

ML-TOMCAT: Machine-Learning-Based Satellite-Corrected Global Stratospheric Ozone Profile ~~Dataset~~Data Set from a Chemical Transport Model

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Abstract.

High quality stratospheric ozone profile ~~datasets~~data sets are a key requirement for accurate quantification and attribution of long-term ozone changes. Satellite instruments ~~obtain~~provide stratospheric ozone profile measurements over typical mission durations of 5-15 years. Various methodologies have then been applied to merge and homogenise the different satellite data in order to create longer term observation-based ozone profile ~~datasets~~data sets with minimal data gaps. However, individual satellite instruments use different measurement methods, sampling patterns and retrieval algorithms which complicate the merging of these different ~~datasets~~data sets. In contrast, atmospheric chemical models can produce chemically consistent long-term ozone simulations based on specified changes in external forcings, but they are subject to the deficiencies associated with incomplete understanding of complex atmospheric processes and uncertain photochemical parameters.

Here, we use chemically self-consistent output from the TOMCAT 3-D chemical transport model (CTM) and a Random-Forest (RF) ensemble learning method to create a merged 42-year (1979-2020) stratospheric ozone profile ~~dataset~~data set (ML-TOMCAT V1.0). The underlying CTM simulation was forced by meteorological reanalyses, specified trends in long-lived source ~~gas~~gases, solar flux and aerosol variations. The RF is trained using the Stratospheric Water and OzOne Satellite Homogenized (SWOOSH) ~~dataset~~data set over the time periods of the Microwave Limb Sounder (MLS) from the Upper Atmosphere Research Satellite (UARS) (1991-1998) and Aura (2005-2016) missions. We find that ML-TOMCAT shows excellent agreement with available independent satellite-based ~~datasets~~data sets which use pressure as ~~the~~-vertical coordinate (e.g. GOZCARDS, SWOOSH for non-MLS periods) but weaker agreement with the ~~datasets which are height-based~~data sets which are altitude-based (e.g. SAGE-CCI-OMPS, SCIAMACHY-OMPS). We find that at almost all stratospheric levels ML-TOMCAT ozone concentrations are well within uncertainties ~~in the observational datasets~~of the observational data sets. The ML-TOMCAT ~~dataset is thus~~(V1.0) data set is ideally suited for the evaluation of chemical model ozone profiles from the tropopause to 0.1 hPa. ~~ML-TOMCAT data and~~ is freely available via <https://zenodo.org/record/4997959#.YNzleUIKiUk> (Dhomse et al., 2021).

1 Introduction

With the successful implementation of the Montreal Protocol, various observations confirm reductions in the concentrations of ~~ozone-depleting~~ halogenated ozone-depleting substances (ODSs) in the atmosphere (WMO, 2014, 2018). Satellite data records also confirm a peak in upper stratospheric HCl (the main chlorine reservoir) around 1997, followed by a steady decline (Anderson et al., 2000; Froidevaux et al., 2006a; Hossaini et al., 2019). Hence, attention has turned towards the detection and attribution of ozone recovery (~~e.g. Dhomse et al., 2006; Solomon et al., 2016; Chipperfield et al., 2017; Dhomse et al., 2018~~) (e.g. Dhomse et al., 2018). However, the accurate quantification of ozone changes is challenging because of the quality of long-term ozone profile ~~datasets~~ data sets, where measurement errors are of similar ~~magnitude or larger or larger magnitude~~ than the long-term ozone trends. In addition, complex coupling between various physical and chemical processes controlling stratospheric ozone concentrations cause large short-term ozone changes. Complications also arise because there are some non-linear changes in stratospheric dynamics as well as chemical constituents. For example, between 2018 and 2021, some of the largest and smallest ozone losses of the recent decades were recorded in both the Arctic and Antarctic polar stratospheres (e.g. Wargan et al., 2020; Wohltmann et al., 2020; Bogner et al., 2021; Weber et al., 2021). Some observational data suggest that there has been a continuous decline in lower stratospheric ozone (Ball et al., 2018, 2020), which could be attributed to changes in stratospheric dynamics (e.g. Chipperfield et al., 2018; Wargan et al., 2018; Orbe et al., 2020; Abalos and de la Cámara, 2020). Atmospheric concentrations of ODSs such as CFC-11 are decreasing at uneven rates (Montzka et al., 2018, 2021) which could induce variability in ozone trends. Additionally, significant positive trends have been detected in very short-lived substances (VSLs) containing chlorine and bromine that are not controlled by the Montreal Protocol (e.g. Hossaini et al., 2015, 2019).

As there ~~is~~ are no long-term ozone profile data from a single satellite instrument, various attempts have been made to merge such data from different instruments. However, individual satellite instruments have different temporal and spatial resolution depending on the measurement techniques and retrieval algorithms (~~e.g. Damadeo et al., 2018~~) (e.g. Sofieva et al., 2014; Damadeo et al., 2014). For example, solar occultation instruments (e.g. Stratospheric Aerosol and Gas Experiment (SAGE, McCormick et al., 1989), Halogen Occultation Experiment (HALOE, Russell III et al., 1993)) provide high quality measurements but are constrained by limited spatial coverage. Limb-scanning instruments such as the Microwave Limb Sounder (MLS, Froidevaux et al., 2006b), Scanning Imaging Absorption Spectrometer for Atmospheric Cartography (SCIAMACHY, Bovensmann et al., 1999), Optical Spectrograph and InfraRed Imager System (OSIRIS, Murtagh et al., 2002) provide better spatial coverage but have coarser vertical resolution. A key constraining factor is that only few satellite ~~datasets~~ data sets cover enough overlapping years to remove inter-instrument biases with minimal uncertainty.

Hence, Randel and Wu (2007) adopted a novel approach to create a gap-free stratospheric ozone profile data for the 1979–2005 time period. They used SAGE (I and II) satellite profile measurements and polar ozonesondes, together with a seasonally varying ozone climatology from Paul et al. (1998) to fill the gaps, to generate multi-variate ~~regression-based~~ regression-based gap-free ozone anomalies. Later, Cionni et al. (2011) used a similar methodology along with climate model simulations to extend the time series backwards to 1850. The Cionni et al. (2011) data ~~was~~ were recommended for the historical CMIP5 simulations ~~in order to enforce time-dependent ozone variations, for the~~ for the climate models that did not

include stratospheric chemistry, in order to enforce time-dependent ozone variations. Hassler et al. (2008) used a different methodology to create a satellite-based long-term ozone profile datasetdata set. Along with SAGE I and II measurements, they used HALOE and POAM (Polar Ozone and Aerosol Measurement) II and III satellite measurements, as well as ozonesonde data from 130 stations, to create a collection of binary data files; also known as the "Binary DataBase of Profiles" (BDBP) version 1.0. Bodeker et al. (2013) ~~;~~updated the BDBP datasetdata set to construct "Bodeker Scientific" or "BS" data. They updated BDBP data by including measurements from the Limb Infrared Monitor of the Stratosphere (LIMS), the Improved Limb Array Spectrometer (ILAS), and ILAS II. They used a multivariate regression model to create different versions of the ozone profile datasetdata set ranging from the surface to 70 km for the 1979-2008 time period. Hassler et al. (2018a) revised ~~the Bodeker et al. (2013) dataset~~ and extended (1979–2016) the "BS" data set by using the TOMCAT chemical transport model (CTM) ozone profiles as a transfer function to capture ozone variability for the period without satellite observations.

Another widely used merged data set is the Global OZone Chemistry And Related trace gas Data records for the Stratosphere (GOZCARDS, Froidevaux et al., 2015). These are monthly mean zonal-meanzonally averaged time series constructed using ozone ~~profiles~~profile measurements from several NASA satellite instruments and the Atmospheric Chemistry Experiment - Fourier Transform Spectrometer (ACE-FTS, Bernath et al., 2005). Merging is done primarily by removing average biases between SAGE II and individual data records for overlap periods (Froidevaux et al., 2015). The GOZCARDS data files contain mixing ratios on a pressure–latitude grid (300 hPa to 0.1 hPa), updated later to GOZCARDS v2.2 (Froidevaux et al., 2019).

Davis et al. (2016) adopted a slightly different approach to construct the Stratospheric Water and Ozone Satellite Homogenized (SWOOSH) data set. SWOOSH merges stratospheric ozone profile data from solar occultation instruments (SAGE-II/III, HALOE, ACE-FTS) as well as limb-scanning instruments (UARS-MLS, and Aura-MLS). The measurements are homogenized by applying corrections that are calculated from data taken during time periods of instrument overlap. The primary SWOOSH data product consists of monthly mean zonal-mean values on a pressure grid at 2.5, 5 and 10 degree resolution. One of the major characteristics of SWOOSH data is that when merging greater weight is given to the instruments that sample more frequently (e.g. Aura-MLS). Filled and unfilled versions of the datasetdata set exist on both geographical and equivalent latitude coordinates.

Several ~~other~~additional attempts have been made to merge satellite time series from limb and occultation instruments. ~~In this study we consider~~ For example, the SAGE–CCI–OMPS datasetdata set, described in Sofieva et al. (2017) ~~;~~which includes SAGE II time series and several limb datasetsdata sets. The OMPS-LP datasetdata set used is produced at the University of Saskatoon (Zawada et al., 2018). ~~The CCI data sets were firstly~~ First, they screened and homogenized CCI data sets in the HARMOZ format ~~and then merged before merging them~~ in terms of ozone anomalies ~~with the SAGE II and OMPS-LP observations~~.

~~Arosio et al. (2019) also~~ Recently, Arosio et al. (2019) created a merged SCIAMACHY-OMPS limb merged datasetdata set (SCIA-OMPS), which combines these two time series produced at the University of Bremen. They ~~use~~used MLS data series as a transfer function to merge SCIAMACHY with OMPS-LP as these instruments share only two months of overlap, but MLS was not included in the merged datasetdata set. This time series is monthly averaged, covers the period 2002-present and is longitudinally resolved, with a 5°–5° latitude × 20°–20° longitude grid. Due to the similarities in the measurement

geometries and techniques, and in the retrieval approaches, ~~a plain de-biasing approach was implemented~~they implemented a plain de-biasing approach for the merging, directly obtaining a long-term ozone time series in appropriate units.

Another widely accepted approach is using data assimilation techniques to create observation-based data (e.g. Feng et al., 2008; Skachko

95 However, only a few instruments such as MLS provide relatively long-term ozone profile measurements. For the pre-MLS time period, very few observations are available that can provide good constraint on the assimilation system. Also, the forward model is generally forced with available (re)analysis dynamical fields so reanalysis data sets are also prone to the inhomogenities in the forcing fields along with any discrepancies in chemical scheme.

In this paper we present a new data-model method for producing a long-term ~~dataset~~data set of stratospheric ozone. We
100 use ozone profile output from a CTM to create a machine-learning-based satellite-corrected long-term chemically (and dynamically) consistent ozone profile ~~dataset~~data set (hereafter, ML-TOMCAT) for the ~~1979–2020~~1979 – 2020 time period. The CTM setup is described in Section 2, followed by our methodology in Section 3. Comparisons of ML-TOMCAT with some of the other available merged ozone profile ~~datasets~~data sets are presented in Section 4, with a summary of our key results in Section 5.

105 2 Model Setup

We use chemically consistent monthly mean zonal mean ozone profiles from the TOMCAT ~~chemical-transport-model (CTM)~~
110 ~~)~~CTM as the basis ~~dataset~~data set. TOMCAT is an off-line three-dimensional (3D) CTM that includes a comprehensive stratospheric chemistry scheme (Chipperfield, 2006). For the present study, the CTM setup is similar to the control simulations used in our recent studies such as Feng et al. (2021); Bogner et al. (2021) and Weber et al. (2021). Briefly, ~~the model~~
TOMCAT is forced with meteorological fields from ERA-5 reanalyses (Hersbach et al., 2020), starting from 1979. Simulations are performed at a 2.8×2.8 degree horizontal resolution with 32 hybrid sigma-pressure levels extending from the surface to about 60 km. For major ODSs and GHGs the model uses time-dependent observed global mean surface mixing ratios (Carpenter et al., 2018) that are treated as well-mixed throughout the troposphere. The model also includes the effects of solar flux variability and heterogeneous chemistry on volcanically enhanced stratospheric aerosol as described in
115 ~~Dhomse et al. (2015, 2016)~~Dhomse et al. (2015, 2016). Solar irradiance data are from the NRL2 (Coddington et al., 2016) empirical model and the ~~sulphate~~sulfate aerosol surface area density (SAD) from Luo (2016). ~~The model~~TOMCAT also includes chlorine and bromine contributions from VSLS as described in Hossaini et al. (2019). ~~The model~~A passive ozone tracer (no chemical ozone loss), generally used to diagnose chemical ozone loss, is initialised every six months from the chemical ozone tracer (1st June and 1st December). TOMCAT has been regularly used to study long-term changes in stratospheric trace
120 gases, showing good agreement with various ground-based and satellite ~~datasets (e.g. Mahieu et al., 2014; Chipperfield et al., 2015; Wales et al., 2018; Harrison et al., 2021; Prignon et al., 2021)~~sets (e.g. Mahieu et al., 2014; Chipperfield et al., 2015; Wales et al., 2018; Harrison et al., 2021; Prignon et al., 2021).

3 Methodology

We use the Random Forest (RF) regression analysis to generate a long-term chemically consistent ~~dataset~~data set. The RF is a supervised machine learning (ML) algorithm that uses an ensemble of decision trees (e.g. Breiman, 2001; Svetnik et al., 2003). ~~Decision trees use a binary recursive classifying algorithm by splitting observations into two homologous groups. The recursive nature of the algorithm means splitting could be repeated until only two observations are left in the final split.~~ A decision tree can be considered as a flow chart used in computer programming (a tree-shaped schematic) that is generally used to show a statistical probability or path of action. A single decision tree in a RF can be considered as a random tree in a forest of decision trees. Each decision tree consists of three components: decision nodes, leaf nodes, and a root node. The root node and decision nodes of the decision tree represent the explanatory variables. The leaf node represents the final output. The explanatory variables used in our analysis are explained at end of this section.

A decision tree algorithm divides the data set into branches (using true and false criterion), which further segregate into other branches until a leaf node (or result node) is reached. Multiple trees are constructed by randomly sampling data points multiple times (e.g. bootstrap method). Hence, an individual tree can be considered as unique tree (hence unique output). RF uses a bagging technique, that means the RF model consists of many individual ~~trees or learners~~decision trees and aggregated predictions are used for the final prognosis. A distinct advantage of RF regression is that it can be relatively accurate and very easy to set up. RF can also behave like a non-linear regression method. As RF adds randomness to the ~~model~~decision procedure, instead of relying on the most important ~~feature~~explanatory variables, it searches for the best ~~feature~~variable among random subsets. This ensures that the final output does not rely heavily on a single ~~learner~~explanatory variable, thereby avoiding ~~overfitting over-fitting~~ (e.g. Kotsiantis, 2013). We use ~~Random Forest (RF)~~ RF Regression from the Python package ~~sklearn~~ Sci-kit learn (Pedregosa et al., 2011) with two options: *random_state=0*, and *bootstrap=True*.

Initially, TOMCAT zonal mean ozone profiles are linearly interpolated in log-pressure space on to 43 equidistant (12 per decade) pressure levels (1000–0.1 hPa, MLS pressure levels), followed by spatial interpolation ~~on to~~onto 72 SWOOSH latitude bins at 2.5° resolution. SWOOSH data ~~is~~are obtained via <https://csl.noaa.gov/groups/csl8/swoosh/>. Then, we calculate the ozone difference (dO_3) between SWOOSH and model ozone profiles for the 1991-1998 and 2005-2016 time periods (total 20 years). For the calculation of dO_3 values, we use the gap-filled SWOOSH data product. SWOOSH data ranges from 316 to 1 hPa (31 pressure levels), hence dO_3 for pressure levels below 316 hPa are linearly ~~interpolated~~extrapolated by approximating dO_3 at 1000 hPa to be about 0.01 ppm. Similarly, for levels above 1 hPa, we use linear interpolation assuming that dO_3 at 0.1 hPa is about -0.1 ppm, based on bias seen with respect to Aura MLS measurements ([Livesey et al., 2020](#)).

For the regression analysis, a 20-year (largely MLS covering) time period is selected in order to avoid heteroscedasticity (i.e. effect of different sampling frequencies/methodologies [\(e.g. Sofieva et al., 2014; Millán et al., 2016\)](#) between different types of satellite ~~datasets~~data sets) as SWOOSH relies heavily on MLS (UARS and Aura) data records. Additionally, it also covers a period when the stratospheric chlorine loading was increasing (1991-1998) and decreasing (2005-2016) and RF has enough sample to include different characteristics of ozone variability. The regression model has 5 terms: Passive ozone (PO_3), HCl mixing ratio (HCl), methane mixing ratio (CH_4) as well as observation-model total column difference ($dTCO$) and Mg II

solar flux term ($MgII$). The PO_3 , HCl and CH_4 terms account for possible biases in CTM profiles due to transport in different stratospheric regions (e.g. Strahan et al., 2011; Feng et al., 2021). $dTCO$ is an ideal learner for the lower stratospheric ozone transport as total column ozone measurements have much smaller retrieval errors (e.g. Petropavlovskikh et al., 2019), hence they provide a good constraint for the possible biases in ERA-5 stratospheric transport (e.g. Ploeger et al., 2021). TOMCAT has 203 spectral bins in the photolysis scheme (e.g. Dhomse et al., 2016). Hence, the $MgII$ solar flux term is included to account for possible biases in the representation of the 11-year solar flux variability (e.g. Haigh et al., 2010; Dhomse et al., 2013) or the use of coarse spectral bins (e.g. Sukhodolov et al., 2016).

Overall, there are ~~thus five learners~~ five explanatory variables in the regression model ~~that for individual grid points and these are taken from TOMCAT output fields.~~ The regression model can be represented as:

$$dO_3 = \beta_1 PO_3 + \beta_2 HCl + \beta_3 CH_4 + \beta_4 dTCO + \beta_5 MgII + residuals \quad (1)$$

where $\beta_1, \beta_2, \beta_3, \beta_4$ and β_5 can be considered as the contribution coefficient for a given explanatory variable and PO_3 , HCl , CH_4 are TOMCAT monthly mean zonal mean tracers. For the calculation of $dTCO$ we use Copernicus Climate Change Service (C3S) total ozone data (1979–2018). The C3S total column product is a combination of total column data from 15 sensors using gap-filling assimilation methods and is obtained via <https://cds.climate.copernicus.eu/cdsapp#!/dataset/satellite-ozone?tab=overview> (last access: 1 May 2021). For the years 2019 and 2020, we use level 3 total column data from the Ozone Monitoring instrument (OMI) V3 that is obtained via <https://search.earthdata.nasa.gov> (last access: 1 May 2021). The Mg II index is obtained from <http://www.iup.uni-bremen.de/UVSAT/datasets/mgii> (last access: 1 May 2021). We assume long-term chemical ozone changes are realistically represented by TOMCAT chemistry (e.g. Feng et al., 2007; Chipperfield et al., 2017; Dhomse et al., 2019) ~~and so,~~ hence all the predictor time series are detrended and normalised between 0 to 1.

4 Results

Atmospheric chemical models are ideal tools for understanding/simulating past (and future) ozone changes, as they combine up-to-date knowledge about various physical and chemical processes using a mathematically consistent framework. Different models use different combinations of chemical and dynamical schemes to represent important processes in the atmosphere. However, some of these processes are computationally expensive, hence they are represented by somewhat simplified parameterisations. For example, many chemical models prescribe observation-based ~~sulphate~~ sulfate surface area density (SAD) to represent the effects of volcanically enhanced stratospheric aerosol ~~to simulate~~ for simulating heterogeneous chemistry which leads to ozone loss (Dhomse et al., 2015). Many models also prescribe surface concentrations of ~~GHGs~~ greenhouse gases (GHGs) and ODSs rather than emission fluxes. CTMs such as TOMCAT use dynamical forcing fields from (re)analyses ~~datasets~~ data sets such as ERA-Interim or ERA-5. Hence CTMs are subject to ~~the~~ possible inhomogeneities due to changes in the number of assimilated observations, as well as other deficiencies (e.g. missing processes) in the forward model used in the assimilation system. On the other hand, observational ~~datasets~~ data sets are also subject to errors associated with the measurement techniques, instrument degradation and retrieval algorithms. Hence, almost all chemical models may be expected to show a bias against

observational data records, either because of model deficiencies or errors in the ~~observational datasets~~observations. However, chemical models do use a consistent chemical scheme, so we can assume that ~~model-observations~~chemical model-observation differences are largely due to uncertainties in the forcing fields such as meteorology (e.g. winds, temperature) and chemical parameterisations (e.g. reaction rates, solar fluxes, photolysis schemes). CTMs have the distinct advantage in terms of dynamics as they are forced with up-to-date reanalysis data, although with the above-noted caveat of possible inhomogeneities in observations used in the assimilation systems. ~~Therefore, Hence, recently initiated SPARC Reanalysis Intercomparison project (S-RIP) is aimed at providing guidance on future reanalysis activity. S-RIP also plans to perform comprehensive evaluation and intercomparison of different reanalysis data products; for details see~~ <https://www.sparc-climate.org/sparc-report-no-10>. ~~Here,~~ we train the ML algorithm on the model-observation differences for the period that has relatively good temporal sampling. Estimated parameters are then used to simulate differences for the entire (1979–2020) time period. In this section, we analyse model-observation biases associated with individual predictors and compare the ML-corrected data against a variety of observation-based ~~datasets~~data sets.

4.1 Model biases

Figure 1 shows climatological (2006–2020) monthly zonal mean differences between ~~SWOOSH and TOMCAT ozone profiles.~~ TOMCAT ~~ozone and SWOOSH ozone profiles (TOMCAT minus SWOOSH).~~ TOMCAT profiles show an almost ~~symmetrical structure with positive~~symmetrically structured negative biases in the upper stratosphere and ~~negative~~positive biases in the lower stratosphere. ~~The largest positive~~The largest negative biases (up to 0.8 ppm) occur in the tropical upper stratosphere (around 3 hPa) and they remain ~~positive~~negative throughout the year. The ozone lifetime at these altitudes is less than a day, hence the observed biases might be associated with deficiencies in the photochemical reactions in the model. At this altitude, ozone production is largely controlled by solar fluxes below 240 nm while longer wavelengths control ozone destruction (e.g. Haigh et al., 2010). Therefore, negative ozone biases in the upper stratosphere are most probably due to uncertainties in the solar irradiances and/or photolysis cross sections that control ozone production (e.g. Brasseur and Solomon, 2006). Furthermore, in this region of the atmosphere, ozone chemistry is ~~very temperature dependent~~mostly temperature dependent (e.g. Stolarski et al., 2010; Dhomse et al., 2013, 2016), hence the model ozone biases could be due to uncertainties in temperature-dependent reaction rates (e.g. Ghosh et al., 1997).

In the lower stratosphere the ozone lifetime ranges from months to years, hence positive biases in the ~~modelled TOMCAT~~ ozone could be due to a combination of both dynamics and chemistry. First, reduced overhead ozone could increase lower stratospheric ozone via the self-healing effect, i.e. increased ultra-violet radiation increases ozone production at lower altitudes (e.g. Haigh, 1994). Second, ozone is primarily produced in the tropical stratosphere, and its downward transport is controlled by the ~~QBO~~quasi-biennial oscillation (QBO) (e.g. Tian et al., 2006), whereas transport towards mid-high latitudes is determined by the strength of the Brewer-Dobson (BD) circulation (~~e.g. Holton et al., 1995; Weber et al., 2003, 2011; Dhomse et al., 2006~~) (e.g. Holton et al., 2006). Hence, ozone biases in the lower stratosphere are likely due to the incomplete representation of various circulation pathways in TOMCAT either due to model resolution or missing representation of key physical process in the ERA-5 reanalysis scheme (e.g. Mitchell et al., 2020) which impacts the meteorology used in the CTM.

4.2 ~~Learner-contributions~~Contribution from explanatory variables

As the exact causes of ~~model-TOMCAT~~ ozone biases are still not well understood, we use the RF model to remove ~~these biases in TOMCAT~~ them. The RF regression model coefficients are derived using 20 years (1991–1998, 2006–2018) for which SWOOSH data includes a large number of observational profiles especially from MLS on the UARS and Aura satellite platforms. The RF regression model uses ~~20-years~~ 20 years of monthly data with 80% and 20% of data points being used for training and testing, respectively. ~~Estimated~~ The estimated RF regression coefficients are then used to calculate model biases for the entire 42-year time period (1979–2020). RF-calculated ozone biases are then added to the TOMCAT time series to create the long-term gap-free ~~dataset~~ data set, hereafter labelled ML-TOMCAT.

Figure 2 shows how much variance (or R^2) of the data the RF regression model is able to explain, along with regression coefficients for individual ~~learner~~ explanatory variables. For example, R^2 ~~equal to value of~~ 0.8 indicates that the RF regression model is able to explain 80% of the biases in ~~modelled~~ TOMCAT ozone relative to SWOOSH data for the 20 years of the training period. R^2 also represents sum of the regression coefficients from individual explanatory variables. Overall, the RF regression model performance is consistently high ($R^2 > 0.8$) throughout the stratosphere, except for the mid-stratosphere which is a transition region where the TOMCAT ozone biases change from positive to negative. At high northern latitudes, mid-stratospheric R^2 values decrease to 0.6. However, since ~~TOMCAT-SWOOSH~~ TOMCAT – SWOOSH differences are much smaller here, a RF-based correction has a minimal impact on the quality of ML-TOMCAT ozone profiles.

Additionally, as expected, the RF regression coefficients are significant in different regions of the stratosphere for various ~~learner~~ explanatory variables. The passive ozone tracer seems to show the largest coefficients in the tropical mid-stratosphere, as well as varying contributions in different regions of the stratosphere. The passive ozone contribution in the tropical mid-stratospheric could be linked to the incomplete representation of NO_x -related chemical changes in TOMCAT and/or seasonal changes in the stratospheric transport in the re-analysis (e.g. Galytska et al., 2019). The HCl tracer shows significant coefficients in the upper stratosphere, where the ClO ozone loss cycle is important. It also shows significant contribution at low-mid-latitude lower stratosphere. HCl can be considered as both a dynamical and chemical proxy, as well as in the tropical-lower-stratosphere. upper stratospheric HCl is primarily produced via degradation of ozone-depleting substances and is transported downwards at high latitudes via the BD circulation (e.g. Mahieu et al., 2014). Therefore, HCl variations in this region can be considered as a proxy for the changes in the strength of the BD circulation as well as horizontal isentropic transport, especially between tropics and mid-latitudes. The CH_4 tracer term seems to show significant coefficients in the lowermost stratosphere (just above the tropopause) as well as a significant contribution around the mid-latitude sub-tropics. The CH_4 tracer contribution resembles a QBO-induced secondary circulation pattern. Interestingly, the solar term shows the largest coefficients in the mid-latitude upper stratosphere rather than in the tropical upper stratosphere, suggesting solar flux variability has only a minor contribution to the ~~model-observation~~ TOMCAT-observation biases. As expected the $dTCo$ term shows the largest contribution in the lowermost stratosphere, especially in the tropical and polar regions. Interestingly, ozone anomalies in these regions show good agreement with various satellite-based data sets (e.g. Chipperfield et al., 2017, 2018; Li

et al., 2020; Feng et al., 2021), and ~~model-TOMCAT~~ biases are much smaller. This means ~~that~~ although $dTCO$ coefficients are largest in the lowermost stratosphere, the overall bias correction contribution remains relatively small.

4.3 Comparison against merged ~~datasets~~data sets

After analysing the regression coefficients, we now present a comparison between ML-TOMCAT and available satellite-based long-term ~~datasets~~data sets. Due to key differences between satellite measurement techniques, ozone profiles are retrieved either at ~~height~~altitude or pressure levels and either in units of mixing ratio or number density. For example, MLS retrieves profiles of ozone mixing ratio on pressure levels whereas SAGE retrieves profiles of number density on ~~height~~altitude levels. Hence, merging these different ~~datasets~~data sets needs pressure, temperature or ~~height~~altitude information at a given co-location from an external source such as reanalysis data. ~~The~~ GOZCARDS and SWOOSH ~~datasets~~data sets use MERRA2 reanalysis data to convert SAGE II ozone number density profiles on fixed pressure levels (Damadeo et al., 2013). ~~Since~~, ML-TOMCAT is based on modelled ozone profiles as a function of pressure, ~~conversion-to-height~~although conversion to altitude (geopotential height) coordinates is straightforward. In particular, ML-TOMCAT data ~~was~~were processed on corresponding grids/units using ERA-5 geopotential ~~height~~height, temperature and pressure fields that are used to drive TOMCAT.

This subsection consists of two parts. First we compare ML-TOMCAT profiles with ~~datasets~~data sets using pressure coordinate systems (e.g. SWOOSH, GOZCARDS), followed by comparisons with ~~height-based datasets~~altitude-based data sets (SAGE-CCI-OMPS, SCIA-OMPS, BSVert).

4.3.1 Comparison with pressure level data

As noted earlier, we used only 20 years of SWOOSH data to train the RF model. Hence, the next obvious step is to compare ML-TOMCAT ozone with SWOOSH over the full time period. Figure 3 compares relative differences (in ~~percentage~~percent) of ML-TOMCAT with GOZCARDS and SWOOSH, respectively, as a function of latitude and pressure. ML-TOMCAT shows slightly positive biases in the middle stratosphere and somewhat negative biases in the upper and lower stratosphere with respect to both SWOOSH and GOZCARDS data. The largest biases (up to 10%) are observed in the tropical lowermost stratosphere as well as polar latitudes. However, these largest differences in the tropical lowermost stratosphere (and upper troposphere) cannot be correctly validated as most satellite retrievals show largest ~~their~~uncertainties in this region (~~Rahpoe et al., 2015; Steinbrecht et al., 2017~~) (~~Rahpoe et al., 2015; Steinbrecht et al., 2017~~). Similarly, ~~the~~for the non-MLS period, the biases in the polar stratosphere could be due to the lack of observational ozone profiles during polar ~~nights~~night.

Figure 4 shows TOMCAT, ML-TOMCAT, SWOOSH and GOZCARDS ozone time series over the equator (0° lat) at 3 pressure levels (1, 10 and 50 hPa). Supplementary Figures S1 to S10 show similar comparisons at 15°N , 15°S , 30°N , 30°S , 45°N , 45°S , 60°N , 60°S , 75°N , and 75°S latitude bins. The grey shaded area indicates the standard deviation of the ozone values within each bin for the GOZCARDS time series. The green shaded areas indicates the root mean square uncertainty of the combined ~~datasets~~data sets for each bin in SWOOSH data (σ_{rms} in Davis et al. (2016)). Overall, there is a good agreement between the ML-TOMCAT, GOZCARDS and SWOOSH time series. As seen in Figure 1, ML-TOMCAT shows significant improvements in the tropical stratosphere compared to TOMCAT.

A peculiar detail of Figure 4 is that the standard deviation in the SWOOSH time series is largest during the 1991-1999 time period, which could be due to a combination of various factors. First, ~~for this time period, SWOOSH largely relies on UARS~~
 290 ~~MLS ozone profiles. As UARS was launched in a near-circular orbit, the MLS latitudinal coverage was either 34°S—80°N~~
~~or 34°N—80°S, depending on the spacecraft yaw position which alternated every 36 days (Barath et al., 1993). Hence, for~~
~~lower latitude ozone profiles during 1990s, SWOOSH relies on SAGE II and HALOE profiles with a low spatial sampling.~~
~~Second, UARS MLS ozone profiles are retrieved at only six levels per pressure decade (Livesey et al., 2003) instead of 12~~
 295 ~~enhancement in the stratospheric aerosol loading following the Mt. Pinatubo eruption in June 1991, led to larger retrieval~~
~~errors. Even with those uncertainties in SWOOSH (and GOZCARDS), ML-TOMCAT is generally close to the satellite-based~~
~~datasets~~ data sets for the entire time period and the agreement with satellite data is greatly improved in comparison ~~with~~
~~to~~ the original TOMCAT ~~model profile~~ data. Supplementary Figures S1 to S10 also show an excellent agreement between
 ML-TOMCAT and ~~the~~ GOZCARDS/SWOOSH ~~datasets~~ data sets for other latitude bands.

300 ~~As discussed above, the ML-TOMCAT training dataset (SWOOSH) has large uncertainties in the 1990s.~~ Next we scruti-
 nise percentage differences between GOZCARDS and ML-TOMCAT on the same pressure levels. Figure 5 shows relative
 differences between TOMCAT, ML-TOMCAT and SWOOSH ozone time series with respect to GOZCARDS. As seen earlier,
 TOMCAT ozone shows up to 40% positive biases in the lower stratosphere and 10% negative biases in the upper stratosphere
 (also seen in Figure 1). In contrast, ML-TOMCAT biases are well below 5% at all levels. At 50 hPa, TOMCAT biases seems
 305 to follow QBO-type oscillations that are correctly removed in ML-TOMCAT. Similarly, at 1 hPa TOMCAT differences show
 some uneven variations that could be linked to the inhomogeneities in the ERA-5 dynamical fields that are used to force TOM-
 CAT. Furthermore, ML-TOMCAT differences show much smaller and almost linear biases at 1 hPa and lie well within the
 spread of GOZCARDS data.

Interestingly, although both GOZCARDS and SWOOSH are created by merging nearly identical ~~datasets~~ data sets, there are
 310 differences between them which are largest for the 1984 to 2004 time period. This indicates that even slight differences in
 merging ~~methods~~ methodology leads to large differences in the merged ~~dataset~~ data set. Although we use completely indepen-
 dent output from a CTM as a basis ~~dataset, data set, GOZCARDS-ML-TOMCAT differences are not significantly different to~~
~~SWOOSH differences~~ differences are within the expected discrepancy between GOZCARDS and SWOOSH data sets, espe-
 cially at 10 and 50 hPa.

315 Another notable feature in Figure 5 is that at 50 hPa, ML-TOMCAT shows largest differences during 2020, which could be
 associated with the biases in ERA-5 dynamics during that period. A TOMCAT sensitivity simulation forced with ECMWF oper-
 ational analysis data shows better agreement with MLS ozone variation during this period (~~not shown~~) (e.g. Chrysanthou et al., 2021).
 In addition, larger differences seen during 1984 (50 hPa), 1988 (10 hPa) and 1996-1999 (1 hPa) are most probably associated
 with SAGE II sampling issues and/or inhomogeneities in ERA-5 dynamical fields. However, a detailed analysis of these biases
 320 is out of scope of this study and it needs further investigation.

4.3.2 ~~Comparison with height level data~~

In this section, we

4.3.2 Comparison with altitude level data

We now compare ML-TOMCAT ozone profiles against ~~height-based merged satellite datasets~~altitude-based merged satellite
325 data sets. Figure 6 shows ~~relative difference~~the relative differences between TOMCAT/ML-TOMCAT vs SAGE-CCI-OMPS
(Sofieva et al., 2017), BSVT (Hassler et al., 2018a) as well as SCIA-OMPS (~~Arosio et al., 2018~~)datasets (~~Arosio et al., 2019~~)data
sets as a function of altitude and latitude. The top panels (a and b) compare the mean relative ~~difference~~differences between
the SAGE-CCI-OMPS ~~dataset~~data set, TOMCAT and ML-TOMCAT, respectively. Here TOMCAT shows large positive biases
(up to 20%) in the lowermost stratosphere and negative biases (up to 15%) in the upper stratosphere. On the other hand, ML-
330 TOMCAT shows only $\pm 10\%$ biases throughout the stratosphere. Larger biases are seen in the Antarctic stratosphere that could
be attributed to the limited observational ozone profiles used to construct the ~~height-based~~altitude-based merged satellite data
products. Interestingly, ML-TOMCAT shows largest biases (up to 30%) w.r.t. the BSVT data set, though TOMCAT profiles
(forced with ERA-Interim) are used ~~to construct the BSVT data set as transfer function while constructing BSVT~~ (Hassler
et al., 2018a). In addition, in the lowermost stratosphere, biases are negative in the ~~SH latitudes and positive biases~~tropics and
335 SH mid-latitudes and positive in the NH mid-high latitudes. Hence, a contributing factor for these hemispherically asymmet-
ric biases with respect to BSVT ozone profiles might be differences between ERA-Interim and ERA-5 reanalysis data (e.g.
Ploeger et al., 2021) that are used to force these two ~~datasets~~The strong data sets. The negative values in relative differences
in the lower tropical stratosphere shown with respect to the SCIA-OMPS ~~dataset~~data set in the fourth panel ~~is~~ systematic
340 throughout the time series and is thought to be related to ~~the two factors~~The first one is the rather coarse vertical grid (corre-
sponding to SCIAMACHY vertical resolution of 3.3 km) ~~and the merging methodology implemented. In the~~which makes it
sensitive to the interpolation onto the TOMCAT grid. The second is the difference in use of merging procedure implemented
for SCIA-OMPS dataset, no anomalies were computed, but OMPS was adjusted to SCIAMACHY mean values in each bin by
using Aura MLS as a transfer function and SWOOSH, so that ML-TOMCAT, trained over the MLS period using SWOOSH,
shows a negative bias w.r.t. SCIA-OMPS, which however does not show such bias w.r.t. MLS (Arosio et al., 2019).
345 Figure 7 compares TOMCAT and ML-TOMCAT profiles with the three ~~height-based ozone datasets~~altitude-based ozone
data sets with a focus on the equator (0° latitude). Supplementary Figures S11 to S20 show similar comparisons for 15°N ,
 15°S , 30°N , 30°S , 45°N , 45°S , 60°N , 60°S , 75°N , and 75°S latitude bins. Figure 8 displays the respective relative differences
with respect to the SAGE-CCI-OMPS ~~dataset~~data set which in this case is taken as a reference. In this way it is possible to
evaluate the improvement introduced by applying the ML algorithm but also have an estimation of the discrepancies between
350 different merged ~~datasets~~data sets, which is expected to be on the order of 5-10%. With respect to the comparison with the
~~datasets~~data sets on pressure vertical coordinate, the scatter between the time series is larger here, due to the larger variety of
different satellites available to produce the merged products and the fact that they have not been used in the ML training.

At about 45km in the tropics the ML algorithm seems to over-correct the negative bias shown by TOMCAT, leading to gen-
erally higher ozone values with respect to the other ~~datasets~~data sets, especially in the first half of the time series. In the middle
355 stratosphere we find the best agreement between SAGE-CCI-OMPS and ML-TOMCAT; here the expected discrepancies among

the merged ~~datasets~~data sets are comparable to the differences observed between ML-TOMCAT and SAGE-CCI-OMPS. At the peak of the ozone number profile around 25 km, we notice generally lower values for ML-TOMCAT, on average by 5%. Similar biases are observed at mid-high latitude as well as seen in Supplementary Figures S11 to S20. The strong seasonal cycle seen in the TOMCAT difference with respect to the merged ~~datasets~~data sets is largely reduced by ML-TOMCAT at this
360 altitude.

4.3.3 Polar regions

The use of ML-TOMCAT helps to fill the observational gaps especially in atmospheric regions with lack of observations and before the beginning of the 21st century, when satellite measurements were sparser. For example, polar regions during local winter cannot be observed by limb observations based on scattered sunlight. Instruments such as Aura MLS and the
365 Sounding of the Atmosphere using Broadband Emission Radiometry (SABER, Rong et al., 2008) ~~are generally~~have generally been used to fill this gap over the last two decades. ~~Also for~~For chemical models, complexities are also associated with the denitrification and dehydration (or chlorine activation) schemes that determine heterogeneous ozone losses (Grooß et al., 2018). Though most of our earlier studies showed that TOMCAT is able simulate to polar ozone losses quite realistically (e.g. Feng et al., 2007; Chipperfield et al., 2015, 2017; Dhomse et al., 2019), some systematic biases in polar stratosphere were noted in
370 ~~Feng et al. (2021); Weber et al. (2021)~~Feng et al. (2021) and Weber et al. (2021). Figure 9 compares ozone at 18 km over the ~~North Pole which also demonstrates~~Arctic demonstrating the good agreement between ML-TOMCAT and MLS in this region ~~, both during for both~~local summer and winter. ~~In the bottom panel,~~seasons. The bottom panel shows the ozone sub-column over the ~~South Pole-Antarctic~~(poleward of 70°S latitude) ~~integrated~~integrated between 12 and 20 km ~~, for~~for TOMCAT, ML-TOMCAT and MLS averaged over September-October months. The good agreement between MLS and ML-TOMCAT during
375 the ozone hole period is observed for most of the years. ML-TOMCAT enables the reconstruction of the large ozone losses which occurred in the 1980s during a phase when ozone depleting substances were on a rapid rise before the implementation of the Montreal Protocol and their phase out.

5 Summary and Conclusions

Stratospheric ozone concentrations are affected by many short- and long-term processes, hence high quality ozone profile
380 ~~datasets~~data sets are needed for accurate attribution studies. Though satellite instruments provide global measurements, due to their short mission durations various merging methodologies have been adopted to create homogeneous and gap-free long-term ozone profile ~~datasets~~data sets. Individual merging methodologies have distinct advantages and disadvantages. Atmospheric chemical models are also able to simulate chemically consistent long-term ~~datasets~~data sets, but they are prone to the deficiencies associated with the simplified parameterisations and uncertain parameters.

385 Here we have used TOMCAT CTM ozone profiles and a ~~RF~~Random Forest (RF) regression model to create gap-free ozone profile ~~dataset~~data set (ML-TOMCAT) for 1979-2020. The RF is ~~trained for~~applied to the ozone difference between the SWOOSH and TOMCAT ozone profiles by selecting 20 years of MLS ~~measurement time periods~~measurements (UARS-MLS

and AURA-MLS) as a training period. RF show consistent performance throughout the stratosphere, except at high latitudes and the mid-latitude mid-stratosphere. Overall, ML-TOMCAT shows excellent agreement with SWOOSH for the entire time
 390 period (1984–2020), though somewhat larger differences are ~~visible~~ apparent for the period where limited ozone measurements are available for SWOOSH construction. We also find that ML-TOMCAT shows better agreement with satellite-based merged ~~datasets~~ data sets which use pressure as the vertical coordinate (e.g. SWOOSH, GOZCARDS) but weaker agreement with the ~~datasets which use height~~ data sets which use altitude (e.g. SAGE–CCI–OMPS, SCIA–OMPS). We find that at almost all stratospheric levels ML-TOMCAT ozone concentrations are well within uncertainties ~~in the observational datasets~~. ~~The of~~
 395 the observational data sets. TOMCAT ozone profiles outside the 316 – 1 hPa range can be considered as (slightly modified) TOMCAT profiles. For the next version of ML-TOMCAT ~~dataset is thus~~, we aim at correcting tropospheric ozone profile biases using merged tropospheric ozone profile data sets described in Tropospheric Ozone Assessment Report (TOAR). Presently, the ML-TOMCAT V1.0 data set is ideally suited for the evaluation of chemical model ozone profiles from the tropopause to 0.1 hPa. ML-TOMCAT V1.0 ~~data~~ ozone profile data on pressure and altitude levels in mixing ratios and number density units is
 400 publicly available via <https://zenodo.org/record/4997959#.YNzleUIKiUk>.

Data availability. We thank Sean Davies for SWOOSH data which is publicly available via <https://csl.noaa.gov/groups/csl8/swoosh/>. We also thank Lucien Froidevaux (Lucien.Froidevaux@jpl.nasa.gov) for GOZCARDS v2 data. SAGE–CCI–OMPS was obtained via <http://www.esa-ozone-cci.org>. SCIA–OMPS data is available upon email request to AR or MW. BSVert data were obtained from <https://zenodo.org/record/1217184> (Hassler et al., 2018b) . ML-TOMCAT V1.0 data is publicly available via <https://zenodo.org/record/#.YNzleUIKiUk> (Dhomse et al., 2021)

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Competing interests. The authors have no competing interests

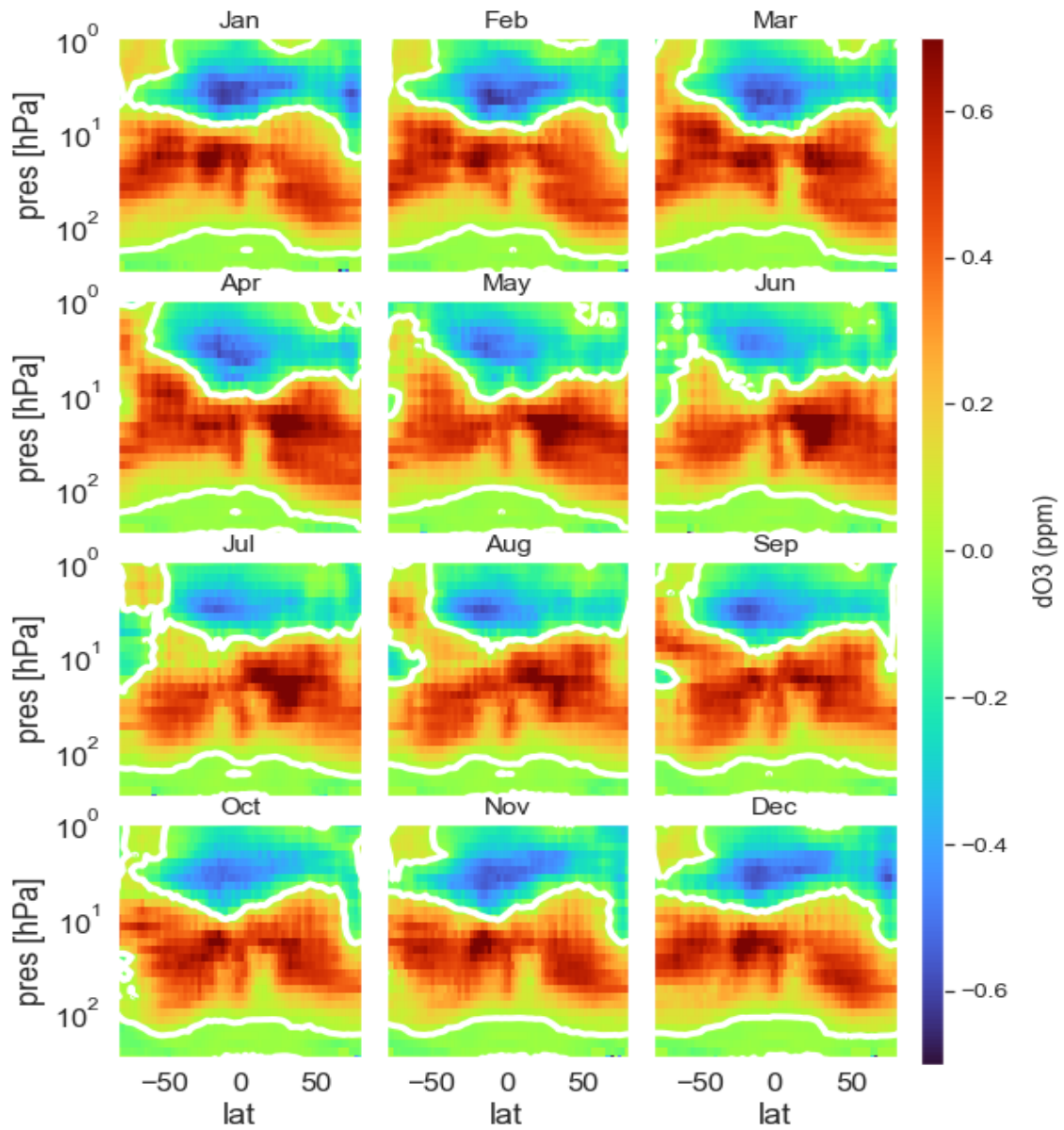


Figure 1. Latitude-pressure cross sections of the climatological (2001-2020/2006-2020) monthly mean difference (ppm) between TOMCAT and SWOOSH (Davis et al., 2016) and TOMCAT ozone profiles.

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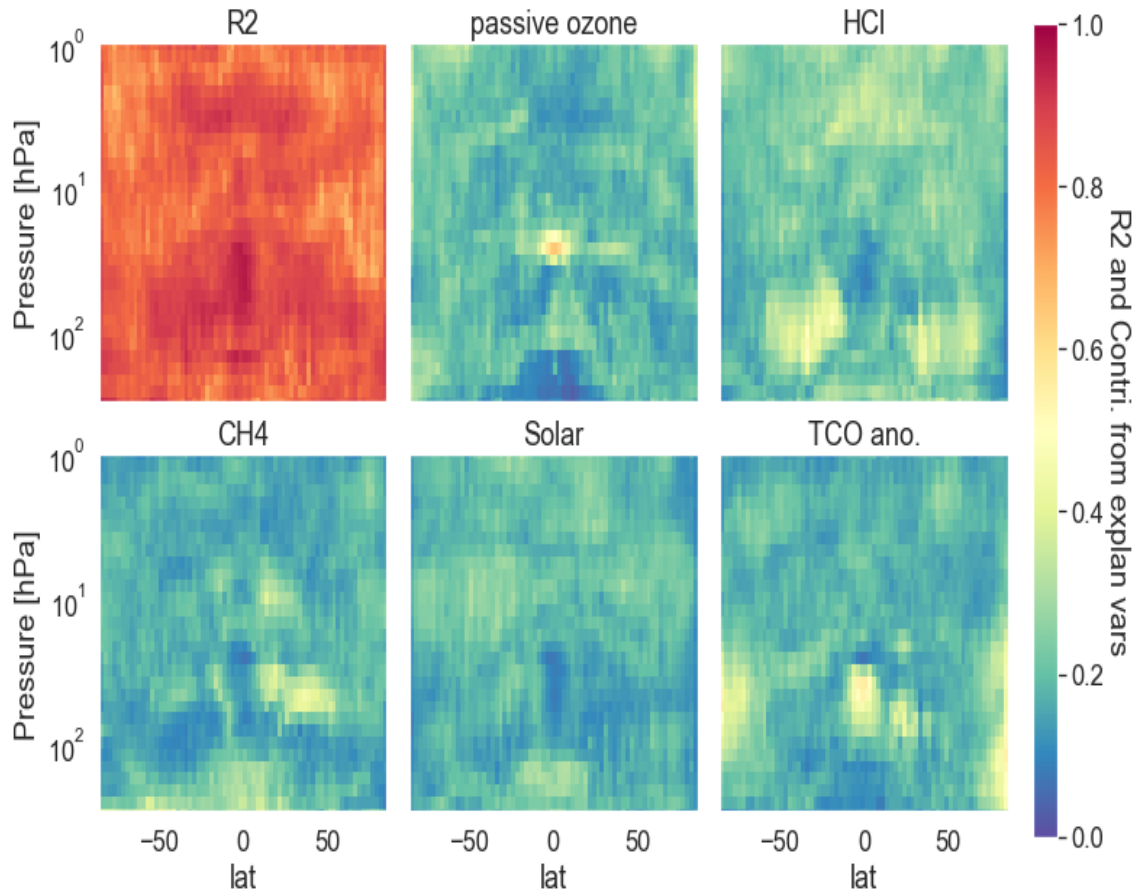


Figure 2. Latitude-pressure cross sections of the variance (R^2) and regression coefficients from passive ozone, HCl , CH_4 , solar and total column ozone anomaly (see main text equation 1).

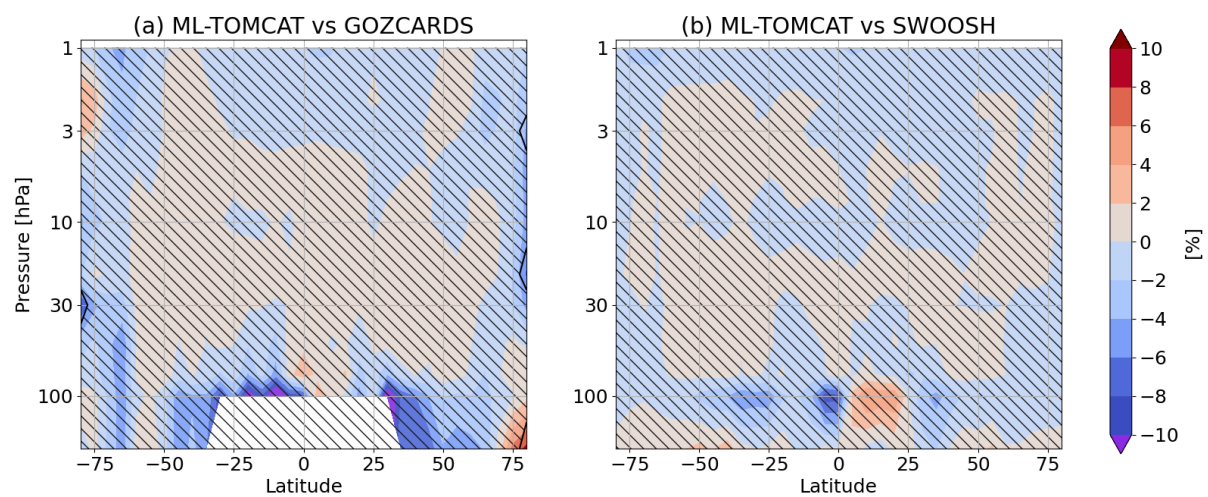


Figure 3. Relative differences (in percent) as a function of pressure and latitude between ML-TOMCAT and (a) GOZCARDS V2 (Froidevaux et al., 2019) and (b) SWOOSH (Davis et al., 2016). Stippling indicates regions where differences are statistically insignificant.

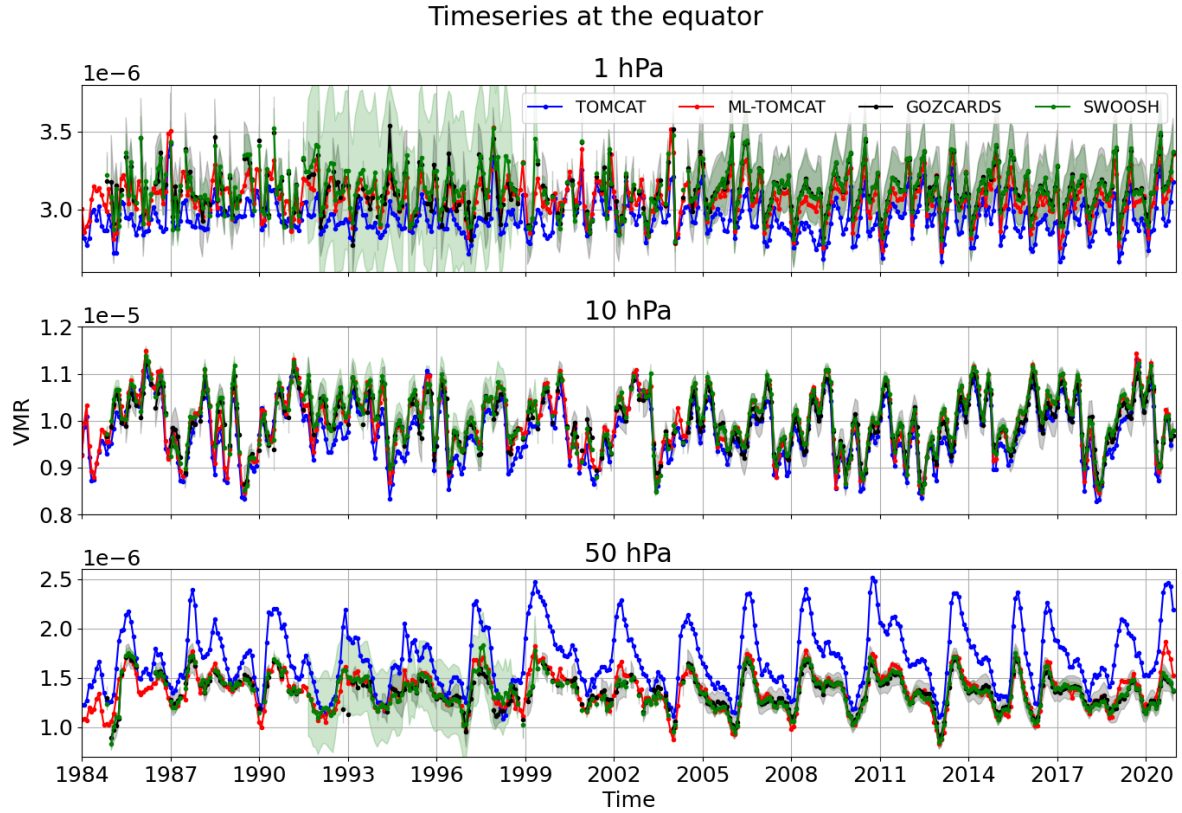


Figure 4. Comparison between TOMCAT (blue lines) and ML-TOMCAT (red lines) ozone mixing ratios over the equator (0°) at (a, top) 1 hPa, (b, middle) 10 hPa and (c, bottom) 50 hPa. Satellite-based ozone mixing ratios from GOZCARDS (Froidevaux et al., 2019) and SWOOSH (Davis et al., 2016) ~~datasets~~data sets along with their uncertainty estimates (shaded) are shown with black and green coloured lines, respectively.

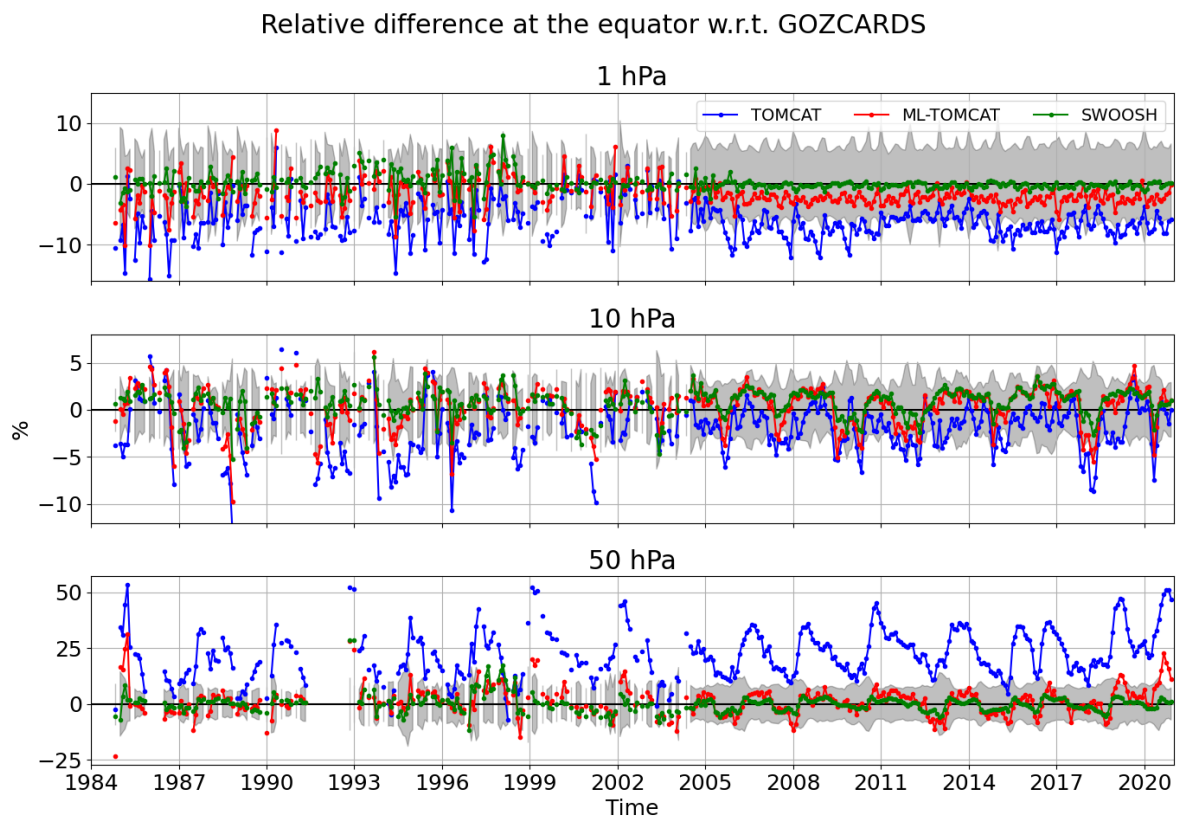


Figure 5. As Figure 4 but for the residuals, i.e. relative differences between SWOOSH (green), TOMCAT (blue) and ML-TOMCAT (red) ozone with respect to GOZCARDS ozone.

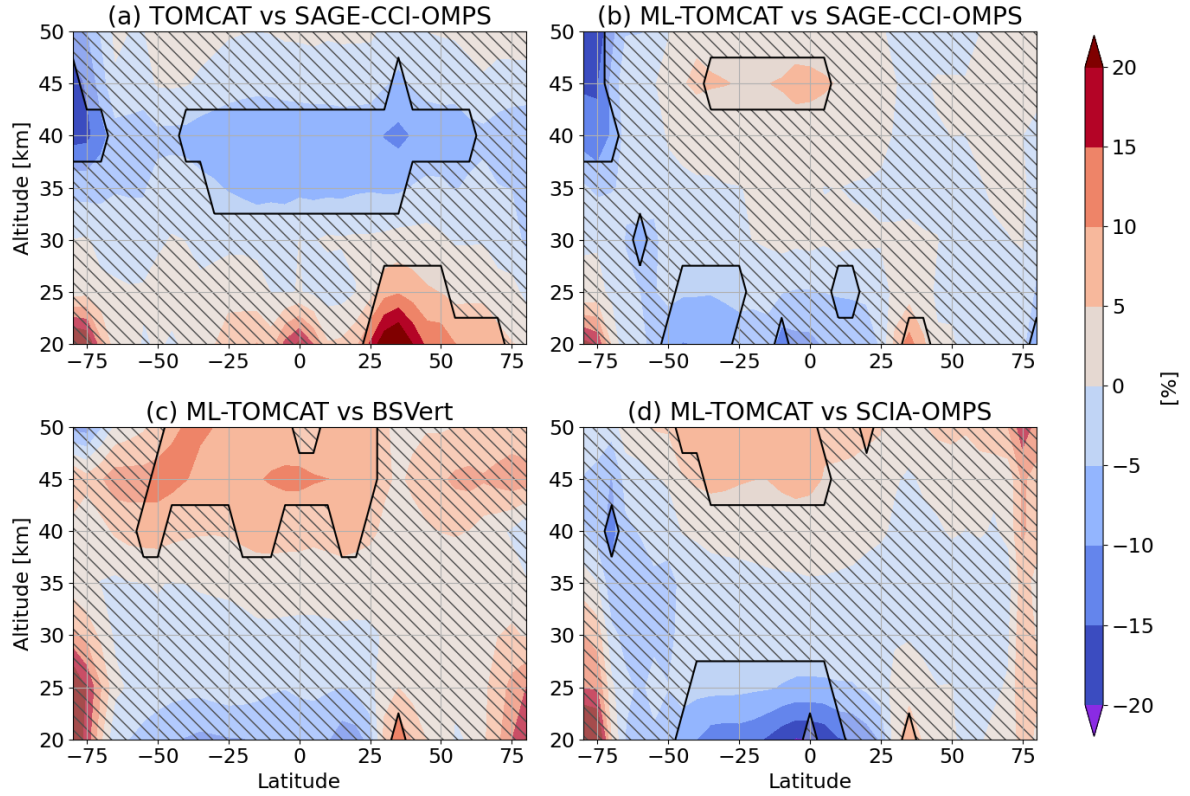


Figure 6. Relative difference (%) as a function of latitude and altitude between (a) TOMCAT, ~~ML-TOMCAT~~ versus SAGE-CCI-OMPS (1985-2019) and ~~the three considered datasets: ML-TOMCAT~~ versus (b) SAGE-CCI-OMPS (1985-2019), (c) BSVert (1985-2017) and (d) SCIA-OMPS (2002-2019), averaged over the respective time series. Stippling indicates regions where differences are statistically insignificant.

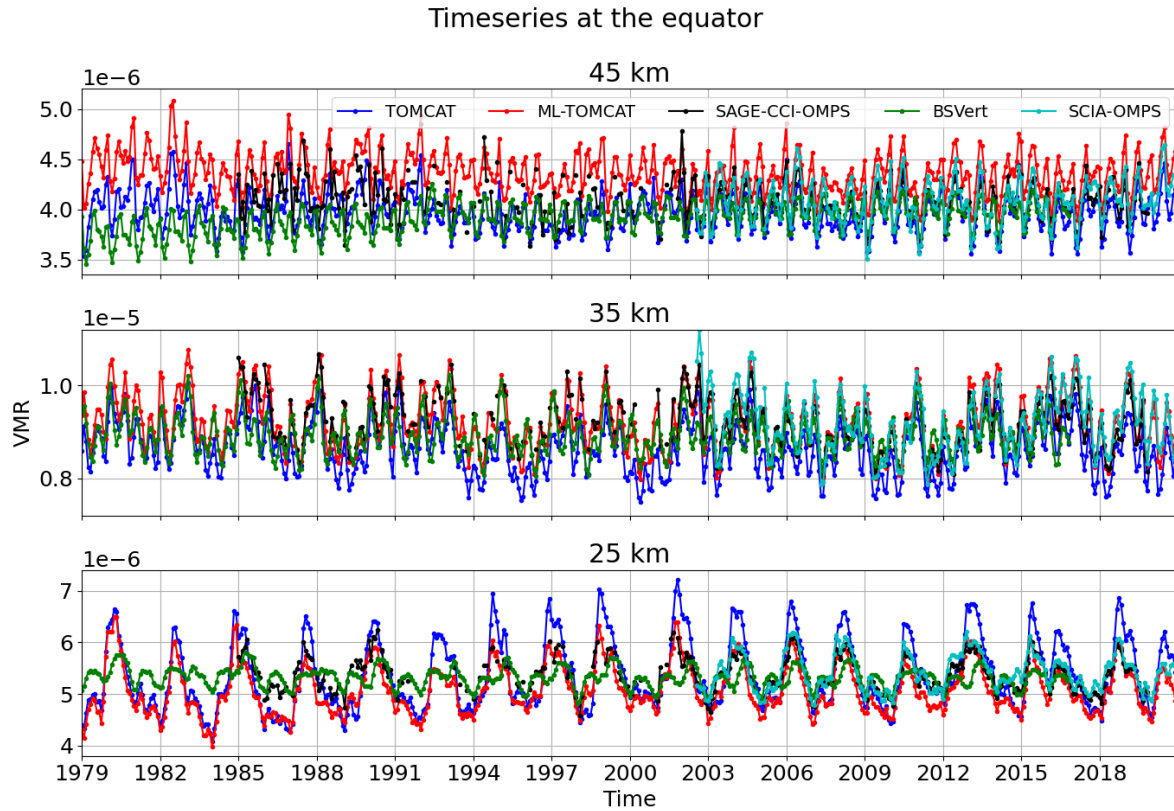


Figure 7. Comparison between TOMCAT (blue lines) and ML-TOMCAT (red lines) ozone mixing ratios over the equator (0°) at (a, top) 45 km, (b, middle) 35 km and (c, bottom) 25 km. Satellite-based ozone mixing ratios from SAGE-CCI-OMPS, BSVert (Hassler et al., 2018a) and SCIA-OMPS (Arosio et al., 2019) ~~datasets~~data sets are shown with black, green and ~~cyan-coloured~~cyan-coloured lines, respectively.

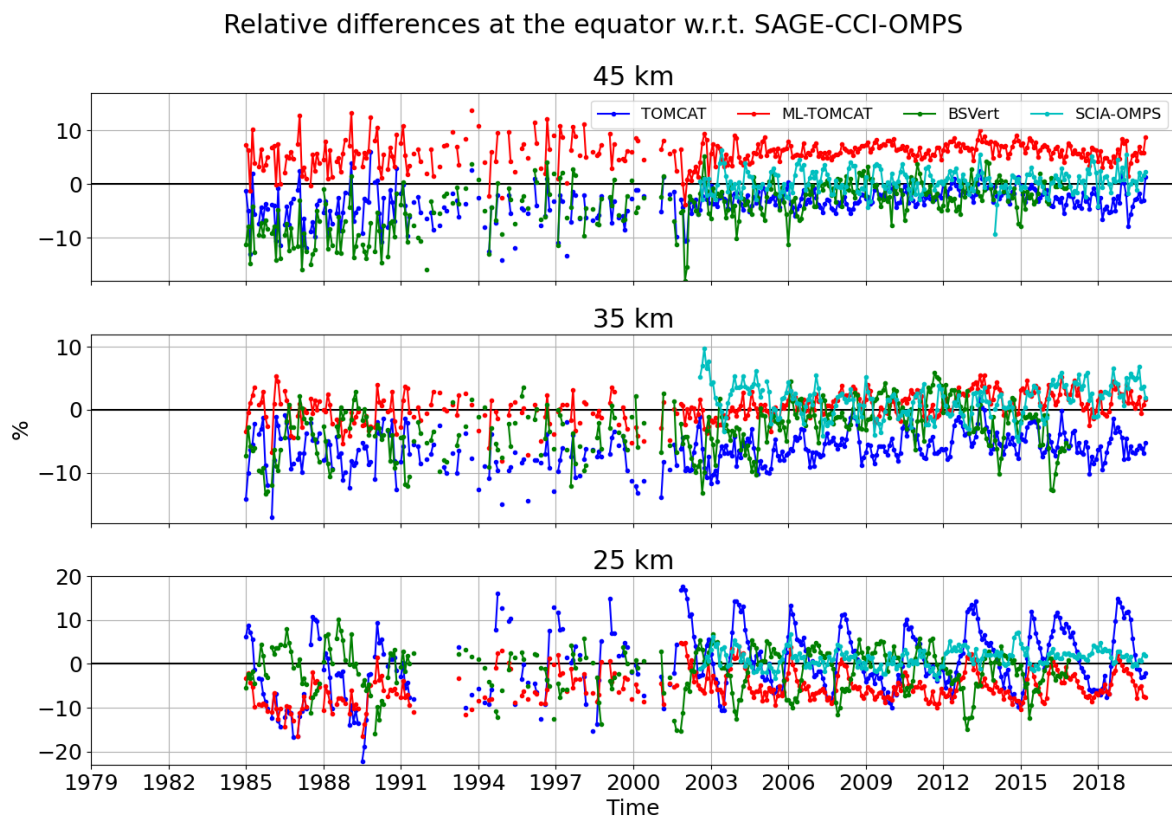


Figure 8. Same as Figure 7 but for the residuals, i.e. relative differences between TOMCAT (blue), ML-TOMCAT (red), BSVert (green) and SCIA-OMPS (cyan) ozone with respect to SAGE-CCI-OMPS.

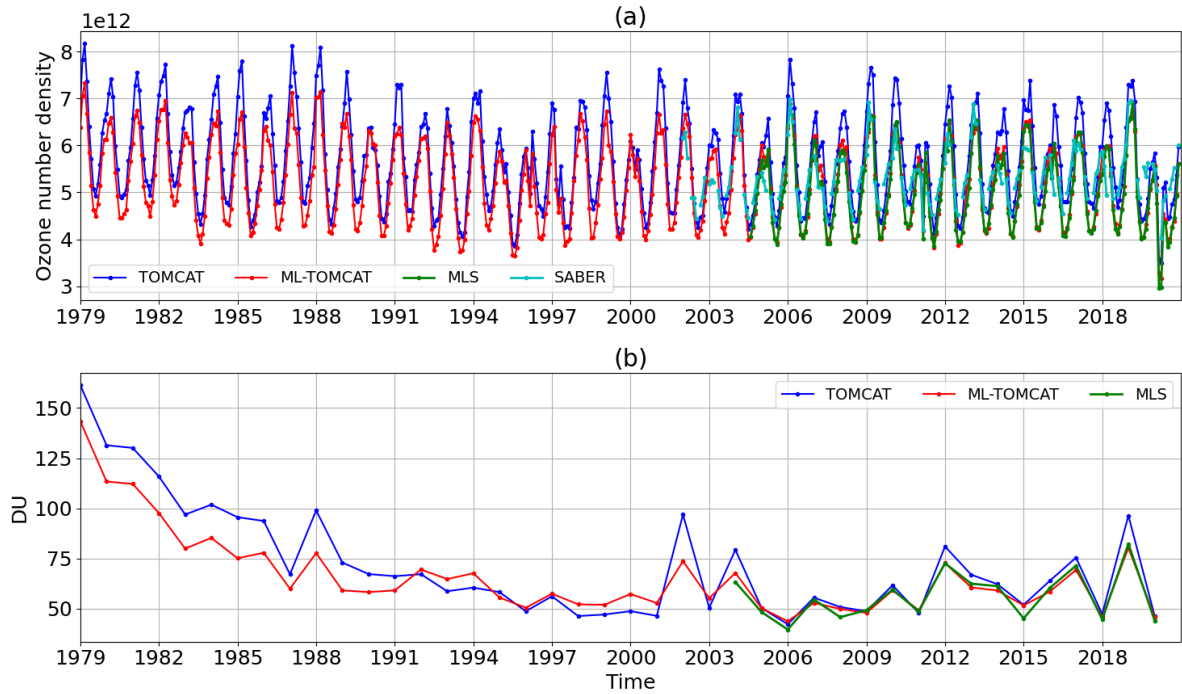


Figure 9. (a) Ozone concentration time series (molecules cm^{-3}) at 18 km over the North Polar-Arctic region (latitudes poleward of 70°N). Aura-MLS and the Sounding of the Atmosphere using Broadband Emission Radiometry (SABER, Rong et al., 2008) data are superimposed on TOMCAT and ML-TOMCAT timeseriestime series. (b) Mean ozone sub-column (DU) between 12-20 km for September and October each year over the South Polar-Antarctic region (latitudes poleward of 70°S).