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2	Radiative sensitivity quantified by a new set of radiation flux kernels based
3	on the ERA5 reanalysis
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5	Han Huang, Yi Huang
6	Department of Atmospheric and Oceanic Sciences, McGill University, Montreal, Canada
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12	Corresponding Authors:
13	Han Huang, han.huang2@mcgill.ca (ORCID: 0000-0002-9143-6453),
14	Yi Huang, yi.huang@mcgill.ca (ORCID: 0000-0002-5065-4198)
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19 Abstract

- 20
- 21 Radiative sensitivity, i.e., the response of the radiative flux to climate perturbations, is essential
- 22 to understanding climate variability. The sensitivity kernels computed by radiative transfer
- 23 models have been broadly used for assessing the climate forcing and feedbacks <u>for</u> global
- 24 warming. As these assessments are largely focused on the top of atmosphere (TOA) radiation
- budget, less attention has been paid to the surface radiation budget or the associated surface
- radiative sensitivity kernels. Based on the fifth generation European Center for Medium-Range
 Weather Forecasts (ERA5) atmospheric reanalysis, we produce a new set of radiative kernels for
- 27 Weather Forecasts (ERA5) atmospheric reanalysis, we produce a new set of r 28 both the TOA and surface radiative fluxes, which is made available at
- 29 http://dx.doi.org/10.17632/vmg3s67568.2 (Huang and Huang, 2023). By comparing with other
- 30 published radiative kernels, we find that the TOA kernels are generally in agreement in terms of
- 31 global mean radiative sensitivity and analyzed overall feedback strength. The unexplained
- 32 residual in the radiation closure tests is found to be generally within 10%, no matter which kernel
- 33 dataset is used. The uncertainty in the TOA feedbacks caused by inter-kernel differences, as
- 34 measured by the standard deviation of the global mean feedback parameter value is much smaller
- 35 than the inter-climate model spread of the feedback values. However, we find relatively larger
- 36 discrepancies in the surface kernels. The newly generated ERA5 kernel outperforms many other
- 37 <u>datasets in closing the surface energy budget, achieving a radiation closure comparable to the</u>
- 38 TOA feedback decomposition, which affirms the validity of kernel method for the surface
- 39 radiation budget analysis. In this paper, we provide a detailed description on how ERA5 kernels
- 40 are generated and considerations to ensure proper use of them in feedback quantifications.
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44 **1. Introduction**

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46 Radiative kernels measure the sensitivity of radiative fluxes to the perturbation of feedback 47 variables, such as temperature, water vapor, albedo and cloud (e.g., Soden and Held, 2006; 48 Huang et al., 2007; Shell et al., 2008; Previdi, 2010; Zelinka et al., 2012; Block and Mauritsen, 49 2013; Yue et al., 2016; Huang et al., 2017; Pendergrass et al., 2018; Thorsen et al., 2018; Kramer 50 et al., 2019b; Smith et al., 2020). Compared to the partial radiative perturbation method (e.g., 51 Wetherald and Manabe, 1988), which is precise but computationally expensive, the kernel 52 method deploys a set of precalculated radiative kernels with simple arithmetic multiplications in 53 feedback quantification and thus is computationally highly efficient, which has greatly facilitated the analysis of radiative feedbacks in global climate models (GCM) (e.g., Soden et al., 2008; 54 55 Jonko et al., 2012; Vial et al., 2013; Zhang and Huang, 2014; Soden and Held, 2006; Dong et al., 56 2020; Zelinka et al., 2020; Chao and Dessler, 2021), as well as in observations (e.g., Dessler, 57 2010; Kolly and Huang, 2018; Zhang et al., 2019; Huang et al., 2021a). These analyses have 58 helped dissect and understand the climate sensitivity differences among the GCMs, such as those 59 in Coupled Model Intercomparison Projects, CMIP5 (Taylor et al., 2012) and CMIP6 (Eyring et 60 al., 2016). For example, Zelinka et al. (2020) attributed the higher climate sensitivity in the 61 CMIP6 models to their more positive extratropical cloud feedback. The kernel-enabled feedback 62 analyses have also provided insights in the energetics of the climate variations such as the El Nino and Southern Oscillation (ENSO, e.g., Dessler et al., 2010; Kolly & Huang 2018; Huang et 63 64 al. 2021a), the Madden-Julian Oscillation (MJO, e.g., Zhang et al. 2019) and the Arctic sea ice 65 interannual variability (e.g., Huang et al., 2019), despite the approximation nature of the kernel 66 method and the known limits of its accuracy (e.g., Colman and Mcavaney, 1997; Huang and 67 Huang, 2021). 68

Multiple sets of radiative kernels have been developed to date, using different radiation 69 codes and based on different atmospheric state datasets ranging from GCMs to global reanalysis 70 and satellite datasets, for both non-cloud variables (e.g., Soden and Held, 2006; Shell et al., 71 2008; Huang et al., 2017; Thorsen et al., 2018; Bright and O'halloran, 2019; Donohoe et al., 72 2020) and cloud properties (e.g., Zelinka et al., 2012; Zhou et al., 2013; Yue et al., 2016; Zhang 73 et al., 2021; Zhou et al., 2022). As the conventional feedback analyses are mostly concerned with 74 the radiation energy budget change at the TOA, most existing kernels have been developed and 75 tested to address that need, i.e., to measure the feedback contributions to the TOA radiation 76 changes. Although the radiative sensitivity depends on the atmospheric states as well as the 77 radiative transfer codes used to compute the kernel values (e.g., Collins et al., 2006; Huang and 78 Wang, 2019; Pincus et al., 2020), it has been noted that the global mean TOA feedback 79 quantification is insensitive to which kernel dataset is used, as the diagnosed feedback values are 80 close to each other when measured by different kernel datasets (e.g., Soden et al., 2008; Jonko et al., 2012; Vial et al., 2013). However, as there is increasing interest in regional climate change 81 82 and associated feedback (e.g., Kolly and Huang, 2018; Huang et al., 2019; Zhang et al. 2019), it 83 becomes important to know how the kernels (dis)agree at regional scales. The generation of the 84 global radiative kernels usually requires radiative transfer computation based on a large number 85 of instantaneous atmospheric profiles. Due to this computational cost, many kernel datasets are 86 generated based on the atmospheric data from an arbitrary calendar year. Given the known 87 interannual climate differences, e.g., between El Niño to La Niña years, this calls into question 88 whether the kernels may differ in important ways for regional feedback assessments.

89 On the other hand, fewer feedback studies have addressed the surface radiation budget,

- 90 although its importance has been recognized for such problems as the precipitation change
- 91 (Previdi, 2010; Pendergrass and Hartmann, 2014; Myhre et al., 2018) and oceanic energy
- 92 transport (e.g., Zhang and Huang, 2014; Huang et al., 2017). The surface budget analysis 93 requires the use of surface kernels, which are not always available from the published kern
- requires the use of surface kernels, which are not always available from the published kernel
 datasets. Few of them have been subject to inter-comparisons or rigorous validation. As
- 94 datasets. Few of them have been subject to inter-comparisons of figorous varidation. As 95 explained below in this paper, the computation and use of them require different care than the
- 96 TOA kernels. Possibly due to the lack of such recognition, there exist considerable discrepancies
- 97 between the existing surface kernels and some surface budget-centered analyses reported
- 98 alarmingly large non-closure in their radiation budget analyses (e.g., Vargas Zeppetello et al.,
- 99 2019), calling into question the validity of kernel method for surface radiation budget analysis.
- Hence, we are motivated to examine the radiative sensitivity quantified by different kernels,especially for the surface budget.
- 102 In this work, we produce a new set of radiative kernels for both the TOA and surface
- 103 radiation fluxes based on the fifth generation European Center for Medium-Range Weather
- 104 Forecasts atmospheric reanalysis (ERA5, Hersbach et al., 2020), which demonstrates superior
- accuracy in the quantification of various atmospheric states, and document the key
- 106 considerations in the kernel computation procedure. We intercompare the kernels computed from
- 107 <u>ERA5</u> to the other <u>previously generated</u> ones, and investigate the interannual variation of the 108 kernel values due to their atmospheric state dependency. In addition, applying a selected sets of
- kernels to analyzing the feedback in the CMIP6 models, we intercompare the discrepancies in
- 110 quantified feedbacks across the GCMs and across different kernels.
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113 2. Construction of ERA5 radiative kernels114

115 **2.1 Radiative transfer model and atmospheric dataset**

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We use the GCM version of the rapid radiative transfer model (RRTMG) (Mlawer et al.,
1997) to calculate the radiative kernels. RRTMG conducts radiative transfer calculations in 16
longwave (LW) spectral bands and 14 shortwave (SW) bands. The accuracy of this model has
been extensively validated against the line-by-line calculations (e.g., Collins et al, 2006).
Input data required by RRTMG, including surface pressure, skin temperature, air

122 temperature, water vapor, albedo, ozone concentration, cloud fraction, cloud liquid water content

123 and cloud ice content, are taken from the instantaneous (as opposed to monthly mean) data of the

- 124 ERA5 reanalysis, with a horizontal resolution of 2.5 degree by 2.5 degree and 37 vertical
- pressure levels <u>between 1 hPa and 1000 hPa</u>. To ensure the accuracy of radiative kernels in upper atmosphere (Smith et al., 2020), we patch five layers of the U.S. standard profile above 1 hPa in
- 127 the LW calculations. Other required input variables, such as the effective radii of cloud liquid
- 128 droplet and ice crystal are taken from the 3-hourly synoptic TOA and surface fluxes and cloud
- product of the Clouds and Earth's Radiant Energy System (CERES) (Doelling et al., 2013) with
- 130 <u>a horizontal resolution of 1 degree and then interpolated to the same resolution as the ERA5 data</u>
- 131 (2.5 degree). A random cloud overlapping scheme is used in our all-sky calculation. Sensitivity
- 132 tests have been conducted to determine the necessary temporal sampling for a proper
- representation of the diurnal cycle and 6-hourly and 3-hourly instantaneous profiles are adopted

- 134 for LW and SW radiative transfer calculations, respectively, to limit the root mean squared error 135 of the computed diurnal mean flux biases to less than one percent.
- 136 137

138 2.2 Radiative kernel computation139

140 Radiative kernels in essence measure the change of radiative flux to unit perturbation of 141 atmospheric variables, i.e., $\frac{\partial R}{\partial X}$, where *R* is either the upwelling irradiance flux at the TOA or 142 <u>upwelling/</u>downwelling irradiance flux at the surface; *X* represents the aforementioned feedback 143 variables; K_X is the radiative kernel of variable *X*. Note that for each radiative flux, K_X varies 144 with the time, geographic and vertical locations of the perturbed variable and is in general a 4-145 dimensional (4-D) data array. Note also that all radiative fluxes and kernel values are defined as 146 downward positive.

147 Following the previous studies, we compute non-cloud radiative kernels including the LW 148 kernels of surface temperature (T_s) , air temperature (T_a) , and water vapor (WV LW), and the SW 149 kernels of surface albedo (ALB) and water vapor (WV SW). To calculate the kernels, we use the 150 partial radiative perturbation experiments, conducting two radiative transfer simulations, one 151 without perturbation (control run) and the other with a perturbation of one atmospheric variable; 152 the difference between these two simulations is used to calculate radiative kernel value. In both experiments, the upward, downward and net radiative fluxes at the TOA and surface are saved at 153 154 each time instance and location. Then ΔR_0 can be obtained by differencing the saved radiative 155 fluxes between the perturbed and unperturbed experiments. Dividing ΔR_0 with the perturbation 156 of variable X (ΔX_0), the instantaneous radiative kernel K_X is calculated as

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 $K_X = \frac{\Delta R_0}{\Delta X_0} \tag{1}$

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Applying such perturbation computations to all the relevant variables (see Appendix for a detailed discussion of the procedure), we obtain instantaneous radiative kernels of these dimensionalities: the surface temperature and albedo kernels are 3-D arrays (time, latitude|73, longitude|144), and the air temperature and water vapor kernels are 4-D arrays (time, level|37, latitude|73, longitude|144).

To account for possible interannual variability of the radiative kernel values, we compute the kernels using atmospheric data of five calendar years: from year 2011 to 2015. Among these years, 2011 is a strong La Niña year, 2015 is a strong El Niño year. Monthly or annual mean kernels are then averaged from the instantaneous computations. For example, the LW annual mean kernel of 2011 is obtained as $K = \frac{1}{365*4} \sum_{i=1}^{365*4} K_i$ (365 is the number of days of a year and 4 is because 6-hourly data are used for LW calculations) and the SW kernels, K = $\frac{1}{365*8} \sum_{i=1}^{365*8} K_i$ (8 is because 3-hourly data are used for SW calculations), where the index *i*

represents the time slices included in the averaging. The analyses in this work are based on multi-year mean kernels if not otherwise stated.

- 174
- 175 **3. Characterization of ERA5 kernels**
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In this section, we first present the <u>all-sky TOA and surface</u> radiative sensitivity <u>kernels</u>
quantified from the ERA5 in Figure 1 to 4 (the clear-sky kernels as well as the <u>atmospheric</u>
<u>kernels are shown in Figure S1-S4 for interested readers</u>). Then, we compare ERA5 kernels with
the other kernel datasets and we examine the interannual variability of the ERA5 kernel values,
due to the dependency of radiative sensitivity on the background atmospheric state. **3.1 Distribution of radiative sensitivity**

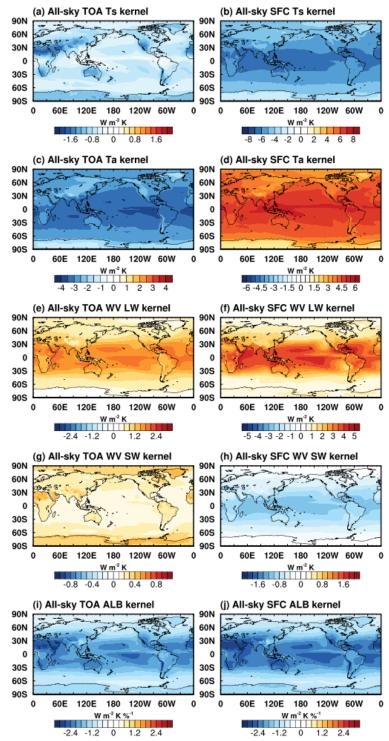


Figure 1. All-sky (left) TOA and (right) surface ERA5 kernels of (a, b) surface temperature (T_s), (c, d) air temperature (T_a), (e, f) water vapor longwave (WV LW), (g, h) water vapor shortwave (WV SW) and (i, j) surface albedo (ALB). Note that for T_a , WV LW, and WV SW kernels, vertically integrated values are shown, which represents the sensitivity of radiative flux to a whole-column atmospheric perturbation.

194 Figure 1 summarizes the spatial distribution of all-sky ERA5 kernels for TOA and surface

- and Figure 2 illustrates the vertical cross-sections of zonal mean air temperature, water vapor
- 196 LW and water vapor SW kernels in all-sky (see Figure S1 and S2 for results in clear-sky). For

surface temperature kernel, an increase of surface temperature leads to more <u>upwelling</u> longwave radiation (i.e., OLR) both at the surface and TOA, therefore the kernel is negative. The TOA flux

sensitivity in clear-sky (Figure S1a) is stronger than that in all-sky (Figure 1a) due to the absence

200 of cloud, and the value increases with latitude, due to the decreasing concentration of water

201 vapor from the tropics to the poles. The all-sky TOA sensitivity is strongly influenced by clouds,

showing, for example, the fingerprint of the ITCZ in the tropical oceans (Figure 1a). The

locations with less atmospheric absorption due to less water vapor or cloud, e.g., in the Tibetan
 Plateau and Sahara Desert regions, show relatively stronger sensitivity (Figure 1a). For the
 surface flux kernels, the increase of surface temperature enhances the upward emission
 according to the Planck function and thus the distribution follows that of surface temperature in

207 both clear-sky and all-sky (Figure 1b).

For air temperature kernel, the increase of air temperature increases the OLR at TOA and also the downwelling flux at surface, so the TOA and surface kernels take negative and positive signs, respectively. The TOA kernel has maximum values in the tropics, due to the higher air temperature (Planck function) and more abundant cloud and water vapor (higher emissivity) there, and generally decreases in magnitude with latitude (Figure 1c). Unlike the TOA flux kernel, which shows comparable sensitivity to air temperature at nearly all vertical levels, the surface flux is mainly sensitive to the bottom layers (Figure 2b).

215 For water vapor LW kernel, an increase of water vapor reduces OLR at TOA and increases 216 downwelling radiation at surface, so that the TOA and surface kernels are both positive in sign. 217 The vertically integrated kernel values (Figure 1e and f) generally follow the temperature 218 distribution, for example, decreasing in magnitude with latitude. In both cases, the kernel 219 magnitude is dampened by clouds in all-sky. The vertically resolved kernels show maximum 220 sensitivity of TOA flux to the upper troposphere (Figure 2c) and maximum sensitivity of surface 221 flux to the bottom layers (Figure 2d), respectively. In terms of the atmospheric radiation (the 222 convergence of the TOA and surface radiation fluxes in the atmosphere), the increase in water 223 vapor concentration absorbs more LW in the upper troposphere than what it emits but the 224 opposite is true in the lower troposphere (Figure S4c). Such features were discussed in previous

225 works (e.g., Huang et al. 2007).

226 For water vapor SW kernel, an increase of water vapor absorbs solar radiation and thus 227 reduces both the upwelling (reflected) SW flux at TOA and the downwelling SW flux at surface. 228 As a result, the two kernels take positive and negative signs, respectively. Note the magnitude of 229 the SW kernels is much weaker than that of the LW kernels, because water vapor absorbs the 230 LW flux more significantly than the SW flux. One noticeable feature of the TOA kernel in clear-231 sky (Figure S1g) is that the magnitude over the land is stronger than that over the ocean, because 232 the relatively higher albedo over the land reflects more SW radiation and thus enhances the 233 absorption by the water vapor in the atmosphere. For this reason, over reflective surfaces such as 234 the Sahara Desert and Tibetan Plateau, as well as the Poles, the sensitivity is maximized. Unlike 235 the TOA kernel, the distribution of surface kernel follows the distribution of background water

236 vapor concentration, with noticeable dampening by clouds (Figure 1h and 2f).

For surface albedo kernel, an increase of surface albedo leads to more upwelling

238 (reflected) SW flux both at surface and TOA; therefore, the kernel is of negative sign. In clear-

sky, the sensitivity strength follows the pattern of solar insolation, with some local maxima, e.g.,

240 in the Sahara Desert and Tibetan Plateau (Figure S1i and j) due to the relatively lower water

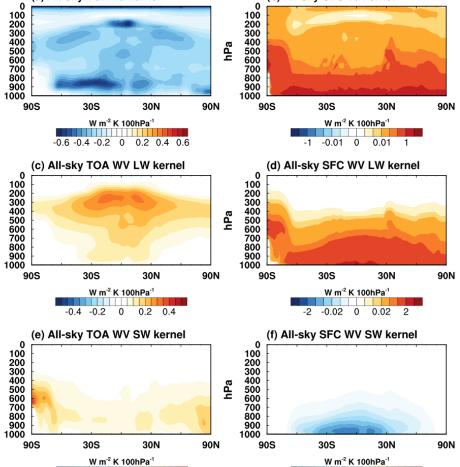
241 vapor concentration. In all-sky, the distribution is again influenced by cloud patterns; for

example, in the ITCZ region, the strength is much reduced as clouds reduce the solar radiation 242

- reaching the surface and thus the sensitivity to surface albedo change (Figure 1i and j). 243 244
 - (a) All-sky TOA Ta kernel 0 100 200 300 400 500 600 700 800 900 1000 hPa 90S 30S 30N W m⁻² K 100hPa⁻ -0.6 -0.4 -0.2 0

hPa

hPa



(b) All-sky SFC Ta kernel

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-0.2 0.2 -0.1 0 0.1 -0.4 -0.2 0 0.2 0.4 Figure 2. All-sky (left) TOA and (right) surface ERA5 vertically resolved and zonally <u>averaged kernels of</u> (a, b) air temperature (T_a) , (c, d) water vapor <u>longwave</u> (WV LW) and (e, f) 247 water vapor shortwave (WV SW), units: W m⁻² K⁻¹ 100hPa⁻¹. Note nonlinear colorbars used for 248 249 surface air temperature and water vapor LW kernels.

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252 3.2 Comparison of ERA5 kernels with other datasets 253

254 To examine the discrepancies between different kernel datasets, we select six previously 255 published ones for comparison. Table 1 summarizes their resolutions and the atmospheric 256 datasets based on which they are computed, including the GCMs: GFDL (Soden et al., 2008). 257 CAM3 (Shell et al., 2008), CAM5 (Pendergrass et al., 2018), and HadGEM3 (Smith et al., 2020),

258 a global reanalysis: ERAi (Huang et al., 2017), and satellite observations: CloudSat/CALIPSO

Table 1. Summary of radiative kernels compared in this work. Datasets with * only have
 TOA kernels.

Radiative kernels	Horizontal resolution (lat*lon)	Vertical resolution	Reference
GFDL*	2x2.5	17 (pressure level)	Soden et al., 2008
CAM3*	2.8x2.8	17 (pressure level)	Shell et al., 2008
ERAi	2.5x2.5	24 (pressure level)	Huang et al., 2017
CAM5	0.94x1.25	30 (hybrid level) or 17 (pressure level)	Pendergrass et al., 2018
<u>CloudSat</u>	<u>2x2.5</u>	<u>17 (pressure level)</u>	<u>Kramer et al., 2019</u> b
HadGEM3	1.25x1.9	85 (hybrid level) or 19 (pressure level)	Smith et al., 2020
ERA5	2.5x2.5	37 (pressure level)	This study

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265 To facilitate an intercomparison, these kernel datasets are interpolated to the same 266 horizontal and vertical resolutions as those of the ERA5 kernel when illustrated in Figure 3 and 4 267 (see Figure S5 and S6 for clear-sky) and are uploaded to the same data repository of ERA5 268 kernels. Note that the CAM5 and HadGEM3 kernels have two versions, with one defined at the 269 raw hybrid levels and the other interpolated to pressure levels. To retain the accuracy of them as 270 much as possible, the hybrid level version is used for the interpolation and comparison in Figures 271 3 and 4, while in Section 4, the pressure level version is used for quantifying the feedbacks of 272 CMIP6 models. The GFDL and CAM3 kernels are only available for TOA fluxes and are 273 excluded for surface kernel comparisons.

Here we use the standard deviation (std) and its normalized value (std^*) to measure the spread of the inter-kernel dataset <u>differences</u>:

$$std_{X} = \sqrt{\frac{1}{n-1}\sum_{i=1}^{n} \left(K_{X}^{i} - \overline{K_{X}}\right)^{2}} \quad (2)$$

$$std_X^* = \frac{std_X}{\overline{K_X}} * 100 \quad (3)$$

where n is the total number of kernel datasets. K_X^i is radiative kernel of variable X from the i^{th} 278 dataset. $\overline{K_X}$ is the multi-dataset mean of radiative kernel K_X . Note that $\overline{K_X}$ does not represent the 279 280 "truth" value, but a reference value used to measure the spread of multi-kernel values. The vertically integrated and the vertically resolved but zonally averaged distributions of fractional 281 282 discrepancy (std^*) are shown in Figures 3 and 4, respectively. The zonal mean kernel values 283 from respective multi-datasets are shown in line plots in Figure 3 and 4. Note that some kernels exhibit abnormal values, such as the surface and air temperature kernel of the surface flux in the 284 285 CAM5 and CloudSat kernels (see Appendix Figure A2), indicating inconsistent computation of their values, and thus are excluded in the corresponding std_x^* statistics in Figures 3 and 4. See 286 287 more discussions in Appendix.

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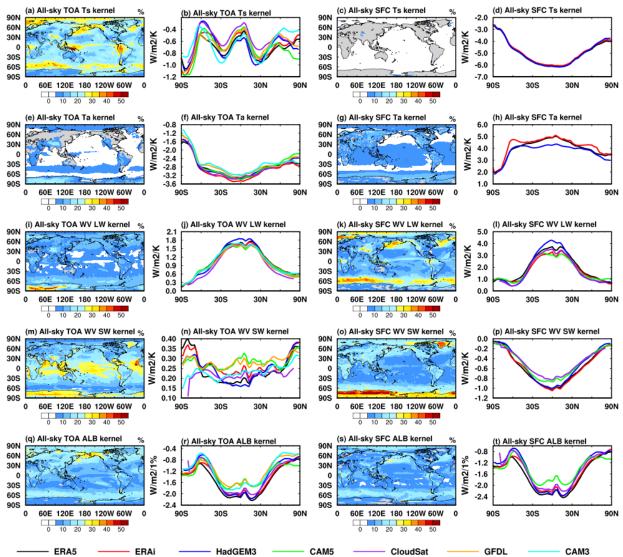


Figure 3. (contour plot) fraction discrepancies as measured by the normalized standard deviation of the kernels by Eq. (3) and (line plot) zonal mean distribution of multi-kernels in all-sky.

293 The comparisons identify the following relatively larger differences in kernel values. 294 Among the TOA kernels, the surface temperature and albedo kernels show relatively large 295 discrepancies in the Arctic, Southern Ocean and over some continental regions in the tropics in 296 all-sky (Figure 3a and q), with the maximum discrepancy exceeding 30%; the air temperature 297 kernel shows larger discrepancies in the lower troposphere and tropical tropopause region 298 (Figure 4a); these kernel differences are likely due to the differences in cloud fields. The water 299 vapor LW kernel also shows noticeable fractional biases, for example, over the Antarctic region 300 (Figure 3i and 4e). The water vapor SW kernel shows differences in the tropical mid-troposphere 301 and over Antarctic in both clear-sky and all-sky (Figure 4i and S6i), leading to strong variations in the vertical integration of sensitivity (Figure 3m and S5m), with a spread exceeding 30%. The 302

303 noticeable periodic equatorial pattern in Figure S5m is caused by the CAM3 kernel, likely due to

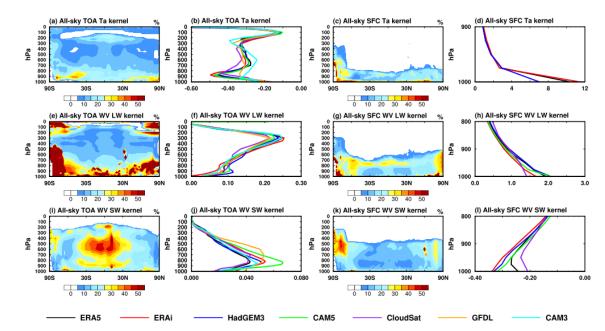
a coarser temporal resolution that does not well resolve the diurnal cycle of solar insolation in the
 kernel computation.

BOG For the surface kernels, the most prominent <u>difference</u>s exist in SW radiative kernels

307 (Figure 3 and 4), especially in the polar regions. The discrepancy in the water vapor SW kernel

reaches 30% for vertically integrated values (Figure $\underline{3}$ o), with noticeable biases through the

- β troposphere (Figure <u>4</u>k). The surface albedo kernel <u>differences</u> are much larger in all-sky than that in clear-sky (Figure 3 and S5), indicating that the cause is in cloud fields, and are also
- bit in clear-sky (Figure 5 and 55), indicating that the cause is in cloud fields, and are also bit noticeable in the Arctic region due to sea ice variations (Figure 3s). In the LW, the water vapor
- kernels exhibit noticeable differences in the Central Pacific, Southern Ocean and Arctic in all-
- sky (Figure 3k), where again the difference in cloud field is likely the cause. The air temperature
- kernels show noticeable discrepancies in the bottom layers (Figure 4d), which may be caused by
- β 15 inconsistency in the kernel computation and vertical resolutions (see the discussions in 316 Appendix).
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Figure 4. (contour plot) Cross-section of fraction discrepancies of the radiative kernels, (line plot) global mean vertically resolved kernels from multi-datasets in all-sky.

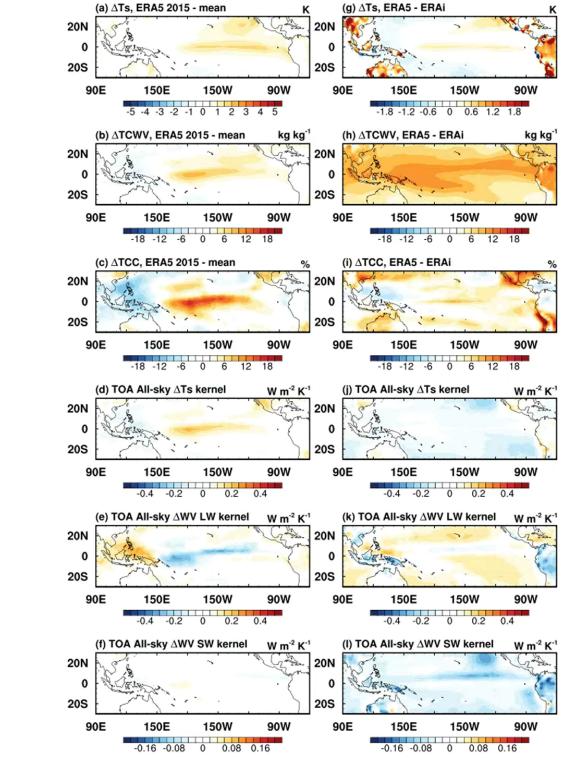
322 In summary, the differences among radiative kernel datasets are generally smaller in clear-323 sky than in all-sky and in most cases, are mostly within 10%. However, there are some notable 324 regional discrepancies, for example, in the surface temperature kernel in the tropics (Figure 3a), in the surface albedo kernel in the Arctic (Figure 3q), and in the water vapor SW kernel in the 325 326 Antarctic region (Figure 3m). As different kernel datasets are calculated using different data 327 sources, the discrepancies detected here are likely due to the state-dependency in the kernels, 328 which differ between the kernel datasets. To ascertain the state-dependency-caused kernel 329 uncertainty, we next examine the ERA5 kernels computed from different years, i.e., from 330 different atmospheric states, to investigate how much difference in radiative sensitivity can result 331 from the change in atmospheric state.

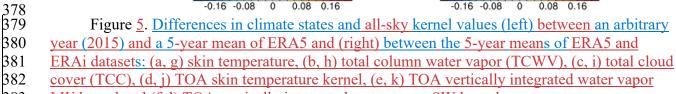
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334 **3.3 Interannual variation of kernel values**

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336 The intercomparison above identified several prominent inter-dataset differences in the 337 kernel values. For example, there are noticeable differences in the values of surface temperature, 338 albedo and water vapor kernels in the Central Pacific and Arctic region. One possible reason that 339 may account for such differences is the atmospheric state-dependency of the kernel values. 340 Besides the inter-model differences in the different GCM climatology, interannual variations of 341 the atmospheric states, such as cloudiness variations in the Central Pacific region during the 342 ENSO cycle, may affect the radiative sensitivity as some radiative kernels are calculated using 343 one arbitrary year's data. To test this hypothesis, we use the ENSO and sea ice loss cases to 344 demonstrate the changes in radiative sensitivity with a focus on Central Pacific and Arctic 345 region, respectively. In the ENSO case, the variation is defined as the difference in annual mean 346 kernel values between 2015 and 5-year mean (from 2011 to 2015), which have the annual mean 347 sea surface temperature anomalies in the Niño 3.4 region (5N-5S, 190-240E) over +2.0K. In the 348 sea ice loss case, the variation is calculated as the difference in September between year 2012 349 and 2013, as the sea ice cover in 2012 was reported to be the lowest level in the satellite 350 observation era. In addition, we further show the comparison between ERA5 and ERAi kernels 351 (in Figure 5), which was also calculated by RRTMG and averaged from 5 years' calculations 352 (2008-2012), to compare the inter-kernel difference and interannual difference in kernel values. 353 To save space, here we only highlight the most prominent differences. Figure 5a-c show 354 the differences in skin temperature, total column water vapor and total cloud cover due to ENSO 355 and Figure 5d-f summarize the corresponding differences in all-sky TOA kernels. As the skin 356 temperature in the Central Pacific warms over 2K (Figure 5a) during ENSO, the increases in 357 water vapor concentration and cloud fraction (Figure 5b and c) reduce the sensitivity of TOA flux to surface temperature change by about 0.2 W m⁻² K⁻¹ (about 33%) (Figure 5d). The 358 359 moistening in the Central Pacific (Figure 5b) enhances the TOA water vapor LW sensitivity in 360 clear-sky (Figure S7b), while in all-sky the enhanced convection and associated total cloud cover 361 in this region lead to a weakened TOA water vapor LW radiative sensitivity (Figure 5e) despite 362 the moist anomaly, and the decrease is almost contributed from the whole troposphere (Figure 363 S8c). The water vapor SW kernel discrepancy is less pronounced (Figure 5f). 364 Comparing the 5-year averaged all-sky ERA5 and ERAi kernels, we find that the 365 atmospheric state differences also exist between the atmospheric datasets on which the kernels 366 are computed from. For example, the ERA5 shows similar, but less pronounced, warming 367 anomalies in sea surface temperature in the Central Pacific compared to ERAi, partly due to the strong El Nino year (2015) included in the ERA5 dataset. ERA5 data also shows more water 368 369 vapor and cloud cover (Figure 5h and i). As a result, the surface temperature kernel computed 370 from ERA5 shows less TOA radiative sensitivity to surface temperature than that from ERAi. It 371 is also noticed that the ERA5 water vapor SW kernel shows lower sensitivity and mainly comes 372 from the contributions in mid-to-low troposphere (Figure S8f), which corresponds to the 373 discrepancy noticed in Figure 4i. The clear-sky kernels are of much less differences (Figure S7), 374 confirming that the difference in clouds is the main cause of the all-sky kernel differences, which 375 also correspond to the discrepancies shown in the multi-kernel comparisons in Figure 3a, i, and 376 m. 377





<u>LW kernel and (f, l) TOA vertically integrated water vapor SW kernel.</u>



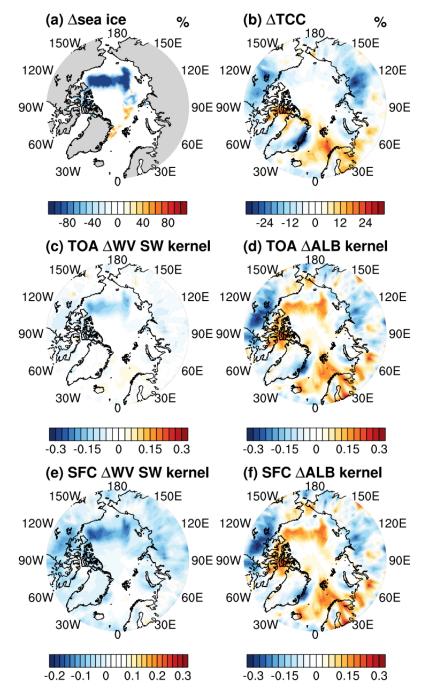


Figure 6. September differences between 2012 and 2013 in (a) sea ice concentration, (b) total cloud cover (TCC), and the differences in (c, e) water vapor SW kernel for TOA and surface fluxes, units: W m⁻² K⁻¹, (d, f) surface albedo kernel for TOA and surface fluxes, units: W m⁻² 1%⁻¹.

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In the sea ice loss case, the reduction of sea ice in the Arctic region (Figure 6a) leads to a
 significant decrease of radiative sensitivity to surface albedo (Figure 6d and f), with the
 maximum difference exceeding 30%. The cloud cover change also contributes to changes in

394 surface albedo kernel values due to the coupling effect between cloud and surface albedo (e.g., 395 see Huang et al., 2021b), which for example is seen in the Siberia and to the west coastline of 396 Europe. The change in sea ice also leads to a significant decrease in the TOA sensitivity and an 397 increase of surface sensitivity to water vapor, respectively (Figure 6c and e), with the maximum 398 changes exceeding 80% for surface. All these results confirm the state-dependency of radiative 399 kernels (e.g., Riihelä et al., 2021). 400 In summary, these quantitatively large interannual differences, as well as their locations, 401 affirm that some discrepancies between the radiative kernels are caused by the difference in 402 atmospheric states and partly explain the inter-dataset kernel differences seen in Figure 3 and 4. 403 Nevertheless, it ought to be noted that the differences are localized and because of that do not 404 cause significant biases in the global mean feedback values (see Section 4). The results above 405 also show that kernel values based on one arbitrary year may be regionally biased. If only one 406 vear's atmospheric profiles are used to generate radiative kernels, we recommend selecting a vear 407 without significant anomalies in atmospheric states, e.g., due to El Nino or severe sea ice loss. 408 409 410 4. Feedback quantification 411 412 In this section, we apply different kernels to quantifying the radiative feedbacks in one 413 quadrupling CO₂ experiment (abrupt4xCO₂) of CMIP6 models. This experiment is selected

because it has been used by a number of studies for forcing and feedback analyses (e.g., Zelinka et al., 2020), which we can compare our results to. The CMIP6 models used in this assessment are listed in Table2. Note that the standard outputs at 19 pressure levels from the models and correspondingly the kernel values, including CAM5 and HadGEM3, provided at the pressure levels are used in this section.

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Table 2. Summary of CMIP6 models used in this study.

Models	Horizontal resolution (lat*lon)	<u>Vertical</u> level <u>s</u>	Reference
CESM2	0.9*1.25	32 levels to 2.26 hPa	Danabasoglu et al. (2020)
CNRM-CM6-1	1.4*1.4	91 levels to 0.01hPa	Voldoire et al. (2019)
EC-Earth3	0.7*0.7	91 levels to 90 km	Döscher et al. (2022)
HadGEM3-GC31-LL	1.25*1.875	85 levels to 85km	Williams et al. (2018)
IPSL-CM6A-LR	1.3*2.5	79 levels to 80km	Boucher et al. (2020)
MPI-ESM1-2-LR	1.875*1.875	47 levels to 0.01hPa	Mauritsen et al. (2019)

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423 **4.1 Analysis procedure**

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To quantify the radiative feedbacks, data from two experiments as documented by Eyring
et al. (2016) and Pincus et al. (2016) are used: abrupt4xCO₂, simulations with an instantaneous
quadrupling of CO₂ concentration of year 1850, piClim-4xCO₂, simulations with SST and sea ice
concentrations fixed at the climatology of pre-industrial control experiment and CO₂

429 concentration quadrupled. In each experiment, a 20-year period at the end of the simulation in

430 each model is used. For example, in the models where the abrupt4xCO₂ simulation is longer than

431 150 years, the simulations from the last 20 years rather than those from years 131 to 150 are used

432 for the calculation. Following the previous studies (e.g., Smith et al., 2020; Zelinka et al., 2020).

radiative feedbacks are diagnosed using the difference of atmospheric variables between the 433

- 434 abrupt4xCO₂ and piClim-4xCO₂ experiments. It is worth noting that the method used in this
- 435 study is slightly different from that in Zelinka et al. (2020), in which piControl simulation was
- 436 used as the climatology baseline and the feedbacks were integrated from the surface to the
- 437 tropopause (as opposed to model top) to remove the rapid adjustment, although the quantitative 438 differences in the global mean feedback values are small.

439 To detail the analysis procedure, firstly, all variables including radiative fluxes and 440 atmospheric variables from CMIP6 models are interpolated to the horizontal and vertical 441 resolution of the kernel itself. Notre that for CAM3, GFDL, CloudSat and CAM5 kernels, they 442 only have 17 pressure levels which are two layers (1hPa and 5hPa) fewer than the CMIP6 443 standard model output. To address this issue, the contribution of the two missing layers is 444 calculated using other kernels (e.g., ERA5) and found to have negligible effect on the global 445 mean feedback value. Hence, when using these three kernels, the contributions from 10hPa 446 above are ignored.

447

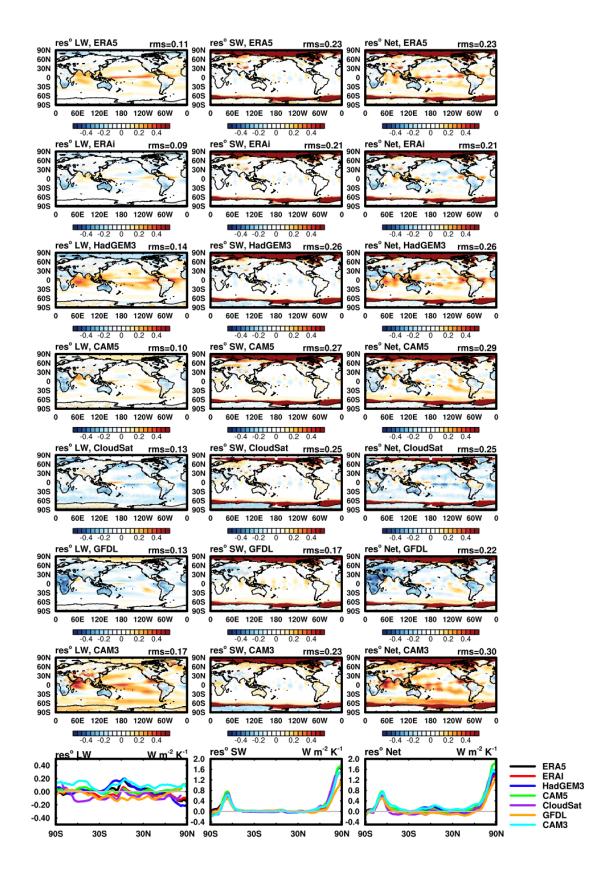
Secondly, the non-cloud radiative feedback of atmospheric variable $X(\Delta R_X)$ is calculated 448 as: 449

 $\Delta R_X = K_X \cdot \Delta X \quad (4)$

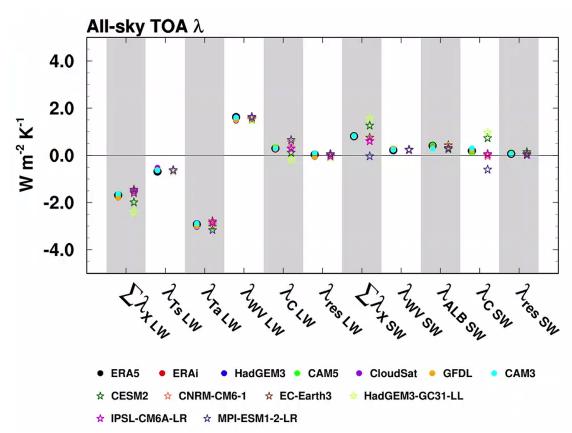
- with units in W m⁻², where K_X is the radiative kernel of variable X and ΔX is the anomaly of X 450 measured by the difference between abrupt4xCO₂ and piClim-4xCO₂, and represents the 451 anomalies of surface temperature (ΔT_s), air temperature (ΔT_a), water vapor (ΔWV) and surface 452 453 <u>albedo (ΔALB)</u>. For the 2D radiative kernels (surface temperature and surface albedo), K_X and 454 ΔX have just single layer values and ΔR_X is simply the product of these two terms. For the 3D radiative kernels (air temperature and water vapor), both K_X and ΔX are vectors of pressure 455 456 levels and ΔR_X is the dot product of K_X and ΔX and is integrated from the TOA to 1000hPa. Note that if K_X is normalized with unit pressure thickness (e.g., W m⁻² K⁻¹ 100hPa⁻¹), the layer 457 thickness must be taken into account when calculating dR_x . See Appendix for further discussion 458 459 on the application of thickness-weighted kernels. 460 Finally, cloud feedbacks are diagnosed using the adjusted cloud-radiative forcing method 461 (Shell et al., 2008). Here we compute the residual term in clear-sky as: $res^o = \sum \Delta R_x^o - \Delta R^o$ 462 (5)463 which represents the unexplained part of radiation budget change, and assuming the all-sky decomposition has the same non-closure residual, the cloud feedback is measured as 464
- 465 $\Delta R_{c} = \Delta R - \sum \Delta R_{x} + res^{o}$ (6)where the superscript o represents clear-sky quantities. $\sum \Delta R_X^o$ and $\sum \Delta R_X$ are the sum of non-466 cloud feedbacks in clear-sky and all-sky, respectively, diagnosed by multiplying the radiative 467
- 468 kernel with the atmospheric responses measured as the difference between abrupt4xCO₂ and
- piClim-4xCO₂ experiments. ΔR^{o} and ΔR are the total radiation change in clear-sky and all-sky, 469 respectively, calculated as the difference in the GCM-simulated radiative fluxes between two 470 471 experiments.
- The feedback parameters, λ_X , in the units of W m⁻² K⁻¹, are then obtained by normalizing 472 the feedback flux changes ΔR_x by the global mean surface temperature change ΔT_s in the 473 474 abrupt4xCO₂ experiment:
- 475 $\lambda_x = \Delta R_x / \Delta T_s$ (7)

478 4.2 TOA feedbacks 479

480 The residual term (res^{o}) measures the unexplained radiation change in the feedback 481 analysis and provides a useful overall indication of the soundness of the feedback quantification. 482 Figure 7 illustrates the residual term for the TOA flux decomposition when different kernels are 483 used to diagnose the multi-model mean feedbacks. In terms of the global mean, all residual terms 484 are of small magnitude, no matter which kernel dataset is used (Figure 8 and Table S1). 485 However, there are some noticeable local residuals, especially for the SW budget, e.g., in the 486 Arctic region and around the Antarctic continent where sea ice changes the most (mid-column in 487 Figure 7). While the non-zero magnitude of the residual is partly due to nonlinearity in the 488 radiation decomposition, e.g., possible coupling between surface albedo and water vapor (Huang 489 et al., 2021b), the spread among the kernel results as evidenced by the line plots of Figure 7 is 490 attributable to the discrepancies in the SW radiative kernels as revealed by the comparisons in 491 Section 3. In the LW, the residual is generally small compared with the total feedback. In 492 summary, the residual terms for the TOA budget are small in terms of the global mean feedback 493 strengths, affirming the validity of the radiative kernels for feedback quantification. Here, we use 494 the spatial root-mean square (RMS) of the residuals to quantify the regional biases, which are 495 shown by the numbers on the right corner of each panel in Figure 7. For LW, results from ERA5, 496 ERAi and CAM5 kernels show relatively smaller regional biases compared to those from 497 HadGM3, CloudSat and CAM3 kernels. For SW, all kernel datasets have similar regional non-498 closures, for example, in the Polar regions (Figure 7 and 8). This is largely caused by the non-499 linearity in albedo feedback and also the coupling effect between water vapor and surface albedo 500 feedbacks (Huang et al., 2021b; Block and Mauritsen, 2013). In summary, these results suggest 501 that for the TOA feedback quantification, the performance of ERA5 kernel is comparable to the 502 other datasets. 503 Figure 8 compares the spreads of feedback values resulted from the differences in kernels 504 and those from the different projections of GCMs. In general, feedbacks from different kernel 505 datasets overlap each other, even for cloud feedbacks, affirming a good consistency between the 506 results computed from different kernel datasets. However, the spread across the GCMs is 507 considerably larger, suggesting the overall feedback uncertainty is dominated by inter-model 508 spread rather than the kernel uncertainty. The values of the feedbacks from each model and 509 kernel datasets are shown in Table S1 and S2 for readers who are interested. These results are 510 consistent with other published results. For example, compared with the results of Zelinka et al. 511 (2020) based on the ERAi kernel, the kernel-diagnosed overall feedback parameter in the two 512 results is -0.87 W m⁻² K⁻¹ and -0.85 W m⁻² K⁻¹ for the CNRM-CM6-1 model and -0.81 W m⁻² K⁻¹ 513 and -0.84 W m⁻² K⁻¹ for the HadGEM3-GC3-LL model. 514 In summary, in terms of TOA feedback values, the inter-kernel differences lead to small 515 uncertainty in the analyzed non-cloud feedbacks; the kernel-induced uncertainty in cloud 516 feedback is relatively larger (Table S2), with the inter-kernel spread in cloud LW feedback 517 almost equally from the spread in surface and air temperature feedback and water vapor LW 518 feedback, and the inter-kernel spread in cloud SW feedback more from the spread in surface 519 albedo feedback than from water vapor SW feedback (not shown); despite this, this uncertainty 520 is considerably less than the inter-GCM cloud feedback spread.



524 Figure 7. The residuals (res^o) in the multi-model mean TOA feedback decomposition 525 when different kernels are used, (left column) LW, (mid-column) SW, (right-column) Net, the sum of LW and SW. The three line-plots in the bottom row are the zonal mean residuals. 526 527 Numbers on the right corner in each panel are the spatial root-mean-square values.



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Figure 8. Global mean TOA feedback parameters in all-sky diagnosed by the kernels list in Table 531 1 across CMIP6 models. Dot marks represent multi-model mean values computed from different 532 kernel datasets. Pentagrams represent the multi-kernel mean results computed from different 533 GCMs.

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536 4.3 Surface feedbacks

538 Next, we examine how the inter-kernel differences lead to uncertainty in the analyzed 539 surface feedbacks.

540 Figure 9 shows the residual distribution. We find that when the ERA5 and ERAi kernels 541 are used for the feedback analysis, the non-closure residual in the surface budget is comparable 542 in magnitude to the TOA analysis. This suggests that the surface kernels afford a valid tool for 543 the surface feedback analysis. However, some prominent biases are noticed for other kernel 544 datasets. For example, the HadGEM3 kernels, show especially an underestimation in air 545 temperature feedback, likely due to a biased sensitivity of the bottom atmospheric layer (see 546 Appendix for more discussions). The <u>sum of global mean surface and air temperature feedback</u>

547 parameter measured by the HadGEM3 kernel is around -3.70 W m⁻² K⁻¹ (Table S4, compared to

548 around -1.0 W m⁻² K⁻¹ measured by the other kernels), and the non-closure residual is as large as 549 -3.0 W m⁻² K⁻¹ (Table S2, compared to 0.1 W m⁻² K⁻¹ in the others). For this reason, the result 550 from HadGEM3 kernel is excluded for the multi-kernel statistics in Figure 10. Table S3 and S4. 551 but listed in a separate row for comparison. From either the spatial distribution of residual term 552 or the spatial RMS residuals, the ERA5 kernel and ERAi kernel show a superior performance 553 than other datasets. The use of ERA5 kernels may be advantageous for diagnosing the surface 554 radiation budget, considering that ERA5 data is a newer version reanalysis dataset from ECMWF 555 compared with ERAi and its data quality has been widely validated. 556 Figure 10 compares the inter-model and inter-kernel spreads for the surface feedbacks. Unlike the results for TOA, the inter-kernel spread can be as large as the inter-model spread, for 557 558 example, in LW surface temperature feedback, air temperature feedback and water vapor 559 feedback. The sum of air temperature and surface temperature feedbacks shows better consistency compared with the respective components, except for HadGEM3 kernel, which is 560 561 due to the reason discussed in the Appendix - possibly wrong quantification of surface 562 temperature effect. In SW, the multi-kernel results are close to each other, showing smaller inter-563 kernel spreads than the inter-model spreads. 564 In summary, we find the surface feedback decomposition can achieve similar level of 565 radiation closure to the TOA analysis when using ERA5 kernels, affirming the validity of kernels 566 for diagnosing the surface radiative feedback. However, the results qualitatively vary depending 567 on which kernel dataset is used, indicating errors in the computation of some kernels. 568

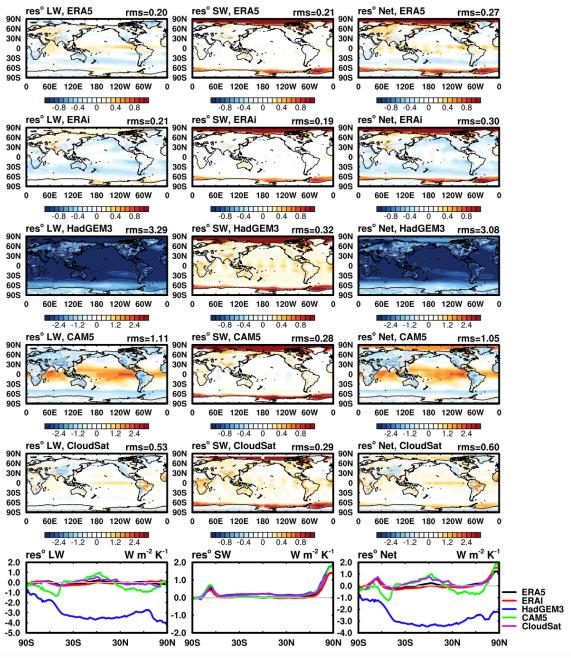
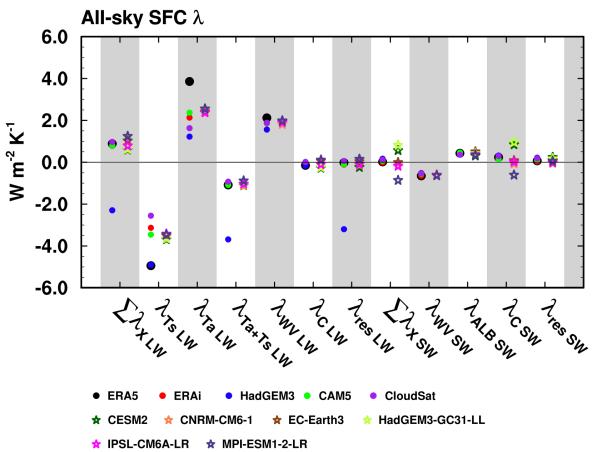




Figure 9. Similar to Figure 8, but for the surface feedback analysis.



573 574 Figure 10. Similar to Figure $\underline{8}$, but for the surface feedback parameter.

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5. Data availability

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579 The datasets contain the multi-year averaged monthly mean TOA and surface kernel for
580 surface temperature, air temperature, surface albedo and water vapor (LW and SW) and are
581 available at: <u>http://dx.doi.org/10.17632/vmg3s67568.2</u> (Huang and Huang, 202<u>3</u>).

583 6. Conclusions and discussions

584

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585 In this paper, we present a newly generated set of ERA5-based radiative kernels of 586 surface and air temperatures, water vapor and surface albedo, for both TOA and surface radiation 587 fluxes. We also compare them with other published kernels, <u>including</u> the kernel values <u>and</u> the 588 <u>kernel-diagnosed</u> radiative feedbacks for both the TOA and surface radiation budgets.

589 For the TOA kernels, the results here affirm general consistency among the different 590 kernel datasets, and the discrepancies are generally within 10% in terms of vertically integrated 591 or globally averaged radiative sensitivity, although some relatively larger regional biases are

592 noticed, including those in the surface temperature kernel in the tropics (Figure 3a), those in the

593 surface albedo kernel in the Arctic (Figure 3q) and those in the water vapor shortwave kernel in

594 the Antarctica (Figure 3m), which is partly due to the dependence of radiative sensitivity on 595 background climate states.

596 For the surface kernels, more prominent inter-kernel <u>differences</u> are found. For example, 597 the <u>differences</u> in the water vapor shortwave kernel in the Antarctic (Figure 3o) and in the 598 surface albedo kernel in the Arctic (Figure 3s) can reach 30%. <u>Some kernels have considerably</u> 599 biased air temperature sensitivity values in the bottom atmospheric layers, which is likely due to 600 improper treatment in the perturbation experiments used for kernel computation. The <u>differences</u> 601 in both TOA and surface kernels discovered here affirm the importance of validating the

602 radiative sensitivity as noted by Huang and Wang (2019) and Pincus et al. (2020).

<u>The investigation of interannual variability in ERA5 kernels affirm the dependence of</u>
 <u>radiative sensitivity on atmospheric state and the further comparison between ERAi and ERA5</u>
 <u>kernel (Figure 5) reveals the effects of clouds on the kernel values, which explains the</u>
 discrepancies of multi-kernel datasets (Figure 3).

- 607 Applying the different kernels to quantifying the TOA and surface radiative feedbacks, 608 we find that for TOA feedback quantification, the ERA5 kernels are as accurate as other kernel
- 609 <u>datasets, while for surface feedback, ERA5 and ERAi kernels show superior accuracy compared</u> 610 with other datasets. Considering the strengths of the ERA5 dataset in representing the
- 611 atmospheric states, we recommend the use of ERA5 kernels.

In addition, we compare the feedback differences caused by using different kernels and
also the inter-GCM spread of the feedback values (when measured by the same kernel). We find
the kernel <u>difference</u> is not a major cause of the inter-GCM <u>TOA</u> feedback spread (<u>Figure 7 and</u>
8). This finding is <u>consistent</u> with the previous assessments (e.g., Soden et al., 2008; Jonko et al.,
2012; Vial et al. 2013).

617 Radiation closure tests show that the unexplained residuals are generally within 10% for 618 both TOA and surface analyses in terms of the global mean feedback, affirming the validity of 619 the kernels for feedback quantification for both budgets. This suggests that the large non-closure 620 residuals reported in some previous studies (e.g., Vargas Zeppetello, et al., 2019) are likely due 621 to kernel inaccuracy rather than the limitation of the kernel method. However, there are more 622 significant local non-closures, for example, in the shortwave in the Arctic region and around the Antarctic continent, which is contributed, but cannot be fully explained, by the kernel 623 624 uncertainty. This points to the accuracy limit of the kernel (linear) method and calls for more 625 advanced methods, such as the neural network method (Zhu et al., 2019), for local feedback

- 626 analysis.
- 627
- 628

629 Author contributions

630

631 HH produced the ERA5 radiative kernel and provided calculations of the inter-kernel

- 632 comparison and feedback analysis. Both HH and YH led the writing of the manuscript.
- 633

634 **Competing interests**

- 635 The authors declare that they have no conflict of interest.
- 636637 Acknowledgements
- 638

639 We thank Mark Zelinka, Ryan Kramer and one anonymous reviewer for their helpful

640 reviews. We acknowledge the grants from the Natural Sciences and Engineering Research

641 Council of Canada (RGPIN-2019-04511) and the Fonds de Recherché Nature et Technologies of

642 Quebec (2021-PR-283823) that supported this research. H. Huang thanks Yonggang Liu, Jun

643 Yang and Qiang Wei for hosting her visit at Peking University and assisting her with computing

- 644 resources.
- 645 646

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793 Appendix 794 795 The ERA5 kernels are computed following Eq. (1) and the approach outlined in Section 796 2.2. 797 798 1. Surface variable kernels 799 800 To execute the partial radiative perturbation computations, the perturbations are prescribed 801 as the following: for the 2D feedback variables, the surface temperature is increased by 1 K and 802 the albedo is increased by 0.01 at each location. Hence, the units of the two kernels, K_{Ts} and K_{ALB} are W m⁻² K⁻¹ and W m⁻² %⁻¹, respectively. When applying them to feedback 803 quantification, their feedbacks are quantified as 804 805 $\Delta R_{Ts} = K_{Ts} \cdot \Delta T_S$ (A1) $\Delta R_{Alb} = K_{ALB} \cdot \Delta Alb \quad (A2)$ 806 where ΔT_s should be measured in the units of K and ΔAlb in <u>absolute values</u>, i.e., the multiply of 807 808 1%. 809 810 2. Water vapor kernel 811 812 For the 3D feedback variables, the perturbations are applied to each of the 37 pressure 813 layers (from 1hPa to 1000hPa) and one layer at a time. For the water vapor kernel, a 10% 814 incremental perturbation of the water vapor concentration is used. To adapt to the convention used in the majority of the existing kernels, we convert the units of the kernels to represent the 815 radiative flux change corresponding to an increase of water vapor concentration that conserves 816 817 the relative humidity of the layer under a 1-K increase in air temperature, i.e., converting the units from W/(m² $\Delta q_0^{+10\%}$ 100hPa) to W/(m² Δq_0^{+1K} 100hPa): $K_q^{+10\%} = \frac{\Delta R_0}{\Delta q_0^{+10\%}}$ (A3) 818 819 $K_q^{+1K} = \frac{\Delta R_0}{\Delta q_0^{+1K}} = K_q^{+10\%} \cdot \frac{\Delta q_0^{+10\%}}{\Delta q_0^{+1K}} = K_q^{+10\%} \cdot \frac{\Delta q_0^{+10\%}}{q_0} \cdot \frac{e_s(T_0)}{e_s(T_0 + 1K) - e_s(T_0)}$ (A4) 820 Where q_0 is the unperturbed water vapor concentration, in units of kg kg⁻¹. $\Delta q_0^{+10\%}$ is a 10% 821 increment in water vapor concentration. $e_s(T)$ is the saturated water vapor pressure under 822 temperature T, and can be measured by empirical formulas; hence, Δq_0^{+1K} can be measured as 823 $q_0\left[\frac{e_s(T_0+1K)}{e_s(T_0)}-1\right]$. Accordingly, when the water vapor kernel is used for water vapor feedback 824 quantification, the feedback is measured as: 825 $\Delta R_q = K_q^{+1K} \cdot \Delta q^{+1K} = K_q^{+1K} \cdot \frac{\Delta q}{\Delta q_0^{+1K}} = K_q^{+1K} \cdot \frac{\Delta q}{q_0} \cdot \frac{e_s(T_0)}{e_s(T_0 + 1K) - e_s(T_0)}$ 826 (A5) where $\Delta q = q - q_0$ measures the change in water vapor concentration and is normalized by 827 Δq_0^{+1K} to give the factor that is multipliable with the K_q^{+1K} kernel value. If using the Clapeyron-828 Clausius relation, the above expression can be further approximated as 829 $\Delta R_q = K_q^{+1K} \cdot \frac{\Delta q}{q_0} \cdot \frac{e_s}{(de_s/dT) \cdot 1K} = K_q^{+1K} \cdot \frac{\Delta q}{q_0} \cdot \frac{R_v}{L_v} \cdot \frac{T_0^2}{1K}$ 830 (A6) where R_v and L_v are the gas constant and specific latent heat of water vapor, respectively. Note 831 that when the kernels are used, T_0 and q_0 typically take their values from the base climate 832 appropriate to the application, e.g., the unperturbed climate of a GCM experiment, not 833 834 necessarily the dataset used for kernel computation.

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836 3. Air temperature kernel

For the air temperature kernel, to be consistent with the "inhomogeneous path treatment"
that accounts for the vertically non-uniform temperature distribution within each discrete
atmospheric layer (Mlawer et al., 1997), perturbations are added not only to the layer-mean
temperature but also the temperature at the exiting boundary of radiative fluxes of interest (i.e.,
the upper boundary of each layer for the TOA flux and the lower boundary for the surface flux),
to appropriately represent the physical temperature perturbation in each layer.

844 A meaningful test to affirm the validity of the air temperature kernel is a vertical sum test, 845 i.e., a linear additivity test to verify the vertical integration of the kernel values reproduce the flux change, either at TOA or surface, in response to a whole-column air temperature increase of 846 1K. Figure A1 shows that the ERA5 kernel well passes this test. However, as shown by Figure 9, 847 848 some kernels (e.g., HadGEM3 kernel) show much weaker radiative response at surface, possibly 849 due to improper treatment of the air temperature perturbation in the kernel computation, which 850 may lead to an underestimated air temperature feedback and large biases in the surface feedback 851 analysis.



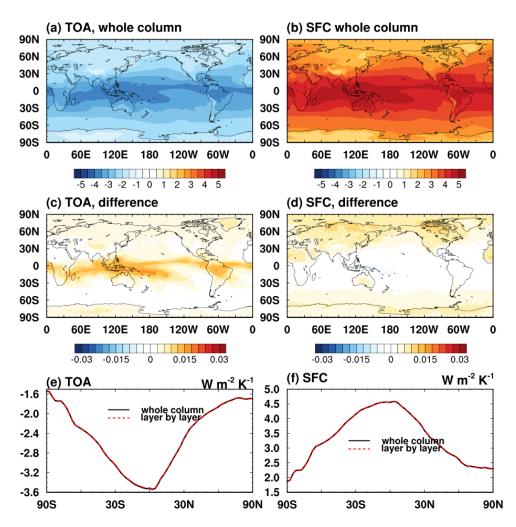


Figure A1. Monthly mean TOA and surface radiation flux change in response to a +1K air temperature perturbation throughout the vertical column: (a, b) computed by a radiation model, RRTMG; (c, d) difference of vertical sum of air temperature kernels compared to truth in (a, b); (e, f) comparison of the zonal mean.

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859 Another challenge in the computation of air temperature kernel for surface flux is that the 860 surface in radiative transfer models is also the lower boundary of the lowermost atmospheric layer. If the effects of the surface temperature perturbation on the emission of the surface and 861 that of the lowermost atmospheric layer are not distinguished, this may lead to improper 862 interpretation and use of the surface temperature kernel. In our ERA5 kernel, the two effects are 863 864 considered separately: according to radiative transfer theory, an increase in surface skin 865 temperature only affects the surface upward emission; an increase in air temperature only affects the downward radiation. In some other kernels such as CAM5, these effects are not 866 867 distinguished, so that the kernel value represents the net effect, i.e., change in the sum of both downward and upward. As a result, in Table S4, we can only report the sum of surface and air 868 869 temperature feedbacks. Figure A2 shows the comparison of vertically integrated air temperature 870 kernels and the sum of surface and air temperature kernels between ERA5, CAM5, HadGEM3 871 and CloudSat. Although the strength of vertically integrated air temperature kernel for CAM5 is 872 much weaker than that for ERA5 (Figure A2a and b), the sum of surface and air temperature 873 kernel between these two datasets are in good agreement (Figure A2c and d), which warns that 874 the seemingly right temperature feedback quantified by some kernels might come from the misattribution of surface temperature contributions. Another noticeable feature in Figure A2 is 875 that the HadGEM3 kernel shows an underestimation in vertical integration of air temperature 876 877 kernel and an overestimation in the sum of surface and air temperature kernel, likely due to 878 mistreatment of the bottom layer, and this accounts for the biased surface feedback analysis as

879 shown in Figure 9. <u>Similar issues were noticed in Kramer et al. (2019a).</u>

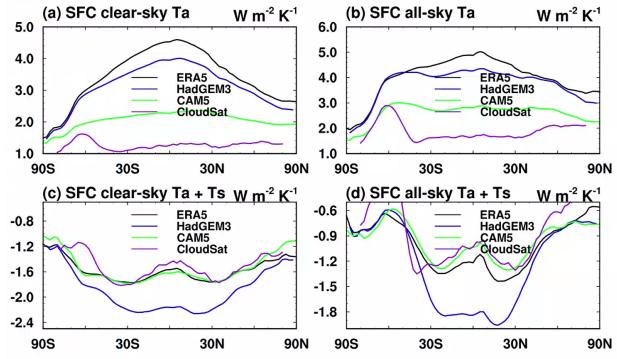


Figure A2. Comparison of annual mean<u>surface</u> kernels for ERA5, CAM5<u>, CloudSat</u> and HadGEM3 for (a, b) the vertically integrated air temperature kernel values, and (c, d) sum of surface and air temperature kernels.

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- 886 4. Time averaging887

As described in Section 2.2, all the kernels provided for feedback analysis are averaged from instantaneous kernel values over each calendar month and, in the ERA5 kernel, over multiple years. This is to ensure proper sampling of radiative sensitivity values under different atmospheric states, so that the kernels are representative of mean radiative sensitivity and thus can be readily multiplied with monthly mean climate responses (ΔX) to evaluate climate feedbacks.

894 If the kernels are computed for fixed pressure levels, and if the pressure of any of these 895 levels of an instantaneous atmospheric profile is higher than the surface pressure (i.e., the level is below the surface) at a time instance, this potentially creates inconsistency in the averaging 896 897 procedure. To address this concern, we set the kernel value to zero (as opposed to missing value) 898 before averaging. This is to ensure that when multiplied with the monthly mean climate response 899 (ΔX) , the contribution of a pressure layer (e.g., that centered at 1000 hPa) is effectively counted only for the fraction of time the layer exists (when surface pressure is higher than 1000 hPa). 900 901 Otherwise, the feedback quantification needs to be further weighted with fraction of time (f)902 when the pressure layer exists. For example, if the surface pressure is larger than 1000hPa only 903 for half of time in a month (f=0.5), the radiation flux anomaly contributed by the layer centered 904 at 1000 hPa is:

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- 906 907

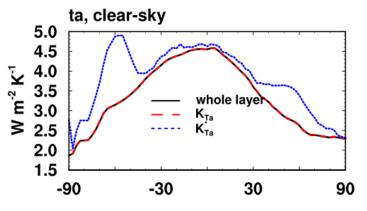
$$\Delta R_{T_{1000hPa}} = K^*_{T_{1000hPa}} \cdot \Delta T_{1000hPa} \cdot f \tag{A7}$$

908 Here, $K_{T_{1000hPa}}^*$ represents the kernel value averaged from the time instances when the layer 909 exists. Our averaging scheme is essentially to provide a kernel $K_{T_{1000hPa}} = K_{T_{1000hPa}}^* \cdot f$, so that 910 it can be simply multiplied with $\Delta T_{1000hPa}$ to obtain the same result.

911 Figure A3 illustrates the differences between $K_{T_a}^*$ and K_{T_a} , in terms of their vertically 912 integrated value. Such difference is pronounced over the Southern Oceans (around 60S), where

913 the surface pressure value varies considerably. This likely explains why Figure 3h shows

noticeable differences in the air temperature kernel in this region.



- 916 Figure A3. Zonal mean monthly mean air temperature kernels for surface flux from 917 ERA5 in clear-sky. Black line is the result from the whole column perturbation computation by 918 RRTMG, providing a "truth" for comparison. Red dashed line is the kernel weighted with fraction of time (K_{T_a}) and blue dotted line represents results without weights $(K_{T_a}^*)$. 919 920 921 5. Layer-specified and layer thickness-normalized radiative kernels 922 923 We generate two versions of vertically resolved air temperature kernel, water vapor LW 924 and SW kernel, one with values corresponding to specified vertical layers, i.e., in the units of W m⁻² K⁻¹, and another with unit-layer thickness (e.g., as shown in Figure 2 and 4), i.e., in W m⁻² K⁻ 925 926 ¹ 100hPa⁻¹. The latter one properly portrays the vertical distribution of radiative sensitivity to 927 perturbations in unit thickness layers, while the former one may be more convenient to use in 928 feedback quantifications. For TOA budget analyses, these two versions of kernels lead to little 929 difference in practice due to limited contributions from the bottom atmospheric layer. However, 930 for surface budget analyses, we recommend using the layer-specified kernels, as the surface 931 kernels typically show strongest sensitivity to the perturbations in the bottom layers, which can 932 be best accounted for in the non-normalized kernels. Otherwise, the difference of surface 933 pressure between ERA5 and GCMs needs to be carefully treated to avoid errors, for example, 934 caused by missing the radiative contribution from the bottom layer of the atmosphere. 935 936 937 938
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