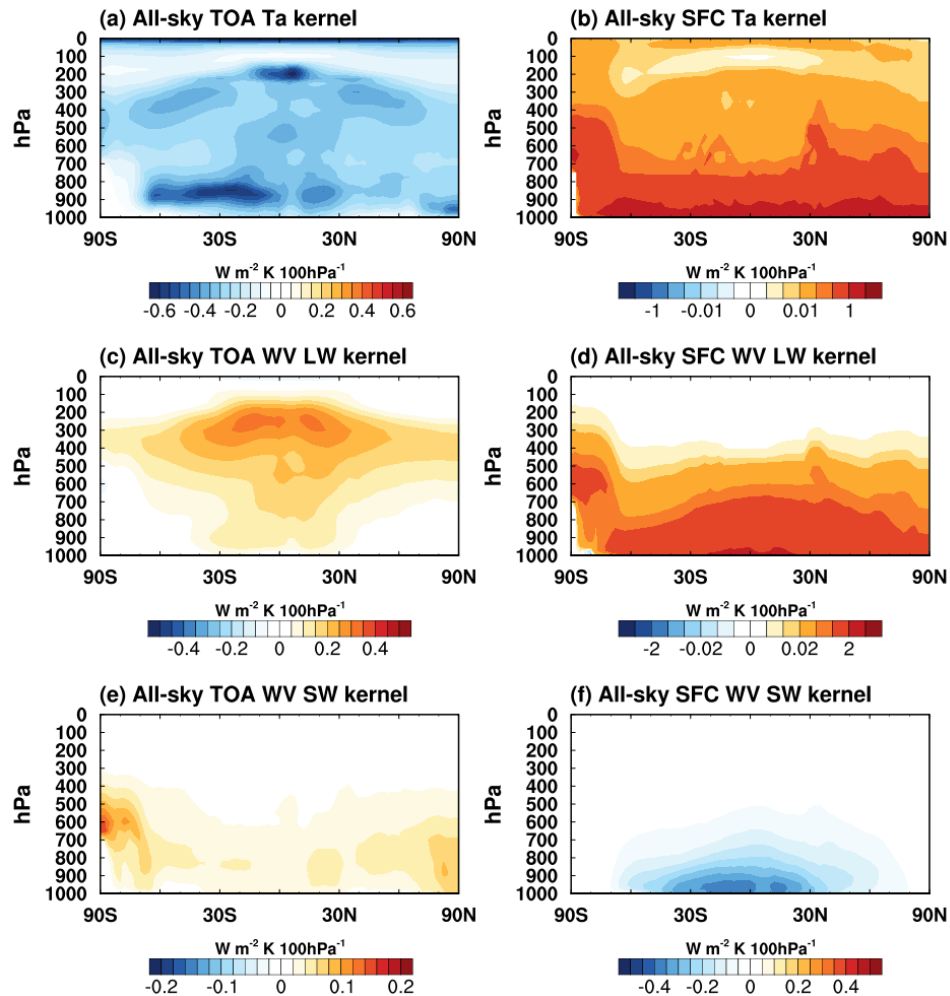
For air temperature kernel, the increase of air temperature increases the OLR at TOA and also the downwelling flux at surface, so the TOA and surface kernels take negative and positive signs, respectively. The TOA kernel has maximum values in the tropics, due to the higher air temperature (Planck function) and more abundant cloud and water vapor (higher emissivity) there, and generally decreases in magnitude with latitude (Figure 1c). Unlike the TOA flux kernel, which shows comparable sensitivity to air temperature at nearly all vertical levels, the surface flux is mainly sensitive to the bottom layers (Figure 2b).

For water vapor LW kernel, an increase of water vapor reduces OLR at TOA and increases downwelling radiation at surface, so that the TOA and surface kernels are both positive in sign. The vertically integrated kernel values (Figure 1e and f) generally follow the temperature distribution, for example, decreasing in magnitude with latitude. In both cases, the kernel magnitude is dampened by clouds in all-sky. The vertically resolved kernels show maximum sensitivity of TOA flux to the upper troposphere (Figure 2c) and maximum sensitivity of surface flux to the bottom layers (Figure 2d), respectively. In terms of the atmospheric radiation (the convergence of the TOA and surface radiation fluxes in the atmosphere), the increase in water vapor concentration absorbs more LW in the upper troposphere than what it emits but the opposite is true in the lower troposphere (Figure S4c). Such features were discussed in previous works (e.g., Huang et al. 2007).

For water vapor SW kernel, an increase of water vapor absorbs solar radiation and thus reduces both the upwelling (reflected) SW flux at TOA and the downwelling SW flux at surface. As a result, the two kernels take positive and negative signs, respectively. Note the magnitude of the SW kernels is much weaker than that of the LW kernels, because water vapor absorbs the LW flux more significantly than the SW flux. One noticeable feature of the TOA kernel in clear-sky (Figure S1g) is that the magnitude over the land is stronger than that over the ocean, because the relatively higher albedo over the land reflects more SW radiation and thus enhances the absorption by the water vapor in the atmosphere. For this reason, over reflective surfaces such as the Sahara Desert and Tibetan Plateau, as well as the Poles, the sensitivity is maximized. Unlike the TOA kernel, the distribution of surface kernel follows the distribution of background water vapor concentration, with noticeable dampening by clouds (Figure 1h and 2f).

For surface albedo kernel, an increase of surface albedo leads to more upwelling (reflected) SW flux both at surface and TOA; therefore, the kernel is of negative sign. In clear-

250 sky, the sensitivity strength follows the pattern of solar insolation, with some local maxima, e.g.,
 251 in the Sahara Desert and Tibetan Plateau (Figure S1i and j) due to the relatively lower water
 252 vapor concentration. In all-sky, the distribution is again influenced by cloud patterns; for
 253 example, in the ITCZ region, the strength is much reduced as clouds reduce the solar radiation
 254 reaching the surface and thus the sensitivity to surface albedo change (Figure 1i and j).
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256
 257 Figure 2. All-sky (left) TOA and (right) surface ERA5 vertically resolved and zonally
 258 averaged kernels of (a, b) air temperature (T_a), (c, d) water vapor longwave (WV LW) and (e, f)
 259 water vapor shortwave (WV SW), units: $W m^{-2} K^{-1} 100hPa^{-1}$. Note nonlinear colorbars are used
 260 for surface air temperature and water vapor LW kernels and the colorbar ranges differ among
 261 panels.
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264 3.2 Comparison of ERA5 kernels with other datasets

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 266 To examine the discrepancies between different kernel datasets, we select six previously
 267 published ones for comparison. Table 1 summarizes their resolutions and the datasets based on
 268 which they are computed, including the GCMs: GFDL (Soden et al., 2008), CAM3 (Shell et al.,
 269 2008), CAM5 (Pendergrass et al., 2018), and HadGEM3 (Smith et al., 2020), a global reanalysis:

270 ERAi (Huang et al., 2017), and satellite observations of cloud fields from CloudSat/CALIPSO
 271 combined with thermodynamic fields from reanalyses (Kramer et al., 2019b). This list is meant
 272 to be representative instead of exhaustive.

273
 274 Table 1. Summary of radiative kernels compared in this work. Datasets with * only have
 275 TOA kernels.

Radiative kernels	Horizontal resolution (lat*lon)	Vertical levels	Reference
GFDL*	2x2.5	17 (pressure level)	Soden et al., 2008
CAM3*	2.8x2.8	17 (pressure level)	Shell et al., 2008
ERAi	2.5x2.5	24 (pressure level)	Huang et al., 2017
CAM5	0.94x1.25	30 (hybrid level) or 17 (pressure level)	Pendergrass et al., 2018
CloudSat	2x2.5	17 (pressure level)	Kramer et al., 2019b
HadGEM3	1.25x1.9	85 (hybrid level) or 19 (pressure level)	Smith et al., 2020
ERA5	2.5x2.5	37 (pressure level)	This study

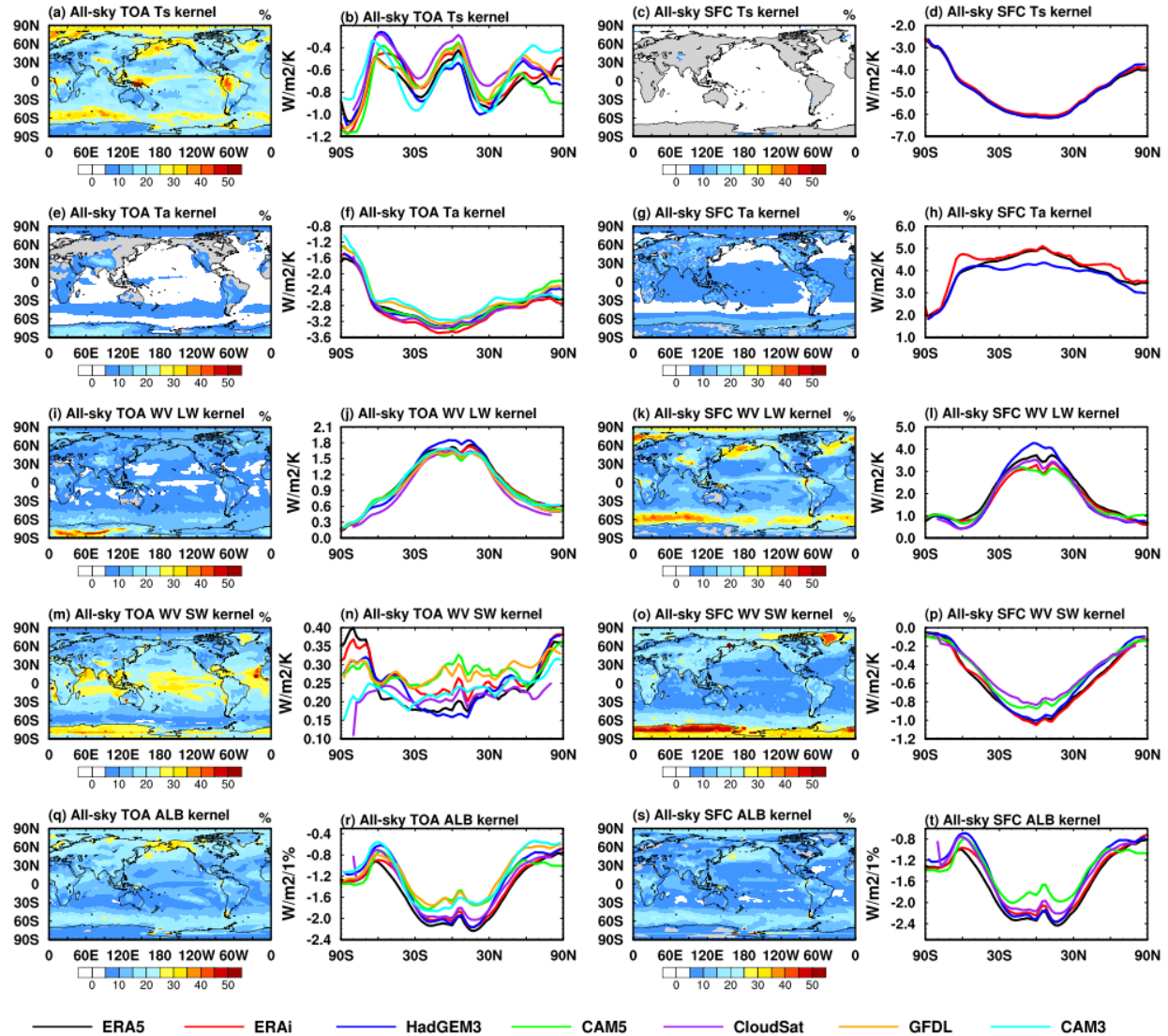
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 278 To facilitate an intercomparison, these kernel datasets are interpolated to the same
 279 horizontal and vertical resolutions as those of the ERA5 kernel when illustrated in Figure 3 and 4
 280 (see Figure S5 and S6 for clear-sky) and are uploaded to the same data repository of ERA5
 281 kernels. Note that the CAM5 and HadGEM3 kernels have two versions, with one defined at the
 282 raw hybrid levels and the other interpolated to pressure levels. To retain the accuracy of them as
 283 much as possible, the hybrid level version is used for the interpolation and comparison in Figures
 284 3 and 4, while in Section 4, the pressure level version is used for quantifying the feedbacks of
 285 CMIP6 models. The GFDL and CAM3 kernels are only available for TOA fluxes and are
 286 excluded for surface kernel comparisons.

287 Here we use the standard deviation (*std*) and its normalized value (*std**) to measure the
 288 spread of the inter-kernel dataset differences:

$$289 \quad std_x = \sqrt{\frac{1}{n-1} \sum_{i=1}^n (K_x^i - \bar{K}_x)^2} \quad (2)$$

$$290 \quad std_x^* = \frac{std_x}{\bar{K}_x} * 100 \quad (3)$$

291 where n is the total number of kernel datasets. K_x^i is radiative kernel of variable X from the i^{th}
 292 dataset. \bar{K}_x is the multi-dataset mean of radiative kernel K_x . Note that \bar{K}_x does not represent the
 293 “truth” value, but a reference value used to measure the spread of multi-kernel values. The
 294 vertically integrated and the vertically resolved but zonally averaged distributions of fractional
 295 discrepancy (*std**) are shown in Figures 3 and 4, respectively. The zonal mean kernel values
 296 from respective multi-datasets are shown in line plots in Figure 3 and 4. Note that some kernels
 297 exhibit abnormal values, such as the surface and air temperature kernel of the surface flux in the
 298 CAM5 and CloudSat kernels (see Appendix Figure A2), indicating inconsistent computation of
 299 their values, and thus are excluded in the corresponding *std** statistics in Figures 3 and 4. See
 300 more discussions in Appendix.

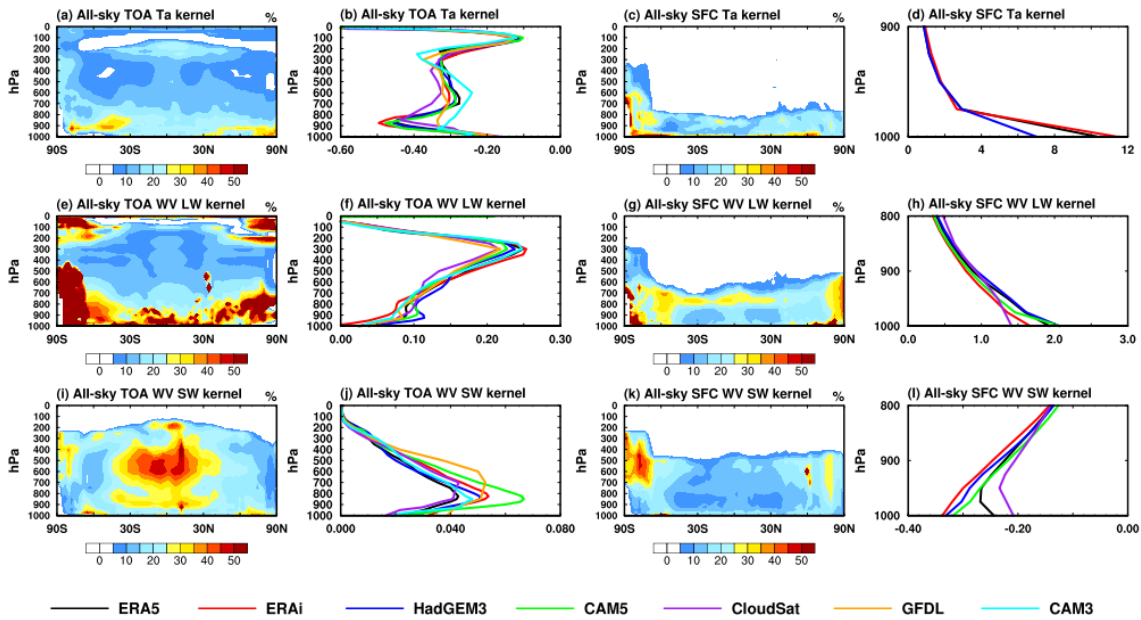


301
 302 Figure 3. (contour plot) fractional discrepancies as measured by the normalized standard
 303 deviation of the kernels by Eq. (3) and (line plot) zonal mean SFC distribution of multi-kernels in all-
 304 sky.

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 307 The comparisons identify the following relatively larger differences in kernel values.
 308 Among the TOA kernels, the surface temperature and albedo kernels show relatively large
 309 discrepancies in the Arctic, Southern Ocean and over some continental regions in the tropics in
 310 all-sky (Figure 3a and q), with the maximum discrepancy exceeding 30%; the air temperature
 311 kernel shows larger discrepancies in the lower troposphere and tropical tropopause region
 312 (Figure 4a); these kernel differences are likely due to the differences in cloud fields. The water
 313 vapor LW kernel also shows noticeable fractional differences, for example, over the Antarctic
 314 region (Figure 3i and 4e). The water vapor SW kernel shows differences in the tropical mid-
 315 troposphere and over Antarctic in both clear-sky and all-sky (Figure 4i and S6i), leading to
 316 strong variations in the vertical integration of sensitivity (Figure 3m and S5m), with a spread
 317 exceeding 30%. The noticeable periodic equatorial pattern in Figure S5m is caused by the CAM3

318 kernel, likely due to a coarser temporal resolution that does not well resolve the diurnal cycle of
 319 solar insolation in the kernel computation.

320 For the surface kernels, the most prominent differences exist in SW radiative kernels
 321 (Figure 3 and 4), especially in the polar regions. The discrepancy in the water vapor SW kernel
 322 reaches 30% for vertically integrated values (Figure 3o), with noticeable differences through the
 323 troposphere (Figure 4k). The surface albedo kernel differences are much larger in all-sky than
 324 that in clear-sky (Figure 3 and S5), indicating that the cause is in cloud fields, and are also
 325 noticeable in the Arctic region due to sea ice variations (Figure 3s). In the LW, the water vapor
 326 kernels exhibit noticeable differences in the Central Pacific, Southern Ocean and Arctic in all-
 327 sky (Figure 3k), where again the difference in cloud field is likely the cause. The air temperature
 328 kernels show noticeable discrepancies in the bottom layers (Figure 4d), which may be caused by
 329 inconsistency in the kernel computation and vertical resolutions (see the discussions in
 330 Appendix).
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 333 Figure 4. (contour plot) Cross-section of fraction discrepancies of the radiative kernels, (line
 334 plot) global mean vertically resolved kernels from multi-datasets in all-sky.
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336 In summary, the differences among radiative kernel datasets are generally smaller in clear-
 337 sky than in all-sky and in most cases, are mostly within 10% of the radiative kernel values.
 338 However, there are some notable regional discrepancies, for example, in the surface temperature
 339 kernel in the tropics (Figure 3a), in the surface albedo kernel in the Arctic (Figure 3q), and in the
 340 water vapor SW kernel in the Antarctic region (Figure 3m). As different kernel datasets are
 341 calculated using different data sources, the discrepancies detected here are likely due to the state-
 342 dependency in the kernels, which differ between the kernel datasets. To ascertain the state-
 343 dependency-caused kernel uncertainty, we next examine the ERA5 kernels computed from
 344 different years, i.e., from different atmospheric states, to investigate how much difference in
 345 radiative sensitivity can result from the change in atmospheric state.
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348 3.3 Interannual variation of kernel values

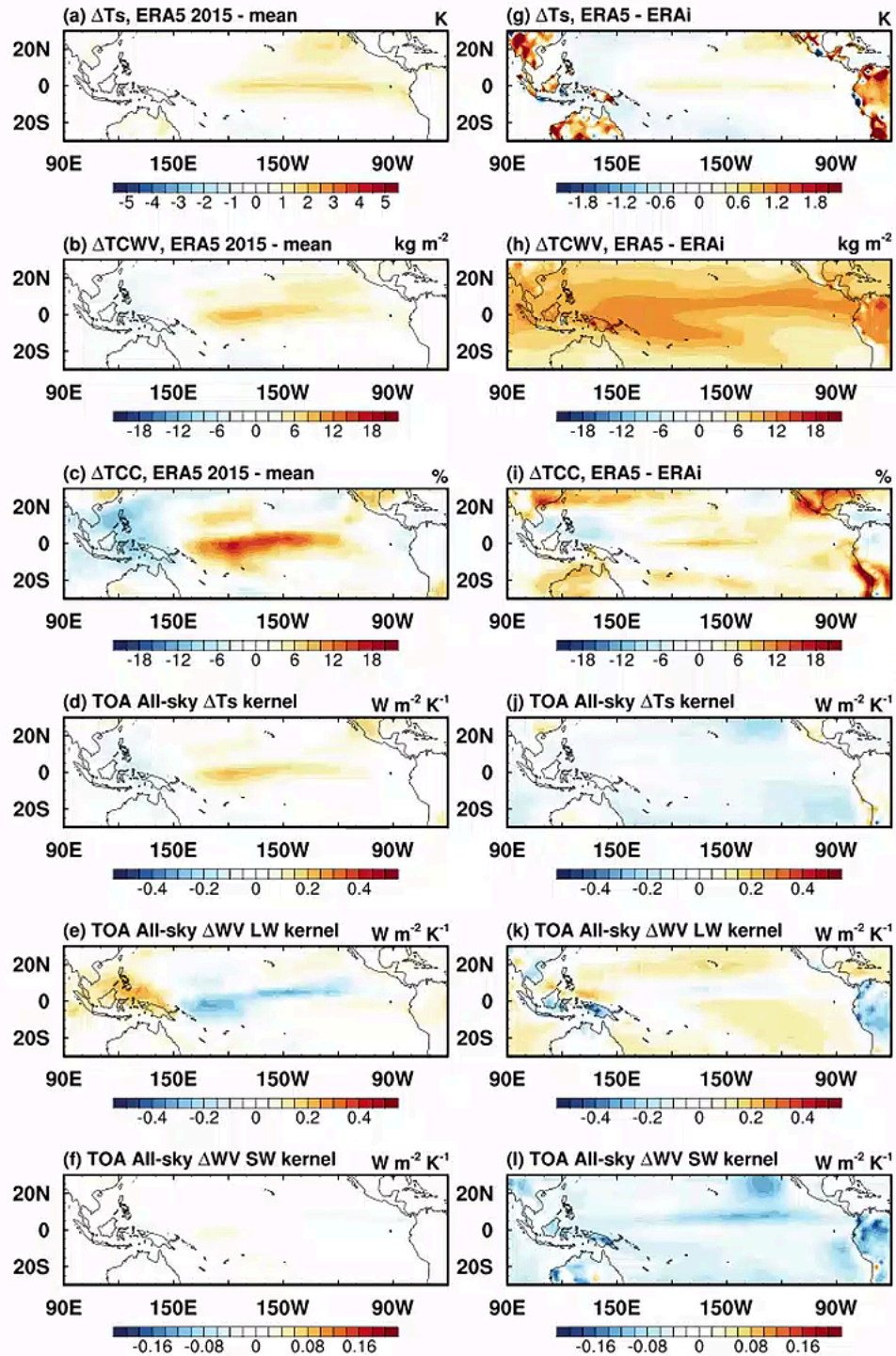
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350 The intercomparison above identified several prominent inter-dataset differences in the
351 kernel values. For example, there are noticeable differences in the values of surface temperature,
352 albedo and water vapor kernels in the Central Pacific and Arctic region. One possible reason that
353 may account for such differences is the atmospheric state-dependency of the kernel values.
354 Besides the inter-model differences in the different GCM climatology, interannual variations of
355 the atmospheric states, such as cloudiness variations in the Central Pacific region during the
356 ENSO cycle, may affect the radiative sensitivity as some radiative kernels are calculated using
357 one arbitrary year's data. To test this hypothesis, we use the ENSO and sea ice loss cases to
358 demonstrate the changes in radiative sensitivity with a focus on Central Pacific and Arctic
359 region, respectively. In the ENSO case, the variation is defined as the difference in annual mean
360 kernel values between 2015 and 5-year mean (from 2011 to 2015), which have the annual mean
361 sea surface temperature anomalies in the Niño 3.4 region (5N-5S, 190-240E) over +2.0K. In the
362 sea ice loss case, the variation is calculated as the difference in September between year 2012
363 and 2013, as the sea ice cover in 2012 was reported to be the lowest level in the satellite
364 observation era. In addition, we further show the comparison between ERA5 and ERAi kernels
365 (in Figure 5), which was also calculated by RRTMG and averaged from 5 years' calculations
366 (2008-2012), to compare the inter-kernel difference and interannual difference in kernel values.

367 To save space, here we only highlight the most prominent differences. Figure 5a-c show
368 the differences in skin temperature, total column water vapor and total cloud cover due to ENSO
369 and Figure 5d-f summarize the corresponding differences in all-sky TOA kernels. As the skin
370 temperature in the Central Pacific warms over 2K (Figure 5a) during ENSO, the increases in
371 water vapor concentration and cloud fraction (Figure 5b and c) reduce the sensitivity of TOA
372 flux to surface temperature change by about $0.2 \text{ W m}^{-2} \text{ K}^{-1}$ (about 33%) (Figure 5d). The
373 moistening in the Central Pacific (Figure 5b) enhances the TOA water vapor LW sensitivity in
374 clear-sky (Figure S7b), while in all-sky the enhanced convection and associated total cloud cover
375 in this region lead to a weakened TOA water vapor LW radiative sensitivity (Figure 5e) despite
376 the moist anomaly, and the decrease is almost contributed from the whole troposphere (Figure
377 S8c). The water vapor SW kernel discrepancy is less pronounced (Figure 5f).

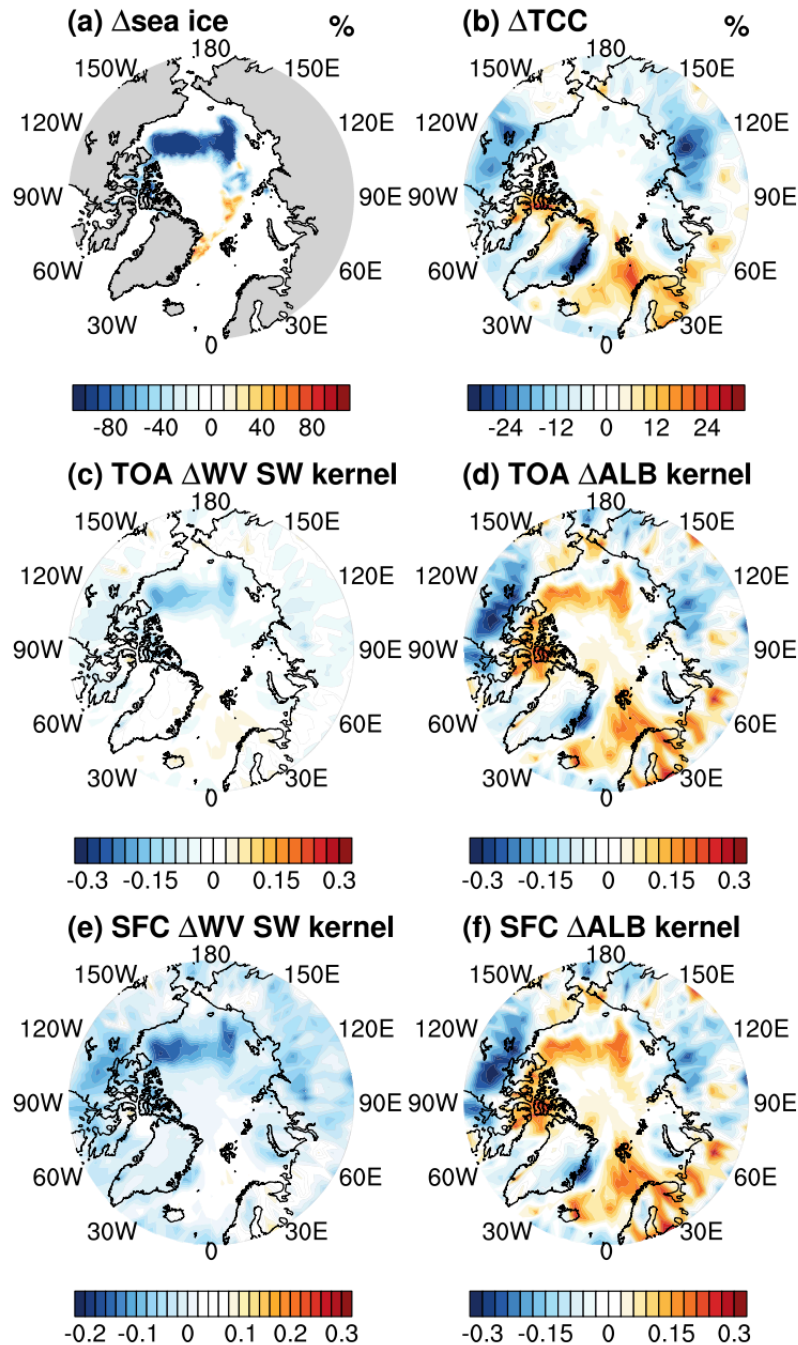
378 Comparing the 5-year averaged all-sky ERA5 and ERAi kernels, we find that the
379 atmospheric state differences also exist between the atmospheric datasets on which the kernels
380 are computed from. For example, the ERA5 shows similar, but less pronounced, warming
381 anomalies in sea surface temperature in the Central Pacific compared to ERAi, partly due to the
382 strong El Nino year (2015) included in the ERA5 dataset. ERA5 data also shows more water
383 vapor and cloud cover (Figure 5h and i). Although the total column water vapor and total cloud
384 cover are higher in the ERA5 (Figure 5h and i), their differences are complex and vertically non-
385 uniform (Figure S8 d and e), which leads to a slight strengthening of surface temperature kernels
386 compared with ERAi (Figure 5j). It is also noticed that the ERA5 water vapor SW kernel shows
387 lower sensitivity (Figure 5l), which mainly comes from the contributions in mid-to-low
388 troposphere (Figure S8f). The difference noticed in Figure S8 f corresponds to the discrepancy
389 noticed in Figure 4i, which are both in the mid-to-low troposphere, and the corresponding clear-
390 sky kernels are of much less differences (Figure S7), suggesting that the difference in clouds
391 might be the main cause of the all-sky kernel differences, which also correspond to the
392 discrepancies shown in the multi-kernel comparisons in Figure 3a, i, and m.

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Figure 5. Differences in climate states and all-sky kernel values (left) between an arbitrary year (2015) and a 5-year mean of ERA5 and (right) between the 5-year means of ERA5 and ERAi datasets: (a, g) skin temperature, (b, h) total column water vapor (TCWV), (c, i) total cloud cover (TCC), (d, j) TOA skin temperature kernel, (e, k) TOA vertically integrated water vapor LW kernel and (f, l) TOA vertically integrated water vapor SW kernel. Note that the colorbar ranges differ among panels.



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Figure 6. September differences between 2012 and 2013 in (a) sea ice concentration, (b) total cloud cover (TCC), and the differences in (c, e) water vapor SW kernel for TOA and surface fluxes, units: $W m^{-2} K^{-1}$, (d, f) surface albedo kernel for TOA and surface fluxes, units: $W m^{-2} 1\%^{-1}$. Note that the colorbar ranges differ among panels.

In the sea ice loss case, the reduction of sea ice in the Arctic region (Figure 6a) leads to a significant decrease of radiative sensitivity to surface albedo in the areas with noticeable sea ice retreats (Figure 6d and f), with the maximum difference exceeding 30% of the radiative kernel value, because of the nonlinear dependency of the reflected solar radiation on the surface albedo

412 (e.g., see Huang et al., 2021b, Fig. 3 and Fig. 6). The cloud cover changes also contribute to
 413 changes in surface albedo kernel values due to the coupling effect between cloud and surface
 414 albedo (see Huang et al., 2021b), which for example is seen in the Siberia and to the west
 415 coastline of Europe. The change in sea ice also leads to a significant decrease in the TOA
 416 sensitivity and an increase of surface sensitivity to water vapor in the sea loss region,
 417 respectively (Figure 6c and e), with the maximum changes exceeding 80% for surface. All these
 418 results confirm the state-dependency of radiative kernels (e.g., Riihelä et al., 2021).

419 In summary, these quantitatively large interannual differences, as well as their locations,
 420 verify that some discrepancies between the radiative kernels are caused by the difference in
 421 atmospheric states and partly explain the inter-dataset kernel differences seen in Figure 3 and 4.
 422 Nevertheless, it ought to be noted that the differences are localized and because of that do not
 423 cause significant differences in the global mean feedback values (see Section 4). The results
 424 above also show that kernel values based on one arbitrary year may be regionally different. If
 425 only one year's atmospheric profiles are used to generate radiative kernels, we recommend
 426 selecting a year without significant anomalies in atmospheric states, e.g., due to El Nino or
 427 severe sea ice loss, so that the computed kernel values better represent the radiative sensitivity
 428 climatology.

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431 **4. Feedback quantification**

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433 In this section, we apply different kernels to quantifying the radiative feedbacks in one
 434 quadrupling CO₂ experiment (abrupt4xCO₂) of CMIP6 models. This experiment is selected
 435 because it has been used by a number of studies for forcing and feedback analyses (e.g., Zelinka
 436 et al., 2020), which we can compare our results to. The CMIP6 models used in this assessment
 437 are listed in Table2. Note that the standard outputs at 19 pressure levels from the models and
 438 correspondingly the kernel values, including CAM5 and HadGEM3, provided at the pressure
 439 levels are used in this section.

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Table 2. Summary of CMIP6 models used in this study.

Models	Horizontal resolution (lat*lon)	Vertical levels	Reference
CESM2	0.9*1.25	32 levels to 2.26 hPa	Danabasoglu et al. (2020)
CNRM-CM6-1	1.4*1.4	91 levels to 0.01hPa	Voltaire et al. (2019)
EC-Earth3	0.7*0.7	91 levels to 90 km	Döscher et al. (2022)
HadGEM3-GC31-LL	1.25*1.875	85 levels to 85km	Williams et al. (2018)
IPSL-CM6A-LR	1.3*2.5	79 levels to 80km	Boucher et al. (2020)
MPI-ESM1-2-LR	1.875*1.875	47 levels to 0.01hPa	Mauritsen et al. (2019)

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444 **4.1 Analysis procedure**

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446 To quantify the radiative feedbacks, data from two experiments as documented by Eyring et al.
 447 (2016) and Pincus et al. (2016) are used: abrupt4xCO₂ simulations with an instantaneous

448 quadrupling of CO₂ concentration of year 1850 and piClim-4xCO₂ simulations with SST and sea
 449 ice concentrations fixed at the climatology of pre-industrial control experiment and CO₂
 450 concentration quadrupled. In each experiment, a 20-year period at the end of the simulation in
 451 each model is used. For example, in the models where the abrupt4xCO₂ simulation is longer than
 452 150 years, the simulations from the last 20 years rather than those from years 131 to 150 are used
 453 for the calculation. To exclude the effect of rapid adjustments, the radiative feedbacks in this
 454 study are measured using the difference of feedback variables between the abrupt4xCO₂ and
 455 piClim-4xCO₂ experiments and vertically integrated from the surface to model top. Note that
 456 these treatments are different from some other studies, e.g., Zelinka et al., 2020, which used
 457 piControl simulation as the climatology baseline and vertical integration from the surface to the
 458 tropopause, although the quantitative differences in the diagnosed global mean feedback values
 459 are small.

460 To detail the analysis procedure, firstly, all variables including radiative fluxes and
 461 atmospheric variables from CMIP6 models are interpolated to the horizontal and vertical
 462 resolution of the kernel itself. Note that for CAM3, GFDL, CloudSat and CAM5 kernels, they
 463 only have 17 pressure levels which are two layers (1hPa and 5hPa) fewer than the CMIP6
 464 standard model output. To address this issue, the contribution of the two missing layers is
 465 calculated using other kernels (e.g., ERA5) and found to have negligible effect on the global
 466 mean feedback value. Hence, when using these three kernels, the contributions from 10hPa
 467 above are ignored.

468 Secondly, the non-cloud radiative feedback of variable X (ΔR_X) is calculated as:

$$469 \quad \Delta R_X = K_X \cdot \Delta X \quad (4)$$

470 with units in W m⁻², where K_X is the monthly mean radiative kernel of variable X and ΔX is the
 471 monthly mean anomaly of X measured by the difference between abrupt4xCO₂ and piClim-
 472 4xCO₂, and represents the anomalies of surface temperature (ΔT_s), air temperature (ΔT_a), water
 473 vapor (ΔWV) and surface albedo (ΔALB). For the 2D radiative kernels (surface temperature and
 474 surface albedo), K_X and ΔX have just single layer values and ΔR_X is simply the product of these
 475 two terms. For the 3D radiative kernels (air temperature and water vapor), both K_X and ΔX are
 476 vectors of pressure levels and ΔR_X is the dot product of K_X and ΔX and is integrated from the
 477 TOA to 1000hPa. Note that if K_X is normalized with unit pressure thickness (e.g., W m⁻² K⁻¹
 478 100hPa⁻¹), the layer thickness must be taken into account when calculating dR_X . See Appendix
 479 for further discussion on the application of thickness-weighted kernels.

480 Finally, cloud feedbacks are diagnosed using the adjusted cloud-radiative forcing method
 481 (Shell et al., 2008). Here we compute the residual term in clear-sky as:

$$482 \quad res^o = \Delta R^o - \sum \Delta R_X^o \quad (5)$$

483 which represents the unexplained part of radiation budget change, and assuming the all-sky
 484 decomposition has the same non-closure residual, the cloud feedback, is measured as

$$485 \quad \Delta R_C = \Delta R - \sum \Delta R_X - res^o \quad (6)$$

486 where the superscript o represents clear-sky quantities. $\sum \Delta R_X^o$ and $\sum \Delta R_X$ are the sum of non-
 487 cloud feedbacks in clear-sky and all-sky, respectively, diagnosed by multiplying the radiative
 488 kernel with the atmospheric responses measured as the difference between abrupt4xCO₂ and
 489 piClim-4xCO₂ experiments. ΔR^o and ΔR are the total radiation change in clear-sky and all-sky,
 490 respectively, calculated as the difference in the GCM-simulated radiative fluxes between two
 491 experiments. It is worth noting that ΔR_C measured according to Eq. (6) is essentially the part of
 492 total radiation change not explained by the non-cloud feedbacks and is equivalent to the other
 493 formulations of the adjusted cloud radiative effect method (e.g., Shell et al. 2008; Soden et al.,

2008). Interested readers can refer to, for example, Huang (2013) for a detailed formulation and explanations of the method.

The feedback parameters, λ_X , in the units of $\text{W m}^{-2} \text{K}^{-1}$, are then obtained by normalizing the feedback flux changes ΔR_X by the global mean surface temperature change ΔT_S in the abrupt4xCO₂ experiment:

$$\lambda_X = \Delta R_X / \Delta T_S \quad (7)$$

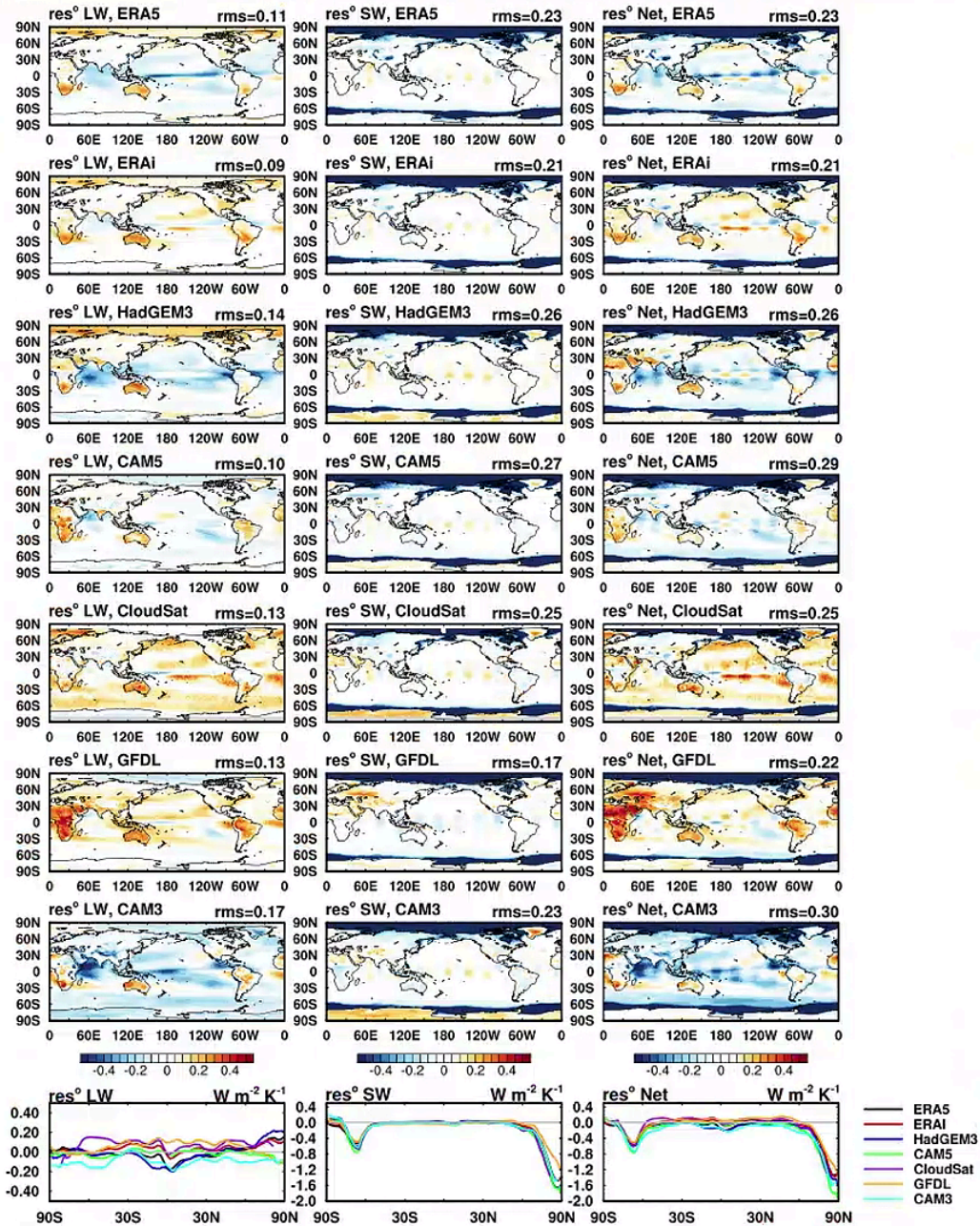
4.2 TOA feedbacks

The residual term (res^o) measures the unexplained radiation change in the feedback analysis and provides a useful overall indication of the soundness of the feedback quantification. Figure 7 illustrates the residual term for the TOA flux decomposition when different kernels are used to diagnose the multi-model mean feedbacks. In terms of the global mean, all residual terms are of small magnitude, no matter which kernel dataset is used (Figure 8 and Table S1). However, there are some noticeable local residuals, especially for the SW budget, e.g., in the Arctic region and around the Antarctic continent where sea ice changes the most (mid-column in Figure 7). While the non-zero magnitude of the residual is partly due to nonlinearity in the radiation decomposition, e.g., possible coupling between surface albedo and water vapor (Huang, Y. et al., 2021), the spread among the kernel results as evidenced by the line plots of Figure 7 is attributable to the discrepancies in the SW radiative kernels as revealed by the comparisons in Section 3. In the LW, the residual is generally small compared with the total feedback. In summary, the residual terms for the TOA budget are small in terms of the global mean feedback strengths, affirming the validity of the radiative kernels for feedback quantification. Here, we use the spatial root-mean square (RMS) of the residuals to quantify the regional biases, which are shown by the numbers on the right corner of each panel in Figure 7. For LW, results from ERA5, ERAi and CAM5 kernels show relatively smaller regional biases compared to those from HadGM3, CloudSat and CAM3 kernels. For SW, all kernel datasets have similar regional non-closures, for example, in the Polar regions (Figure 7 and 8). This is largely caused by the non-linearity in albedo feedback and also the coupling effect between water vapor and surface albedo feedbacks (Huang et al., 2021b; Block and Mauritsen, 2013). In summary, these results suggest that for the TOA feedback quantification, the performance of ERA5 kernel is comparable to the other datasets.

Figure 8 compares the spreads of feedback values resulted from the differences in kernels and those from the different projections of GCMs. In general, feedbacks from different kernel datasets overlap each other, even for cloud feedbacks, indicating a good consistency between the results computed from different kernel datasets. However, the spread across the GCMs is considerably larger, suggesting the overall feedback uncertainty is dominated by inter-model spread rather than the kernel uncertainty. The values of the feedbacks from each model and kernel datasets are shown in Table S1 and S2 for readers who are interested. These results are consistent with other published results. For example, compared with the results of Zelinka et al. (2020) based on the ERAi kernel, the kernel-diagnosed overall feedback parameter in the two results is $-0.87 \text{ W m}^{-2} \text{ K}^{-1}$ and $-0.85 \text{ W m}^{-2} \text{ K}^{-1}$ for the CNRM-CM6-1 model and $-0.81 \text{ W m}^{-2} \text{ K}^{-1}$ and $-0.84 \text{ W m}^{-2} \text{ K}^{-1}$ for the HadGEM3-GC3-LL model.

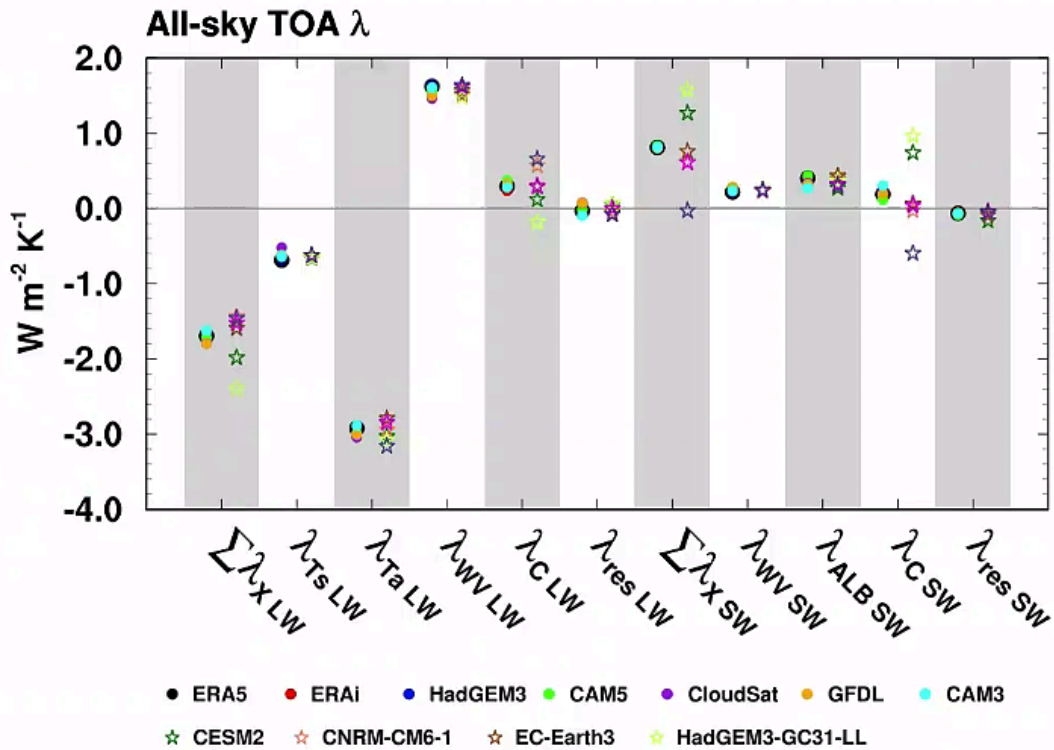
In summary, in terms of TOA feedback values, the inter-kernel differences lead to small uncertainty in the analyzed non-cloud feedbacks; the kernel-induced uncertainty in cloud

540 feedback is relatively larger (Table S2), with the inter-kernel spread in cloud LW feedback
 541 almost equally from the spread in surface and air temperature feedback and water vapor LW
 542 feedback (as measured by the $\sum \Delta R_X^0 - \sum \Delta R_X$ terms in Eq. (6)), and the inter-kernel spread in
 543 cloud SW feedback more from the spread in surface albedo feedback than from water vapor SW
 544 feedback (not shown); despite this, this uncertainty is considerably less than the inter-GCM
 545 cloud feedback spread.
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 549 Figure 7. The residuals (res^0) in the multi-model mean TOA feedback decomposition
 550 when different kernels are used, (left column) LW, (mid-column) SW, (right-column) Net, the

551 sum of LW and SW. The three line-plots in the bottom row are the zonal mean residuals.
 552 Numbers on the right corner in each panel are the spatial root-mean-square values.
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 555 Figure 8. Global mean TOA feedback parameters in all-sky diagnosed by the kernels listed in
 556 Table 1 across CMIP6 models. Dot marks represent multi-model mean values computed from
 557 different kernel datasets. Stars represent the multi-kernel mean results computed from different
 558 GCMs.

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 561 **4.3 Surface feedbacks**

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 563 Next, we examine how the inter-kernel differences lead to uncertainty in the analyzed
 564 surface feedbacks.

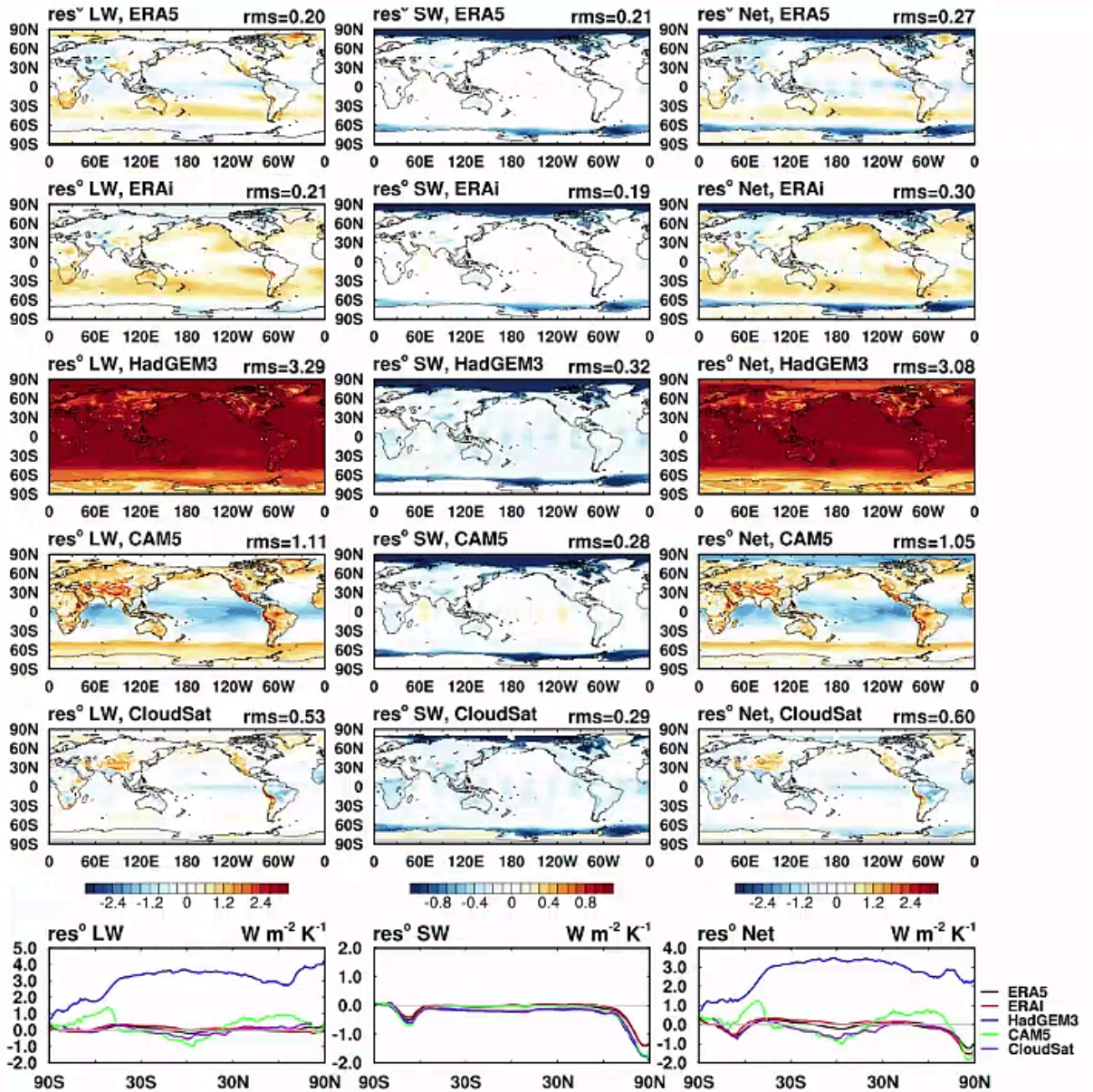
565 Figure 9 shows the residual distribution. We find that when the ERA5 and ERAi kernels
 566 are used for the feedback analysis, the non-closure residual in the surface budget is comparable
 567 in magnitude to the TOA analysis. This suggests that the surface kernels afford a valid tool for
 568 the surface feedback analysis. However, some prominent biases are noticed for other kernel
 569 datasets. For example, the HadGEM3 kernels, show especially an underestimation in air
 570 temperature feedback, likely due to a biased sensitivity of the bottom atmospheric layer (see
 571 Appendix for more discussions). The sum of global mean surface and air temperature feedback
 572 parameter measured by the HadGEM3 kernel is around $-3.70 \text{ W m}^{-2} \text{ K}^{-1}$ (Table S4, compared to
 573 around $-1.0 \text{ W m}^{-2} \text{ K}^{-1}$ measured by the other kernels), and the non-closure residual is as large as
 574 $3.0 \text{ W m}^{-2} \text{ K}^{-1}$ (Table S3, compared to $-0.1 \text{ W m}^{-2} \text{ K}^{-1}$ in the others). For this reason, the result
 575 from HadGEM3 kernel is excluded for the multi-kernel statistics in Figure 10, Table S3 and S4,
 576 but listed in a separate row for comparison. From either the spatial distribution of residual term

577 or the spatial RMS residuals, the ERA5 kernel and ERAi kernel show a superior performance
578 than other datasets. The use of ERA5 kernels may be advantageous for diagnosing the surface
579 radiation budget, considering that ERA5 data is a newer version reanalysis dataset from ECMWF
580 compared with ERAi and its data quality has been widely validated.

581 Figure 10 compares the inter-model and inter-kernel spreads for the surface feedbacks.
582 Unlike the results for TOA, the inter-kernel spread can be as large as the inter-model spread, for
583 example, in LW surface temperature feedback, air temperature feedback and water vapor
584 feedback. The sum of air temperature and surface temperature feedbacks shows better
585 consistency compared with the respective components (except for the HadGEM3 kernel), and the
586 respective air temperature and surface temperature feedbacks quantified by the ERA5 kernel are
587 stronger than the results from the other kernels. These discrepancies are due to the reason
588 discussed in the Appendix – a possibly wrong quantification of surface temperature effect. In
589 SW, the multi-kernel results are close to each other, showing smaller inter-kernel spreads than
590 the inter-model spreads.

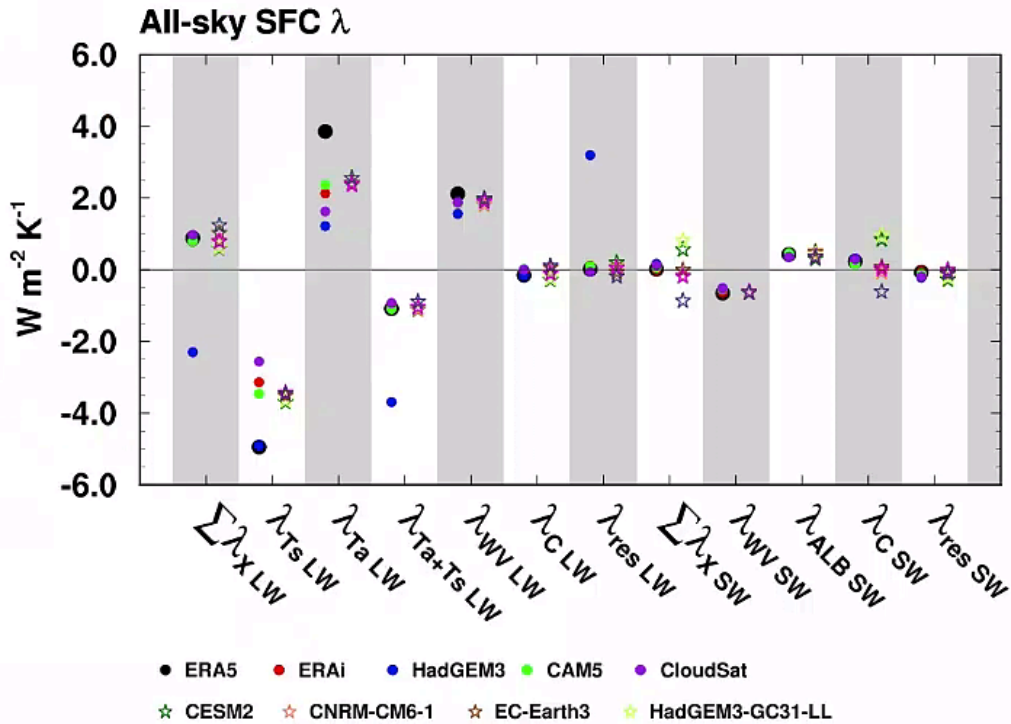
591 In summary, we find the surface feedback decomposition can achieve similar level of
592 radiation closure to the TOA analysis when using ERA5 kernels, affirming the validity of kernels
593 for diagnosing the surface radiative feedback. However, the results qualitatively vary depending
594 on which kernel dataset is used, indicating errors in the computation of some kernels.

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Figure 9. Similar to Figure 7, but for the surface feedback analysis.



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601 Figure 10. Similar to Figure 8, but for the surface feedback parameter.
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604 **5. Data availability**
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606 The datasets containing the multi-year averaged monthly mean TOA and surface kernels
607 for surface temperature, air temperature, surface albedo and water vapor (LW and SW) are
608 available at: <http://dx.doi.org/10.17632/vmg3s67568.3> (Huang and Huang, 2023). Other kernel
609 datasets used in this study, interpolated to the same horizontal and vertical grids of the ERA5
610 kernels, are also provided at this link.

611
612 **6. Conclusions and discussions**
613

614 In this paper, we present a newly generated set of ERA5-based radiative kernels of
615 surface and air temperatures, water vapor and surface albedo, for both TOA and surface radiation
616 fluxes. We also compare them with other published kernels, including the kernel-diagnosed
617 radiative feedbacks for both the TOA and surface radiation budgets.

618 For the TOA kernels, the results here demonstrate general consistency among the
619 different kernel datasets, and the discrepancies are generally within 10% in terms of vertically
620 integrated or globally averaged radiative sensitivity, although some relatively larger regional
621 differences are noticed, including those in the surface temperature kernel in the tropics (Figure
622 3a), those in the surface albedo kernel in the Arctic (Figure 3q) and those in the water vapor
623 shortwave kernel over Antarctica (Figure 3m), which is partly due to the dependence of radiative
624 sensitivity on background climate states.

625 For the surface kernels, more prominent inter-kernel differences are found. For example,
626 the differences in the water vapor shortwave kernel in the Antarctic (Figure 3o) and in the

627 surface albedo kernel in the Arctic (Figure 3s) can reach 30% of the kernel value itself. Some
628 kernels have considerably biased air temperature sensitivity values in the bottom atmospheric
629 layers, which is likely due to improper treatment in the perturbation experiments used for kernel
630 computation (see Appendix). The differences in both TOA and surface kernels discovered here
631 emphasize the importance of validating the radiative sensitivity as noted by Huang and Wang
632 (2019) and Pincus et al. (2020).

633 The investigation of interannual variability in ERA5 kernels validates the dependence of
634 radiative sensitivity on atmospheric state and the further comparison between ERAi and ERA5
635 kernel (Figure 5) reveals the effects of clouds on the kernel values, which might explain the
636 discrepancies of multi-kernel datasets (Figure 3).

637 Applying the different kernels to quantifying the TOA and surface radiative feedbacks,
638 we find that for TOA feedback quantification, the ERA5 kernels are as accurate as other kernel
639 datasets, while for surface feedback, ERA5 and ERAi kernels show superior accuracy compared
640 with other datasets. Considering the strengths of the ERA5 dataset in representing the
641 atmospheric states, we recommend the use of ERA5 kernels.

642 In addition, we compare the feedback differences caused by using different kernels and
643 also the inter-GCM spread of the feedback values (when measured by the same kernel). We find
644 the kernel difference is not a major cause of the inter-GCM TOA feedback spread (Figure 7 and
645 8). This finding is consistent with the previous assessments (e.g., Soden et al., 2008; Jonko et al.,
646 2012; Vial et al. 2013).

647 Radiation closure tests show that the unexplained residuals are generally within 10% of
648 the total feedback for both TOA and surface analyses in terms of the global mean feedback,
649 confirming the validity of the kernels for feedback quantification for both budgets. This suggests
650 that the large non-closure residuals reported in some previous studies (e.g., Vargas Zeppetello, et
651 al., 2019) are likely due to kernel inaccuracy rather than the limitation of the kernel method.
652 However, there are more significant local non-closures, for example, in the shortwave in the
653 Arctic region and around the Antarctic continent, which is contributed, but cannot be fully
654 explained, by the kernel uncertainty. This points to the accuracy limit of the kernel (linear)
655 method and calls for more advanced methods, such as the neural network method (Zhu et al.,
656 2019), for local feedback analysis.

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659 **Author contributions**

660

661 HH and YH led the design of this research and the writing of the manuscript. HH produced the
662 ERA5 radiative kernel and provided calculations of the inter-kernel comparison and feedback
663 analysis.

664

665 **Competing interests**

666 The authors declare that they have no conflict of interest.

667

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669

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825

826 **Appendix**

827

828 The ERA5 kernels are computed following Eq. (1) and the approach outlined in Section
829 2.2.

830

831 1. Surface variable kernels

832

833 To execute the partial radiative perturbation computations, the perturbations are prescribed
834 as the following: for the 2D feedback variables, the surface temperature is increased by 1 K and
835 the albedo is increased by 0.01 at each location. Hence, the units of the two kernels, K_{TS} and
836 K_{ALB} are $\text{W m}^{-2} \text{K}^{-1}$ and $\text{W m}^{-2} \%^{-1}$, respectively. When applying them to feedback
837 quantification, their feedbacks are quantified as

838
$$\Delta R_{TS} = K_{TS} \cdot \Delta T_S \quad (\text{A1})$$

839
$$\Delta R_{Alb} = K_{ALB} \cdot \Delta Alb \quad (\text{A2})$$

840 where ΔT_S should be measured in the units of K and ΔAlb in absolute values divided by 1%.

841

842 2. Water vapor kernel

843

844 For the 3D feedback variables, the perturbations are applied to each of the 37 pressure
845 layers (from 1hPa to 1000hPa) and one layer at a time. For the water vapor kernel, a 10%
846 incremental perturbation of the water vapor concentration is used. To adapt to the convention
847 used in the majority of the existing kernels, we convert the units of the kernels to represent the
848 radiative flux change corresponding to an increase of water vapor concentration that conserves
849 the relative humidity of the layer under a 1-K increase in air temperature, i.e., converting the
850 units from $\text{W}/(\text{m}^2 \Delta q_0^{+10\%} 100\text{hPa})$ to $\text{W}/(\text{m}^2 \Delta q_0^{+1K} 100\text{hPa})$:

851
$$K_q^{+10\%} = \frac{\Delta R_0}{\Delta q_0^{+10\%}} \quad (\text{A3})$$

852
$$K_q^{+1K} = \frac{\Delta R_0}{\Delta q_0^{+1K}} = K_q^{+10\%} \cdot \frac{\Delta q_0^{+10\%}}{\Delta q_0^{+1K}} = K_q^{+10\%} \cdot \frac{\Delta q_0^{+10\%}}{q_0} \cdot \frac{e_s(T_0)}{e_s(T_0+1K) - e_s(T_0)} \quad (\text{A4})$$

853 Where q_0 is the unperturbed water vapor concentration, in units of kg kg^{-1} . $\Delta q_0^{+10\%}$ is a 10%
854 increment in water vapor concentration. $e_s(T)$ is the saturated water vapor pressure under
855 temperature T , and can be measured by empirical formulas; hence, Δq_0^{+1K} can be measured as
856 $q_0 \left[\frac{e_s(T_0+1K)}{e_s(T_0)} - 1 \right]$. Accordingly, when the water vapor kernel is used for water vapor feedback
857 quantification, the feedback is measured as:

858
$$\Delta R_q = K_q^{+1K} \cdot \Delta q^{+1K} = K_q^{+1K} \cdot \frac{\Delta q}{\Delta q_0^{+1K}} = K_q^{+1K} \cdot \frac{\Delta q}{q_0} \cdot \frac{e_s(T_0)}{e_s(T_0+1K) - e_s(T_0)} \quad (\text{A5})$$

859 where $\Delta q = q - q_0$ measures the change in water vapor concentration and is normalized by
860 Δq_0^{+1K} to give the factor that is multipliable with the K_q^{+1K} kernel value. If using the Clapeyron-
861 Clausius relation, the above expression can be further approximated as

862
$$\Delta R_q = K_q^{+1K} \cdot \frac{\Delta q}{q_0} \cdot \frac{e_s}{(de_s/dT) \cdot 1K} = K_q^{+1K} \cdot \frac{\Delta q}{q_0} \cdot \frac{R_v}{L_v} \cdot \frac{T_0^2}{1K} \quad (\text{A6})$$

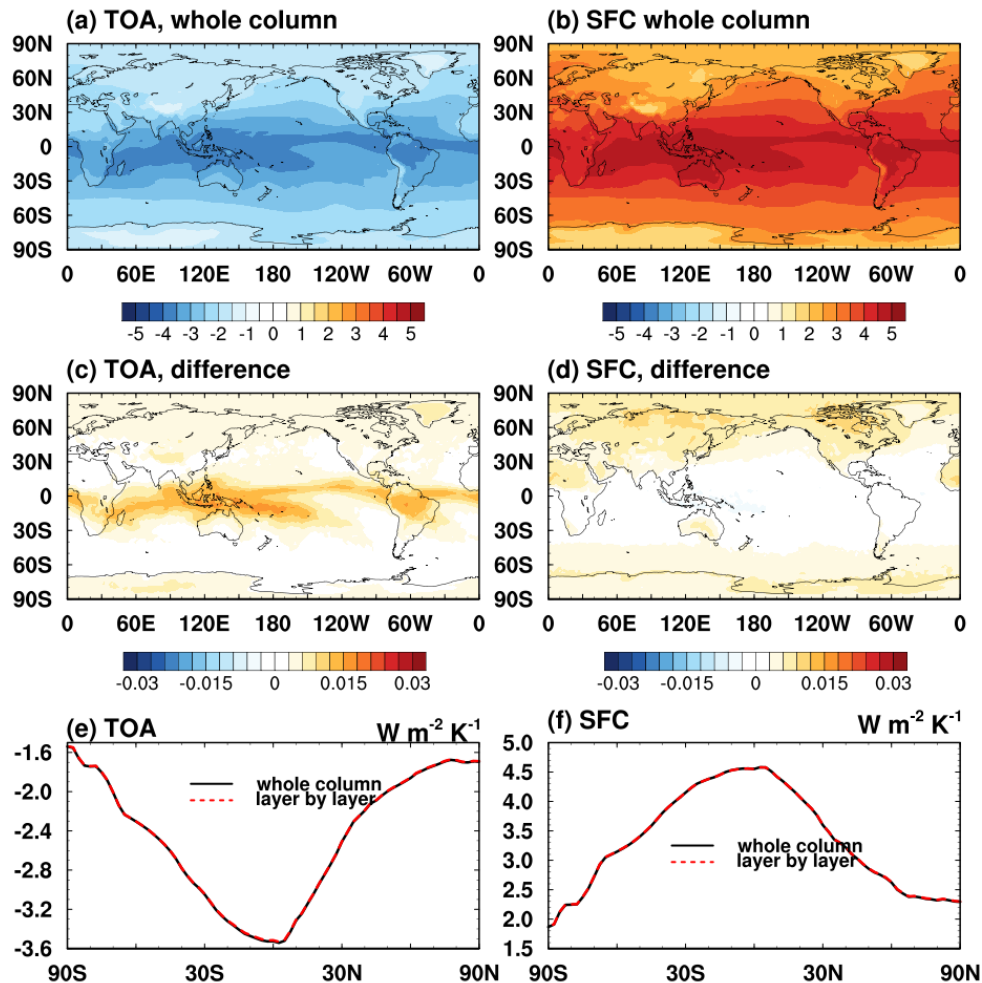
863 where R_v and L_v are the gas constant and specific latent heat of water vapor, respectively. Note
864 that when the kernels are used, T_0 and q_0 typically take their values from the base climate
865 appropriate to the application, e.g., the unperturbed climate of a GCM experiment, not
866 necessarily the dataset used for kernel computation.

867

868 3. Air temperature kernel
 869

870 For the air temperature kernel, to be consistent with the “inhomogeneous path treatment”
 871 that accounts for the vertically non-uniform temperature distribution within each discrete
 872 atmospheric layer (Mlawer et al., 1997), perturbations are added not only to the layer-mean
 873 temperature but also the temperature at the exiting boundary of radiative fluxes of interest (i.e.,
 874 the upper boundary of each layer for the TOA flux and the lower boundary for the surface flux),
 875 to appropriately represent the physical temperature perturbation in each layer.

876 A meaningful test to validate the validity of the air temperature kernel is a vertical sum
 877 test, i.e., a linear additivity test to verify the vertical integration of the kernel values reproduce
 878 the flux change, either at TOA or surface, in response to a whole-column air temperature
 879 increase of 1K. Figure A1 shows that the ERA5 kernel well passes this test. However, as shown
 880 by Figure 9, some kernels (e.g., HadGEM3 kernel) show much weaker radiative response at
 881 surface, possibly due to improper treatment of the air temperature perturbation in the kernel
 882 computation, which may lead to an underestimated air temperature feedback and large biases in
 883 the surface feedback analysis.
 884

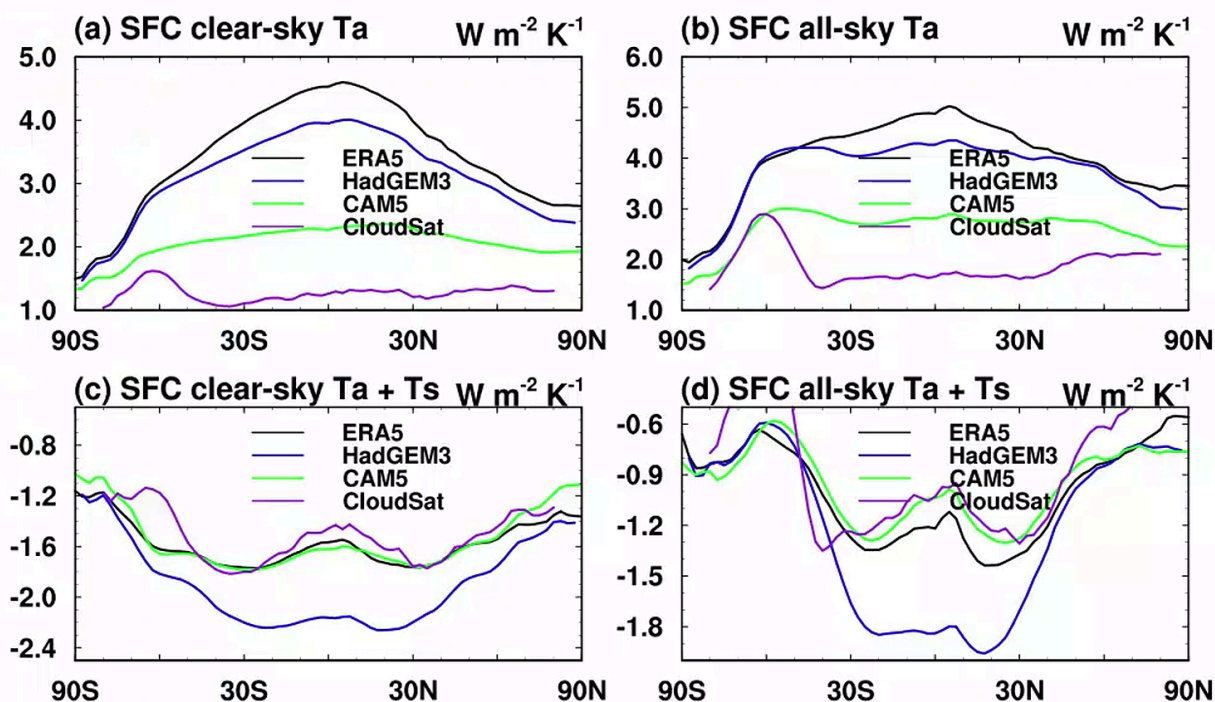


885
 886 Figure A1. Monthly mean TOA and surface radiation flux change in response to a +1K
 887 air temperature perturbation throughout the vertical column: (a, b) computed by a radiation

888 model, RRTMG; (c, d) difference of vertical sum of air temperature kernels compared to truth in
 889 (a, b); (e, f) comparison of the zonal mean.

890
 891 Another challenge in the computation of air temperature kernel for surface flux is that the
 892 surface in radiative transfer models is also the lower boundary of the lowermost atmospheric
 893 layer. If the effects of the surface temperature perturbation on the emission of the surface and
 894 that of the lowermost atmospheric layer are not distinguished, this may lead to improper
 895 interpretation and use of the surface temperature kernel. In our ERA5 kernel, the two effects are
 896 considered separately: according to radiative transfer theory, an increase in surface skin
 897 temperature only affects the surface upward emission; an increase in air temperature only affects
 898 the downward radiation. In some other kernels such as CAM5, these effects are not
 899 distinguished, so that the kernel value represents the net effect, i.e., change in the sum of both
 900 downward and upward. As a result, in Figure 10, we see stronger air temperature and surface
 901 temperature feedbacks quantified from ERA5 kernels than those from other kernels and in Table
 902 S4, we can only report the sum of surface and air temperature feedbacks.

903
 904 Figure A2 shows the comparison of vertically integrated air temperature kernels and the
 905 sum of surface and air temperature kernels between ERA5, CAM5, HadGEM3 and CloudSat.
 906 Although the strength of vertically integrated air temperature kernel for CAM5 is much weaker
 907 than that for ERA5 (Figure A2a and b), the sum of surface and air temperature kernel between
 908 these two datasets are in good agreement (Figure A2c and d), which warns that the seemingly
 909 right temperature feedback quantified by some kernels might come from the misattribution of
 910 surface temperature contributions. Another noticeable feature in Figure A2 is that the HadGEM3
 911 kernel shows an underestimation in vertical integration of air temperature kernel and an
 912 overestimation in the sum of surface and air temperature kernel, likely due to mistreatment of the
 913 bottom layer, and this accounts for the biased surface feedback analysis as shown in Figure 9.
 914 Similar issues were noticed in Kramer et al. (2019a).



915

916 Figure A2. Comparison of annual mean surface kernels for ERA5, CAM5, CloudSat and
 917 HadGEM3 for (a, b) the vertically integrated air temperature kernel values, and (c, d) sum of
 918 surface and air temperature kernels.
 919

920
 921 4. Time averaging
 922

923 As described in Section 2.2, all the kernels provided for feedback analysis are averaged
 924 from instantaneous kernel values over each calendar month and, in the ERA5 kernel, over
 925 multiple years. This is to ensure proper sampling of radiative sensitivity values under different
 926 atmospheric states, so that the kernels are representative of mean radiative sensitivity and thus
 927 can be readily multiplied with monthly mean climate responses (ΔX) to evaluate climate
 928 feedbacks.

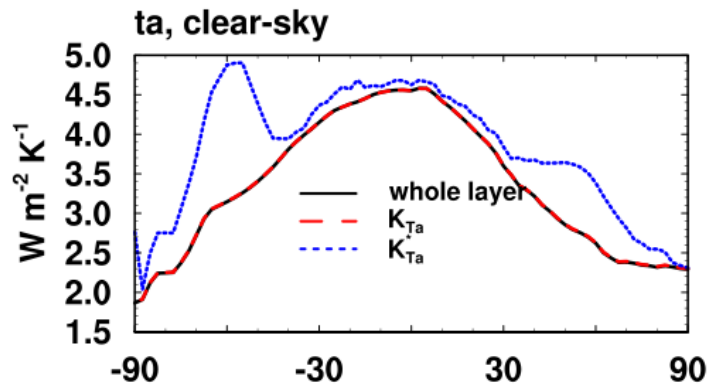
929 If the kernels are computed for fixed pressure levels, and if the pressure of any of these
 930 levels of an instantaneous atmospheric profile is higher than the surface pressure (i.e., the level is
 931 below the surface) at a time instance, this potentially creates inconsistency in the averaging
 932 procedure. To address this concern, we set the kernel value to zero (as opposed to missing value)
 933 before averaging. This is to ensure that when multiplied with the monthly mean climate response
 934 (ΔX), the contribution of a pressure layer (e.g., that centered at 1000 hPa) is effectively counted
 935 only for the fraction of time the layer exists (when surface pressure is higher than 1000 hPa).
 936 Otherwise, the feedback quantification needs to be further weighted with fraction of time (f)
 937 when the pressure layer exists. For example, if the surface pressure is larger than 1000hPa only
 938 for half of time in a month ($f=0.5$), the radiation flux anomaly contributed by the layer centered
 939 at 1000 hPa is:

940
 941
$$\Delta R_{T_{1000hPa}} = K_{T_{1000hPa}}^* \cdot \Delta T_{1000hPa} \cdot f \quad (A7)$$

 942

943 Here, $K_{T_{1000hPa}}^*$ represents the kernel value averaged from the time instances when the layer
 944 exists. Our averaging scheme is essentially to provide a kernel $K_{T_{1000hPa}} = K_{T_{1000hPa}}^* \cdot f$, so that
 945 it can be simply multiplied with $\Delta T_{1000hPa}$ to obtain the same result.

946 Figure A3 illustrates the differences between $K_{T_a}^*$ and K_{T_a} , in terms of their vertically
 947 integrated value. Such difference is pronounced over the Southern Oceans (around 60S), where
 948 the surface pressure value varies considerably. This likely explains why Figure 3h shows
 949 noticeable differences in the air temperature kernel in this region.



950

951 Figure A3. Zonal mean monthly mean air temperature kernels for surface flux from
952 ERA5 in clear-sky. Black line is the result from the whole column perturbation computation by
953 RRTMG, providing a "truth" for comparison. Red dashed line is the kernel weighted with
954 fraction of time (K_{T_a}) and blue dotted line represents results without weights ($K_{T_a}^*$).
955

956 5. Layer-specified and layer thickness-normalized radiative kernels 957

958 We generate two versions of vertically resolved air temperature kernel, water vapor LW
959 and SW kernel, one with values corresponding to specified vertical layers, i.e., in the units of W
960 $m^{-2} K^{-1}$, and another with unit-layer thickness (e.g., as shown in Figure 2 and 4), i.e., in $W m^{-2} K^{-1}$
961 $100hPa^{-1}$. The latter one properly portrays the vertical distribution of radiative sensitivity to
962 perturbations in unit thickness layers, while the former one may be more convenient to use in
963 feedback quantifications. For TOA budget analyses, these two versions of kernels lead to little
964 difference in practice due to limited contributions from the bottom atmospheric layer. However,
965 for surface budget analyses, we recommend using the layer-specified kernels, as the surface
966 kernels typically show strongest sensitivity to the perturbations in the bottom layers, which can
967 be best accounted for in the non-normalized kernels. Otherwise, the difference of surface
968 pressure between ERA5 and GCMs needs to be carefully treated to avoid errors, for example,
969 caused by missing the radiative contribution from the bottom layer of the atmosphere. To
970 illustrate this issue in an example, consider a location (latitude-longitude grid point) where the
971 surface pressure is 960 hPa in a GCM and the lowermost level of non-zero value of ERA5 air
972 temperature kernel is located at 975 hPa. Had the air temperature change been set to zero or NaN
973 value due to the GCM ground level being above 975 hPa, the contribution to the surface
974 radiation change from the air temperature change in the bottom layer of the atmosphere would
975 not be included, which may lead to a biased quantification of the feedback. We recommend
976 interpolating the air temperature changes from the GCM vertical coordinate to the kernel vertical
977 coordinate, using surface values to replace the missing levels (e.g., the 975 hPa level in the above
978 example) before multiplying with the kernel values, when computing the feedbacks of air
979 temperature and water vapor.

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