



1 **The Rocklea Dome 3D Mineral Mapping Test Data Set**

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9

10 **ABSTRACT**

11 The integration of surface and subsurface geoscience data is critical for efficient and
12 effective mineral exploration and mining. Publicly accessible datasets to evaluate the various
13 geoscience analytical tools and their effectiveness for characterisation of mineral assemblages
14 and lithologies or discrimination of ore from waste are however scarce. The open access
15 Rocklea Dome 3D Mineral Mapping Test Data Set ([Laukamp, 2020;](https://doi.org/10.25919/5ed83bf55be6a)
16 <https://doi.org/10.25919/5ed83bf55be6a>) provides an opportunity for evaluating proximal and
17 remote sensing data, validated and calibrated by independent geochemical and mineralogical
18 analyses, for exploration of channel-iron deposits (CID) through cover. We present
19 hyperspectral airborne, surface and drill core reflectance spectra collected in the visible-near
20 infrared and shortwave infrared wavelength ranges (VNIR-SWIR; 350 to 2500 nm), as well as
21 whole rock geochemistry obtained by means of X-Ray fluorescence analysis and loss on
22 ignition measurements of drill core samples.

23 The integration of surface with subsurface hyperspectral data collected in the frame of
24 previously published Rocklea Dome 3D Mineral Mapping case studies demonstrated that about
25 30% of exploration drill holes were sunk into barren ground and could have been of better use,
26 located elsewhere, if airborne hyperspectral imagery had been consulted for drill hole planning.
27 The remote mapping of transported Tertiary detritals (i.e. potential hosts of channel iron ore



28 resources) versus weathered in situ Archaean geology (i.e. barren ground) has significant
29 implications for other areas where “cover” (i.e. regolith and/or sediments covering bedrock
30 hosting mineral deposits) hinders mineral exploration. Hyperspectral remote sensing represents
31 a cost-effective method for regolith landform mapping required for planning drilling programs.
32 In the Rocklea Dome area, vegetation unmixing methods applied to airborne hyperspectral
33 data, integrated with subsurface data, resulted in seamless mapping of ore zones from the
34 weathered surface to the base of the CID – a concept that can be applied to other mineral
35 exploration and mineral deposit studies. Furthermore, the associated, independent calibration
36 data allowed to quantify iron oxide phases and associated mineralogy from hyperspectral data.
37 Using the Rocklea Dome data set, novel geostatistical clustering methods were applied to the
38 drill core data sets for ore body domaining that introduced scientific rigour to a traditionally
39 subjective procedure, resulting in reproducible objective domains that are critical for the
40 mining process.

41 Beyond the already published case studies, the Rocklea Dome 3D Mineral Mapping
42 Test Data Set has the potential to develop new methods for advanced resource characterisation
43 and develop new applications that aid exploration for mineral deposits through cover. The here
44 newly presented white mica and chlorite abundance maps derived from airborne hyperspectral
45 highlight the additional applications of remote sensing for geological mapping and could help
46 to evaluate newly launched hyper- and multispectral spaceborne systems for geoscience and
47 mineral exploration.

48

49 **Key words:** Channel Iron ore Deposits, regolith, hyperspectral remote sensing, hyperspectral
50 drill core sensing, geochemistry



51 **1. INTRODUCTION**

52

53 The three dimensional (3D) geologic case history of the Rocklea Dome located in the
54 Hamersley Province (Western Australia) targeted the use of reflectance and emission
55 spectroscopy for measuring mineralogy and geochemistry specific to the exploration and
56 characterisation of economic Tertiary channel iron ore deposits in a terrain obscured by
57 weathered, transported materials. This public case history was generated by CSIRO's Western
58 Australian Centre of Excellence for 3D Mineral Mapping (C3DMM), which was operated from
59 2009 to 2012 and had the primary aim of generating and demonstrating the capabilities for
60 “scalable” 3D mineral mapping from the continental to the prospect scales (Cudahy, 2016).
61 The Rocklea Dome project was established in collaboration with Murchison Metal Ltd, who
62 granted C3DMM access to their drill hole dataset, consisting of 14 diamond cores and 180
63 reverse circulation drill holes. These drill holes were designed using traditional exploration
64 mapping technologies, such as published geology maps and geophysical data (magnetics and
65 radiometrics).

66 Key achievements of the Rocklea Dome 3D Mineral Mapping case study include:

- 67 • Based on the kaolin crystallinity index derived from surface and sub-surface
68 hyperspectral data (Cudahy, 2016), drill holes were identified that were sunk at surface
69 into barren (i.e. bedrock) weathered material. That is, approximately one third of the drill
70 holes need not have been drilled or would have been located differently had surface
71 mineral mapping data, e.g. airborne hyperspectral imagery, been used during drill hole
72 planning. This represents potential significant savings in time, money and environmental
73 disturbance.
- 74 • Characterisation of clay mineralogy associated with distinct domains of the CID and its
75 cover (i.e. kaolin group vs. Al-smectites vs. Fe-smectites) suggested that clay mineral



76 assemblages as well as calcrete atop buried CIDs have a different composition when
77 compared to regolith covering adjacent areas. That could represent useful information
78 when exploring for CIDs through regolith cover.

- 79 • Quantification of iron oxide phases and associated mineralogy derived from
80 hyperspectral data and validated using X-ray diffractometry and geochemistry (Haest et
81 al., 2012 a,b):
- 82 ○ iron (oxyhydr-)oxide content: RMSE of 9.1 weight % Fe
 - 83 ○ Al clay content: RMSE 3.9 weight % Al_2O_3
 - 84 ○ hematite/goethite ratio: RMSE 9.0 weight % goethite
 - 85 ○ spatial characterisation of vitreous vs. ochreous goethite
- 86 • Geological modelling the iron ore resource of the Rocklea Dome CID (Haest et al., 2012
87 a,b; Cudahy, 2016; Fouedjio et al., 2018), which was reported by Dragon Resources in
88 2012 to be 72.6Mt (at 53% Fe cut-off) with 54.4% Fe, 7.2% SiO_2 , 2.7% Al_2O_3 , 0.031%
89 P and 11.2% LOI.
- 90 • Improvement of quality of mineral maps by application of vegetation unmixing methods
91 (Haest et al, 2013), resulting in seamless mapping of ore zones from the weathered
92 surface to the base of deposit (Cudahy, 2016).

93 All the above points showcase how hyperspectral data can be used for critical parts of
94 the mining cycle, especially exploration and 3D resource characterisation.

95 This article aims to provide an overview of the publicly available hyperspectral data set
96 of the Rocklea Dome, which ought to be used as a test data set for 1) data mining for exploration
97 and mining, 2) integration of independent geoscience data sets (i.e. hyperspectral,
98 geochemical), 3) resource modelling, and 4) different approaches for routine processing of
99 hyperspectral data.



100 The geological setting of the Rocklea Dome area, as well as analytical and processing
101 methods will be discussed first, after which the publicly available test data are listed as a table.
102 Example applications of the geochemical and mineralogical data for exploration, 3D mineral
103 mapping and Resource estimation are summarised briefly in the discussion.

104

105 **2. GEOLOGICAL SETTING**

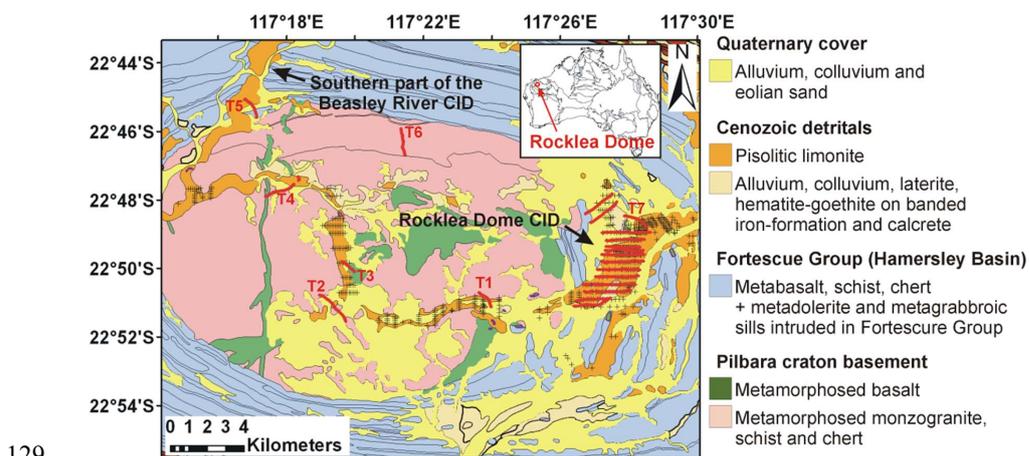
106

107 The Rocklea Dome Channel Iron Deposit is located in the Hamersley Province, which
108 is the dominant source of Australia's iron ore exports. Channel Iron Deposits (CID) are
109 economically significant formations, providing a substantial percentage of the iron ore mined
110 in Australia. A detailed overview of the geology of the Rocklea Dome and the formation of the
111 CID was provided by Haest et al. (2012b) and is briefly summarised here. The bedrock geology
112 of the Rocklea Dome comprises a monzogranite pluton and cross-cutting mafic and ultramafic
113 intrusives that form part of the Pilbara Craton. The Archean age pluton is overlain by Archean
114 to Proterozoic metasedimentary and volcanic rocks of the Hamersley Basin, enveloping the
115 central monzogranite dome (Fig. 1; Thorne & Tyler, 1996). Folding is attributed to
116 development of both the Ophthalmia and the Ashburton Fold Belt (Thorne & Tyler, 1996).

117 A meandering Tertiary palaeochannel passes over the Archean and Proterozoic rocks,
118 containing locally Channel Iron ore Deposits (CID), such as the Beasley River CID, which
119 crosscuts the north-western part of the Rocklea Dome. Channel iron ore was also drilled along
120 8 km strike length of a palaeochannel on the eastern side of the Rocklea Dome, which was
121 described by Haest et al. (2012a,b; 2013) as the Rocklea Dome CID (Figure 1). The bedrocks
122 and Tertiary channel are covered partly by regolith (e.g. Quaternary alluvium). Green
123 vegetation and dry vegetation (mostly Spinifex grass and bushes) cover the area partly.



124 A mixture of Fe-Ox pelletoids and ferruginised wood fragments below 10 mm in size
125 represent the major components of CIDs (Morris & Ramanidou, 2007). In CID systems, the
126 base of the channel often consists of a clay horizon of variable composition. The CID is capped
127 in places by calcrete and silcrete.
128



130 Figure 1: Geological map of the core of the Rocklea Dome (Haest et al., 2013; Thorne and
131 Tyler, 1996). Validation transects are indicated as red lines. T1 to T7 refer to transects
132 described in Haest et al. (2012a). The black crosses identify the position of all reverse
133 circulation drill cores intersecting the palaeochannel (i.e. pisolitic limonite in map) in the core
134 of the Rocklea Dome.

135



136 **3. METHODS AND MATERIALS**

137

138 *3.1. VNIR-SWIR drill core spectroscopy*

139 Reflectance spectra of 180 rock chips (RC) and 14 diamond drill cores (RKD) were
140 measured using CSIRO's HyChips™ system (cf. Huntington et al., 2004), which comprises a
141 TerraSpec™-based spectrometer (Malvern-Panalytical) system. In total, 7,520 reflectance
142 spectra were collected from RC samples and 66,853 reflectance spectra were collected from
143 RKD samples (Haest et al., 2012a). An automated X-Y table moves the drill core tray in a
144 snake-like pattern below the TerraSpec optical fibre at a distance of ~6 to 13 cm (depending
145 on sample type, i.e., diamond core or drill chips), while the spectral data are collected. Each
146 sample spectrum is collected from a 1×1 cm area. Four light globes are positioned 40 cm
147 above and at a small angle (off the backscatter/specular angle) to the measurement/sample
148 point. In addition to hyperspectral data, high spatial resolution (0.1 mm pixel) images are
149 collected from the core or chip tray and the sample height in the tray is measured using a laser
150 profilometer. Reflectance spectra were calibrated using a Teflon/Spectralon™ panel (see Haest
151 et al., 2012a for more details). TerraSpec™ spectra were collected in the visible to near-
152 infrared (VNIR: 380–1,000 nm) and shortwave-infrared (SWIR: 1,000–2,500 nm), with
153 sampling intervals of 1.4 nm in the VNIR and 2.0 nm in the SWIR and a wavelength accuracy
154 of ± 1 nm. The spectral resolution is 5 nm in the VNIR and between 11 and 12 nm in the SWIR.
155 The TerraSpec™ radiance spectra of each sample are first converted to apparent bidirectional
156 reflectance using the Teflon signal, which is collected at the beginning/end of each drill
157 core/drill chip tray measurement cycle. This signal is then converted to absolute reflectance,
158 based on the measurement of a Spectralon panel.

159

160 *3.2. Remote Sensing*



161 Airborne VNIR–SWIR imagery was collected using the Airborne Multispectral
162 Scanner (AMS), which is an earlier version of HyVista Corporation’s HyMap™ system (Cocks
163 et al., 1998). The AMS system collects 96 bands over the VNIR-SWIR, excluding the
164 atmospheric bands from ~1000 nm to ~1400 nm and from ~1800 to ~1950 nm, respectively.
165 For each spectral band, the average spacing of collected bands is 15 nm and the average full
166 width at half maximum is 17 nm. The AMS data over Rocklea Dome were collected in a north–
167 south direction between the 31st of July 2000 and the 2nd of August 2000, comprising a set of
168 14 flight lines, totalling in a combined length of ~280 km at a pixel size of approximately 7 m.
169 Atmospheric correction was done using MODTRAN5 (Berk et al., 2004, 2006) and SODA
170 (Rodger, 2011), based on a combination of the AMS at-sensor radiance with in-scene flight
171 parameters (e.g. latitude, longitude, sensor height, etc.). For more details about georeferencing
172 and mosaicking the single flight lines, see Haest et al. (2013).

173

174 3.3. *X-ray fluorescence analysis (XRF)*

175 XRF analysis of 11,900 RC samples (1m interval) for weight percentages of Fe, P, S,
176 SiO₂, Al₂O₃, Mn, CaO, K₂O, MgO, and TiO₂ was conducted by Kalassay Ltd. (now Bureau
177 Veritas Minerals Pty Ltd., Western Australia). A Bruker Pioneer X-ray fluorescence instrument
178 with an end window 4 kW rhodium X-ray tube was used. Sample preparation included drying
179 at 105°C for 12 hr or for 1 hr, depending on whether the sample was wet or dry, respectively.
180 Samples were then crushed to a nominal 90% passing 75 µm. The sample powders were fused
181 in a Herzog automated (RF energized) fusion furnace and cast into 40 mm diameter beads using
182 a 12:22 flux containing 5% sodium nitrate. Matrix corrections were applied using a calculated
183 alpha correction for this combination of flux, tube, and instrument geometry. Previously
184 determined weight ranges were used for both the sample and the flux weight. Kalassay Ltd.
185 used lab duplicates, internationally certified reference materials, and reference materials of the



186 same ore type as standards and reported a precision better than 0.01% for all analyses. In order
187 to evaluate the accuracy of XRF analyses undertaken by Kalassay Ltd., duplicate samples were
188 also sent to the Amdel laboratory in Cannington (Western Australia), with good correlation
189 observed (Haest et al., 2012a).

190

191 3.4. *Loss on ignition (LOI)*

192 In order to characterise the mineral assemblages present in the samples in more detail,
193 loss on ignition (i.e., LOI) measurements were undertaken on 11,900 RC samples to record the
194 mass loss of samples on heating (Haest et al., 2012a). A pre-dried portion of all samples was
195 heated in an electric furnace to 1,000°C. During this process, goethite will release its strongly
196 bonded water and its OH groups between 260° and 425°C (Strezov et al., 2010), organic matter
197 will be completely ignited by 550°C (Dean, 1974), aluminosilicate clay materials will
198 decompose between 530° and 605°C (Strezov et al., 2010), and inorganic carbon will be
199 oxidized and lost as CO₂ between 700° and 850°C (Dean, 1974).

200

201 3.5. *Sample storage*

202 Drill core trays, field samples and XRF standards are all stored at the Australian
203 Resources Research Centre (ARRC) in Kensington (Western Australia). Samples can be
204 viewed and investigated at the ARRC, using local analytical facilities.

205

206 4. SOFTWARE AND PROCESSING METHODS

207 4.1. *Processing of hyperspectral drill core data*

208 Hyperspectral drill core data were analysed using the CSIRO's The Spectral Geologist
209 software (TSG™) by interpreting the abundance, composition and/or crystallinity of selected



- 210 mineral groups and species using the Multiple Feature Extraction Method. A list of scripts
211 applied to the hyperspectral drill core and rock chip data can be found in Table 1.



212 Table 1: Base scripts and multiple feature extraction method scripts used for the Rocklea Dome
 213 3D Mineral Mapping project.

214

Product name	Minerals detected	Base algorithm	Filters/Masks	Lower stretch limit	Upper stretch limit (only applicable for composition products)	related publication
Ferric oxide abundance (Ferric_oxide_abundance.txt)	Hematite, goethite, jarosite, "limonite"	Continuum removed depth of the 900 nm absorption calculated using a fitted 2 nd order polynomial between 776 and 1050 nm 900D	R450 > R1650	0.04: low content	further developed on the basis of Haest et al. (2012a,b), which used a 4th order polynomial or 4 band ratio approach	
Hematite-goethite distribution (Hematite-goethite_dist.txt)	Hematite-goethite ratio (Cudahy and Ramaniidou, 1997)	Continuum removed wavelength of the 900 nm absorption calculated using a fitted 2nd order polynomial between 776 and 1050 nm. 900W	R450 > R1650 + 900D > 0.025	-890nm: more hematite	further developed on the basis of Haest et al. (2012a,b), which used a 4th order polynomial or 4 band ratio approach	
Ferrous iron abundance (Ferrous iron_abundance.txt)	Fe ²⁺ in silicates & carbonates, (Fe-chlorites, Fe-amphibole, Fe-pyroxene, Fe-olivine, Fe-carbonate)	(R ₆₃₀ /R ₆₅₀)(R ₁₀₂₀ /R ₁₂₃₅) Ferrans		-1.005: low content	Lankamp et al. (2012)	
opaque2 (opaque2inv.txt)	"Reduced" materials such as carbon black, sulphides and magnetite as well as Mn oxides.	albedo @ 1650 nm /650	OPAQUES. 450D(1650 > 0.25; albedo @ 1650 nm 1650 < 30%)	2: low content		
White mica and Al-smectite abundance (vmAlsmct.txt)	Abundance of white micas (e.g. illite, muscovite, paragonite, brammalite, phengite, lepidolite, margarite) and Al-smectites (montmorillonite, beidellite)	Relative absorption depth of the 2200 nm absorption for which the continuum is removed between 2120 and 2245, determined using a 3 band polynomial fit around the band with the lowest reflectance. 2200D	(R ₁₃₈ +R ₁₉₀)/(R ₁₃₆ -R ₂₁₉) 2160D2190 < 1.063	0.02: low content	further developed on the basis of Sonntag et al. (2012), which used a 4th order polynomial or 4 band ratio approach	
White mica and Al-smectite composition (vmAlsmct.txt)	Tschermak substitution of white micas, ranging from paragonite, brammalite, to illite, muscovite to phengite, and Al-smectites, ranging from beidellite to montmorillonite.	Minimum wavelength of the 2200 nm absorption for which the continuum is removed between 2120 and 2245, determined using a 3 band polynomial fit around the band with the lowest reflectance. 2200W	(R ₁₃₈ +R ₁₉₀)/(R ₁₃₆ -R ₂₁₉) 2160D2190 < 1.063	2180 nm: Al-rich mica (muscovite, illite, paragonite, brammalite, lepidolite)	further developed on the basis of Sonntag et al. (2012), which used a 4th order polynomial or 4 band ratio approach	
Kaolin abundance index	Kaolin group minerals, namely kaolinite halloysite, dickite and nacrite	2200D (Normalized depth of a fitted 4th order polynomial between 2120 and 2245 nm)	2160D (R ₂₁₃₈ +R ₂₁₉₀)/(R ₂₁₅₆ +R ₂₁₇₉) > 1.005	0.02: low content	Sonntag et al. (2012), Haest et al. (2012a,b)	
Kaolin composition index	Composition and crystallinity of kaolin group minerals ranging from well-ordered kaolinite to halloysite to dickite (and nacrite)	(R ₂₁₃₈ -R ₂₁₇₃)/R ₂₁₅₆ [(R ₂₁₅₆ +R ₂₁₉₀)/R ₂₁₇₃]	2200D > 0.005	low values = low crystallinity	Sonntag et al. (2012), Haest et al. (2012a,b)	
Carbonate abundance (carb3_jfit.txt)	carbonates w. MgOH-bearing silicates, based on left-asymmetry of CO3 feature @ 2340	Relative absorption depth of the 2340 nm absorption for which the continuum is removed between 2270 and 2370, determined using a 3 band polynomial fit around the band with the lowest reflectance. 2340D	2340D > 0.04, 2295 nm < 2340W < 2360 nm, 2290D < 0.025, 2380D > 0.117 2340D < 0.0002. Asymmetry of the 2340 absorption using a fitted 4th order polynomial between 2120 and 2370: 2340. left_asym > 1.13	0.05: low content	further developed on the basis of Sonntag et al. (2012), which used a 4th order polynomial or 4 band ratio approach	
Carbonate composition (carb3_jfit.txt)	separating calcite, dolomite, siderite, ...	Minimum wavelength of the 2340 nm absorption for which the continuum is removed between 2270 and 2370, determined using a 3 band polynomial fit around the band with the lowest reflectance. 2340W	2340D > 0.04, 2295 nm < 2340W < 2360 nm, 2290D < 0.025, 2380D > 0.117 2340D < 0.0002. Asymmetry of the 2340 absorption using a fitted 4th order polynomial between 2120 and 2370: 2340. left_asym > 1.13	2343 nm: calcite 2326 nm: dolomite	further developed on the basis of Sonntag et al. (2012), which used a 4th order polynomial or 4 band ratio approach	
White mica (+Al-smectite) abundance, refined for airborne hyperspectral imagery	Abundance of white micas (e.g. illite, muscovite, paragonite, brammalite, phengite, lepidolite, margarite) and Al-smectites (montmorillonite, beidellite)	Relative absorption depth of the 2200 nm absorption for which the continuum is removed between 2120 and 2245, determined using a 3 band polynomial fit around the band with the lowest reflectance. 2200D	(R ₁₃₈ +R ₁₉₀)/(R ₁₃₆ -R ₂₁₉) 2160D2190 < 1, 2200D 2220D > 1.5	0.04: low content	further developed on the basis of Sonntag et al. (2012), which used a 4th order polynomial or 4 band ratio approach	
Chlorite (epidote, -beotite) abundance, refined for airborne hyperspectral imagery	Abundance of chlorite (e.g. clinoclchlore, chamosite), as well as members of the epidote and biotite mineral groups	(R ₂₂₂₇ +R ₂₂₇₅)/(R ₂₂₄₁ +R ₂₂₅₉), 2250D	2250D > 1.04 & 2306 < 2320W < 2.42 & 2240 < 2250W < 2.260	1.04: high content	further developed on the basis of Sonntag et al. (2012)	



215

216 4.2. *Image Processing*

217 The processing strategy for generating geoscience products from AMS data, such as the
218 Kaolin Crystallinity (Table 1) builds on the quality control of the acquired data (Cudahy et al.,
219 2008). Well calibrated radiance-at-sensor or surface reflectance data are required for the
220 processing of airborne hyperspectral imagery. Commonly applied levelling and statistics-based
221 methods were avoided as these introduce undesirable scene-dependencies, making a
222 comparison of image products from different areas impossible. Physics-based reduction
223 models were applied to the remote sensing data, using the image processing software ENVI™.
224 Complicating effects were removed in their order of development (i.e. 1. instrument, 2.
225 atmospheric, 3. surface effects) through either normalization or offsets.

226

227 4.3. *The Multiple Feature Extraction Method*

228 In hyperspectral proximal (e.g. HyLogging™) and remote (e.g. AMS) sensing
229 technologies, the VNIR, SWIR, and thermal infrared (TIR: ca. 6000 – 14500 nm) wavelength
230 ranges are used to infer abundance and composition of various geological materials. The
231 relative intensity and wavelength position of absorption features in the reflectance spectra
232 relate to the physicochemical characteristics of the various minerals. Feature extraction
233 methods can be used to determine the mineralogy of a sample material (Cudahy et al., 2008).
234 The advantage the multiple feature extraction method (MFEM) is that the associated scripts
235 are not biased on a training dataset or spectral reference libraries, but are based only on the
236 visible and/or infrared active functional groups of minerals (see Laukamp et al., 2011, for more
237 details). As they are instrument independent, the same scripts can be applied to remote sensing
238 and proximal hyperspectral data, easing the integration of, for example, surface (e.g. HyMap)
239 and subsurface data (e.g. HyLogging™) for the purpose of visualisation in 3D or advanced



240 data analytics. Interferences of mineralogical information with other surface materials such as
241 vegetation can be evaluated by using a multiple linear regression model for unmixing
242 vegetation from hyperspectral remote sensing data (Rodger & Cudahy, 2009; Haest et al.,
243 2013). Other complications, such as spectrally overlapping materials, are removed by the
244 application of thresholds.

245

246 5. DATA PRODUCTS AND APPLICATIONS

247 Publicly accessible data of the Rocklea Dome 3D Mineral Mapping project can be
248 found on CSIRO's Data Access Portal:
249 <https://data.csiro.au/collections/#collection/CIcsiro:44783>
250 (<https://doi.org/10.25919/5ed83bf55be6a>) and are listed in Table 2. Data and other content on
251 this site are scientific research data collected by CSIRO and third parties and are made available
252 on an 'as is' basis. If any data or other material are downloaded from this site, the user does so
253 at own risk and acknowledges that such data or other content: 1) may contain general
254 statements based on scientific research and may be incomplete and not applicable to all
255 situations; 2) is not professional, scientific, medical, technical or expert advice and is subject
256 to the usual uncertainties of scientific and technical research; and 3) should not be relied upon
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258 professional scientific and technical advice should be sought prior to acting in reliance on data
259 and other material from this site. To the extent permitted by law, CSIRO excludes all liability
260 to any person for any consequences, including but not limited to all losses, damages, costs,
261 expenses and any other compensation, arising directly or indirectly from using and any
262 information or material contained in it.

263



264 Table 2: Publicly accessible data of the Rocklea Dome 3D Mineral Mapping project
 265 (<https://data.csiro.au/collections/#collection/CIsiro:44783>;
 266 <https://doi.org/10.25919/5ed83bf55be6a>)

main directory	sub directory	file name	type of data	source/IP	
DTM		DTM Rocklea 50k.00t	Digital terrain model	GWSA	
		DTM Rocklea 50k.dxf			
		DTM Rocklea 50k.evf			
		DTM Rocklea 50k.zip			
			Hardey HR DTM.00t	Digital terrain model of 100K mapsheet Hardey 2252	GWSA
			Hardey HR DTM.dxf		
			Hardey HR DTM.evf		
			Topography ENVI		
		Topography ENVI.hdr	Digital elevation model	GWSA	
		dem plus collars.csv	Digital elevation model	GWSA	
drill hole data	RC_hyperspectral_geochem	GeoscienceProductDescriptions_Pr oximalHyperspectral.xlsx	table describing multiple feature extraction scripts applied to hyperspectral data for interpretation of mineralogy	CSIRO	
		RC data.tsg	TSG-file	CSIRO	
		RC data.ini	TSG-file	CSIRO	
		RC data.bip	TSG-file	CSIRO	
		RC data cras.bip	TSG-file	CSIRO	
		RC_data_tsgexport.CSV	spectral and geochemical data exported from TSG	CSIRO	
	RKD		RKD5-7-9.tsg	TSG-file	CSIRO
			RKD5-7-9.ini	TSG-file	CSIRO
			RKD5-7-9.bip	TSG-file	CSIRO
			RKD5-7-9 cras.bip	TSG-file	CSIRO
remote sensing data	GeoTIFF AMS/	2200D_Mstd.tfw	AMS product "2200D", showing the relative abundance of Al-clays	CSIRO	
		2200D_Mstd.tif		CSIRO	
		2200WAR_2190-2205.tfw	AMS product "2200W", indicating compositional changes of Al-smectites and white micas (Al ^{VI} Al ^{IV} (Fe,Mg) ₋₁ Si ₋₁)	CSIRO	
		2200WAR_2190-2205.tif		CSIRO	
			2250_MStd.tfw	AMS product "2250D", showing the relative abundance of chlorite, epidote and/or biotite	CSIRO
			2250_MStd.tif		CSIRO
			2330_2250-2380.tfw	AMS product "Carbonate abundance", showing the relative abundance of carbonates	CSIRO
			2330_2250-2380.tif		CSIRO
			KC_NoSM_22D+216DM_3MeFi.tfw	AMS product "Kaolin crystallinity"	CSIRO
			KC_NoSM_22D+216DM_3MeFi.tif		CSIRO
	TXT AMS/		2320D_vegunm.txt	AMS product "2320D", vegetation unmixed	CSIRO
			AIOHAbVegunm.txt	AMS product "Al-clay abundance index", vegetation unmixed	CSIRO
			FeOxVegUnm.txt	AMS product "Ferric Oxide Abundance Index", vegetation unmixed	CSIRO
			SRTM_RockleaDome+HardeyRiver.txt	Digital elevation model	GWSA



Rocklea Dome exercise	StudentExercises_Rocklea.docx	Exercises for analysis of HyLogging data	CSIRO
	Answers_CIDexercises.docx	Suggested answers to exercises for analysis of HyLogging data	CSIRO
	MinSpec_Workshop_7RockleaDomeTSG_HandsOn.pptx	PPT-presentation summarising Rocklea Dome exercise and results	CSIRO

267

268 The following chapters briefly describe examples of how the provided hyperspectral
269 and geochemical proximal and remote sensing data sets can be used to address challenges for
270 the mineral resources sector.

271

272 5.1. Drill core mineralogy and geochemistry

273 Reflectance spectra collected from rock chips (RC) and diamond drill cores (RKD)
274 using CSIRO's HyChips™ system presented a cost-effective way to spatially map the major
275 ore (i.e. goethite +/- hematite) and gangue minerals (i.e. kaolinite, smectite, carbonate), apart
276 from quartz, in detail. To achieve this, the relative intensity of mineral-diagnostic absorption
277 features was calculated using a suite of batch scripts ("Geoscience Products" in Haest et al.,
278 2012a). The relative intensity of the respective absorption features correlates with the relative
279 abundance of the respective mineral, whereas the wavelength position of key absorption
280 features relates to mineral speciation (e.g. ochreous versus vitreous goethite) or determining
281 the mineral chemistry. For example, the relative abundance of iron oxides was calculated from
282 the relative depth of the ferric iron-related absorption at around 900 nm (Cudahy &
283 Ramanaidou, 1997), whereas goethite was distinguished from hematite by tracking the
284 wavelength position of the same absorption feature (Table 2 in Haest et al., 2012a).

285 Whole rock geochemistry obtained from the same drill core material showed significant
286 correlations with the Geoscience Products. Haest et al. (2012) determined an RMSE of 9.1
287 weight % Fe for the correlation between the hyperspectrally-derived iron oxide abundance and
288 the XRF weight % Fe data and an RMSE of 3.9 weight % Al₂O₃ for correlation between the

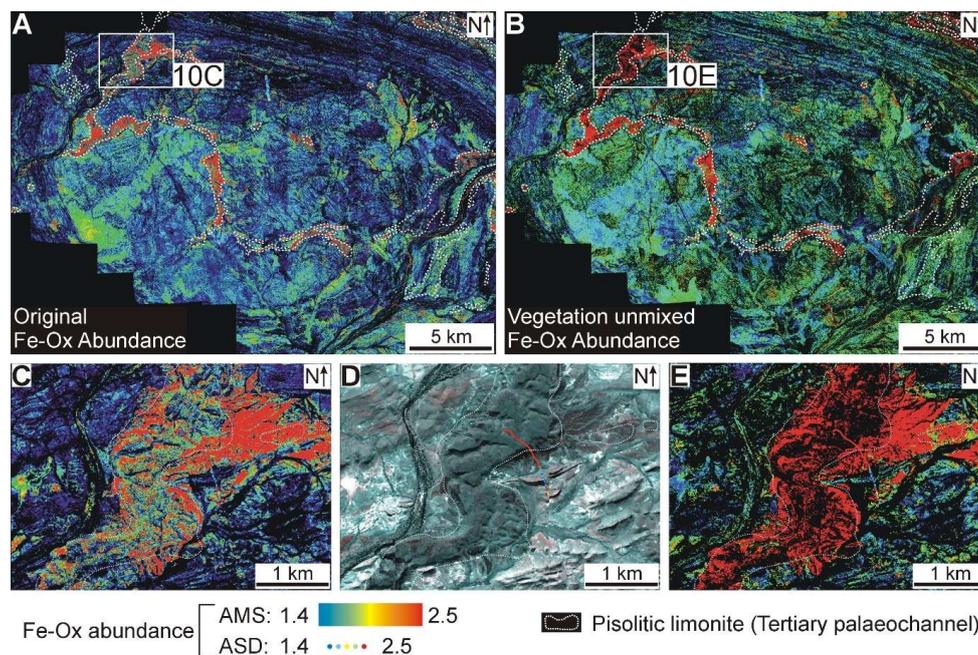


289 hyperspectrally-derived Al-clay abundance and the XRF weight % Al_2O_3 data. The errors
290 associated with the correlations were found to be due to a combination of grain size variations
291 and the transopaque behaviour of iron oxides and/or different amounts of silica, causing
292 variations in the optical depth of sample material.

293

294 5.2. *Surface mineral mapping*

295 Airborne hyperspectral surveys provide spatially contiguous mineralogical information
296 of the Earth's surface at high spatial resolution (down to circa 1 m). The relative intensity of
297 mineral diagnostic absorption features and their wavelength positions can be used to infer the
298 relative abundance of the respective minerals and even variations of single mineral species in
299 terms of their cation composition, crystallinity and hydroxylation. The Rocklea Dome case
300 study data set was used by Haest et al. (2013) to demonstrate how quantitative mineral maps
301 can be produced by validation of airborne hyperspectral data against field data, including
302 reflectance spectra and XRF data collected from surface samples. The effect of both green and
303 dry vegetation cover was unmixed at the pixel-level using the Normalised Difference
304 Vegetation Index (NDVI; e.g. Tucker, 1979) and the continuum-removed depth of the
305 cellulose-lignin absorption centred at around 2100 nm, respectively. The resulting mineral
306 mapping products have a higher spatial continuity, as well as higher accuracy of, for example,
307 mineral abundance or composition values shown in single pixels. This proved to be especially
308 useful in areas with outcropping CID, which appeared to be sub-economic from the original
309 iron oxide abundance mineral maps but showed as potentially economic CID resources when
310 the vegetation cover was unmixed (Figure 2).



311

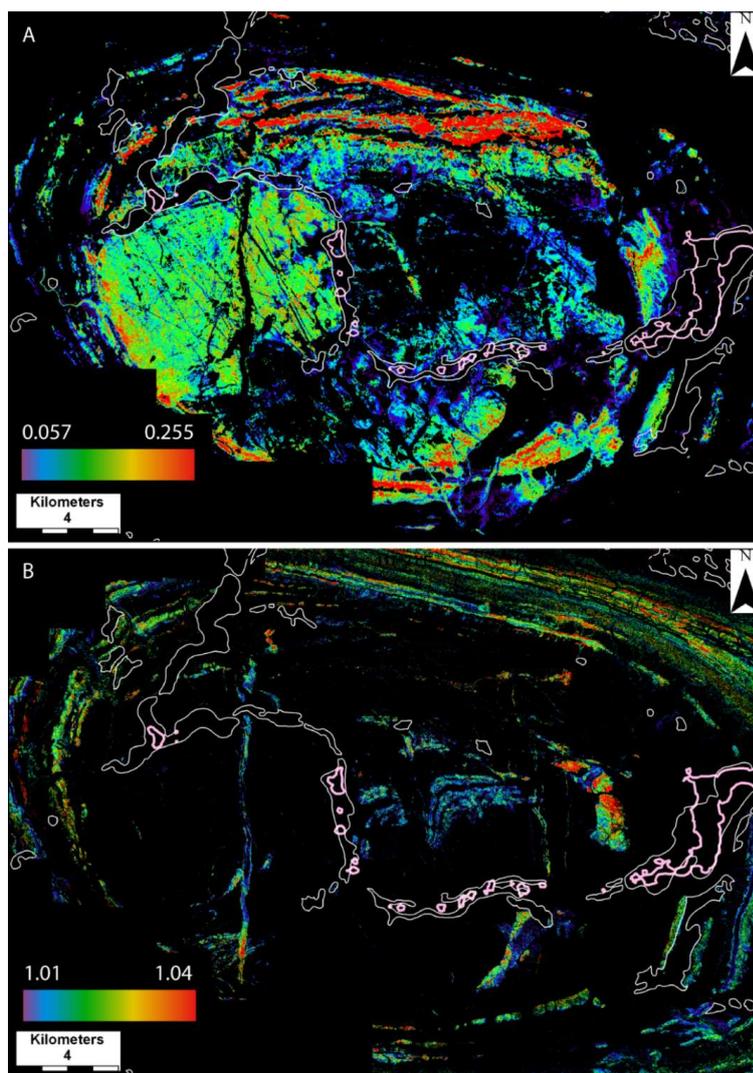
312 Figure 2: A–B: Fe-(oxyhydr-)oxide (Fe-Ox) abundance maps of the Rocklea Dome without
313 (A) and with (B) vegetation unmixing. C–E: Fe-Ox abundance maps of the southern part of
314 the Beasley River CID with (C) and without (E) vegetation removal and the false colour
315 image of this area (D). The Beasley River CID has a plateau like surface expression, with the
316 edges of the plateau clearly visible in the false colour image. These edges were mapped by
317 the Geological Survey of Western Australia as representing the boundary of the pisolitic
318 limonite (Fe-rich palaeochannel; white stippled line) (the Fe-Ox abundance measurements
319 collected along transects 1 to 7 with the TerraSpec™ are also shown for reference).

320

321 Beyond the iron oxide, kaolin and carbonate mineral maps published by Haest et al.
322 (2013), airborne hyperspectral data can be used to create numerous additional mineral mapping
323 products that can be used to address other geoscientific questions. For example, the Rocklea
324 Dome presents a wide variety of igneous units that are part of the Proterozoic basement of the
325 Pilbara Craton (Figure 1). These include 1) metamorphosed monzogranite, schist and chert, 2)
326 metamorphosed basalt, and 3) amphibolite dykes. According to the white mica abundance
327 derived from airborne hyperspectral data (green in Figure 3a), the metamorphosed
328 monzogranite contains less white mica, when compared to the metamorphosed schists which



329 are striking East-West and occur in the northern part of the Rocklea Dome (red in Figure 3a).
330 In the eastern half of the Proterozoic basement in the Rocklea Dome, white mica is much less
331 abundant to absent. This coincides with elevated amounts of chlorite (folded lithologies in the
332 centre of Figure 3b), which map out metamorphosed basalt (Figure 1). The North-South
333 striking occurrence of chlorite in the Western half of the Proterozoic basement traces an
334 amphibolite dyke. Both the white mica abundance and chlorite abundance maps can also be
335 used to map out different lithologies in the metasediments and metabasalts of the Fortescue
336 that crop out to the North and South of investigated area, demonstrating how the airborne
337 hyperspectral data can be used to map out all major lithologies occurring in the Rocklea Dome
338 case study area.



339

340 Figure 3: A & B: White mica (+Al-smectite) and chlorite (+epidote, +biotite) abundance
341 maps of the Rocklea Dome area in A and B, respectively, calculated from airborne
342 hyperspectral data using algorithms described in Table 1. Warm colours represent high
343 abundance and cool colours low abundance of the respective minerals. Black pixels have
344 been masked out as relative intensity of the absorption feature mapped in the respective
345 mineral map is below a given threshold (Table 1) and/or because of non-mineralogical effects
346 (e.g. vegetation, clouds). A shows monzogranites in the western part of the dome in green
347 colours and the Fortescue Group in the northern fringe of the dome in red colours. B
348 highlights Archean metamorphosed basalts in the eastern part of the dome structure and an N-
349 S trending amphibolite dyke in the western part of the dome. White lines indicate the surface
350 extension of the Tertiary channel as mapped by Thorne & Tyler (1996). Pink lines indicate
351 the horizontal extension of the Tertiary channel as mapped by the hyperspectral data.



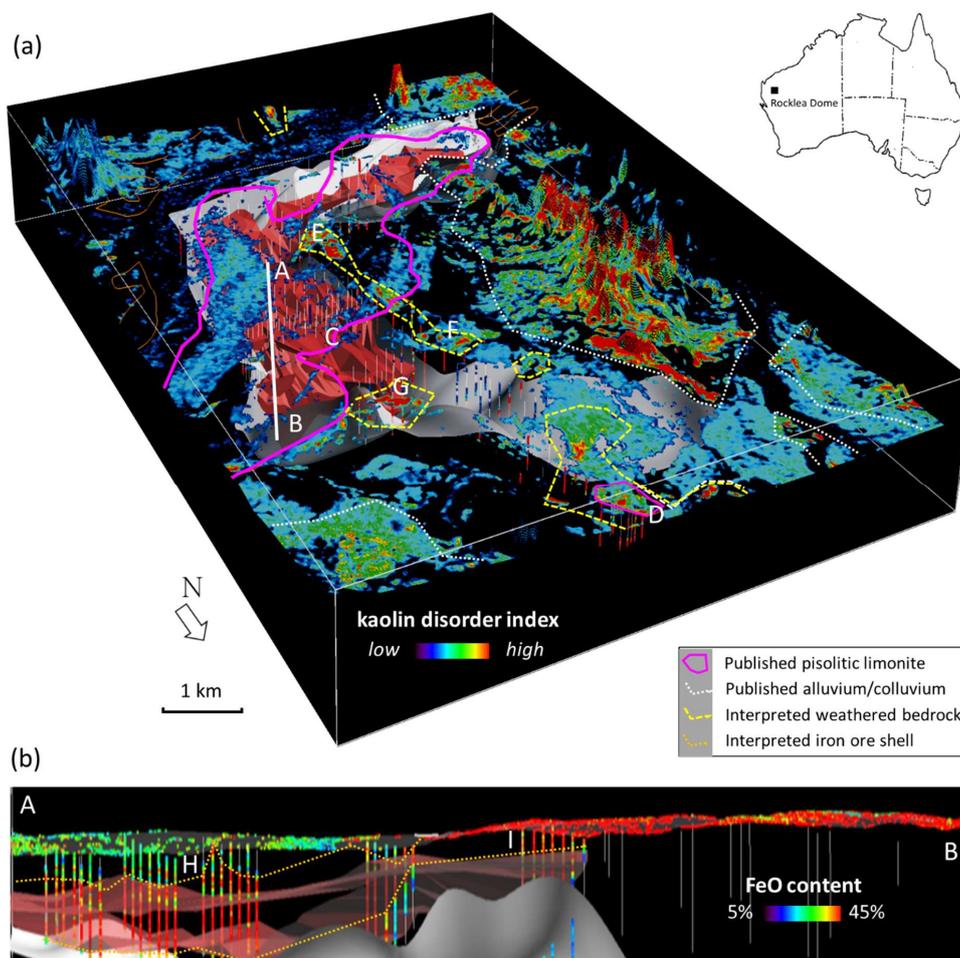
352

353 5.3. *3D Mineral Mapping*

354 The hyperspectral drill core data can be combined with airborne hyperspectral data into
355 a seamless 3D mineral model of the Rocklea Dome CID (Figure 4). For this, all hyperspectral
356 data were resampled to the same spatial resolution and imported into the 3D modelling software
357 GoCad/SKUA™. The channel basement contact that was delineated at depth using the kaolin
358 crystallinity products could also be delineated at the surface from the airborne hyperspectral
359 image. A combination of both provided a seamless surface of the channel bottom (grey surface
360 in Figure 4) that separates the basement characterised by well-crystalline kaolinite from the
361 tertiary channel sediments characterised by poorly-crystalline kaolinite. The here identified
362 channel basement contact deviates at the surface significantly from the area mapped by the
363 geological survey as palaeochannel. This suggests that drilling patterns could have been much
364 better defined if the airborne hyperspectral-based surface outline would have been available
365 prior to drilling (Cudahy, 2016).

366 As part of their 3D Geomodel Series, the Geological Survey of Western Australia
367 (GSWA) provides access to 3D models of the Rocklea Dome area via their online portal:
368 [https://dmpbookshop.eruditetechnologies.com.au/product/rocklea-inlier-2016-3d-geomodel-](https://dmpbookshop.eruditetechnologies.com.au/product/rocklea-inlier-2016-3d-geomodel-series.do)
369 [series.do](https://dmpbookshop.eruditetechnologies.com.au/product/rocklea-inlier-2016-3d-geomodel-series.do). The data can be viewed in three different formats (3D PDF, Geoscience Analyst,
370 GOCAD).

371



372

373 Figure 4: 3D mineral models of the Rocklea Dome area (Cudahy, 2016). Scene centre is
374 approximately 22.8216° latitude 117.4652° longitude. (a) A southwest oblique 3D view of
375 the Rocklea Dome study area showing kaolin disorder measured using airborne HyMap™
376 (surface) and drill core HyLogger™ (coloured vertical pegs) reflectance spectra. Warmer
377 colours (well-ordered kaolin) relate to weathered, in situ bedrock, while cooler colours
378 (poorly-ordered kaolin) relate to transported (alluvium/colluvium) materials. The interpolated
379 model of the base of the channel iron system calculated using the 3D kaolin crystallinity map
380 is shown by the shaded grey surface. The CID, which was calculated from the XRF-derived
381 % FeO (Haest et al., 2012a), is shown by a shaded red volume. Areas of weathered bedrock
382 (as mapped by Haest et al., 2012b and Cudahy, 2016) are highlighted by yellow-coloured
383 hashed lines and highlight which drill cores were sunk into barren ground. A white straight
384 line shows the location of the cross-section (A–B) presented in (b); (b) Cross section A–B in
385 (a) of the % FeO measured from the drill core and airborne imagery, which was vegetation
386 unmixed (Haest et al., 2013).

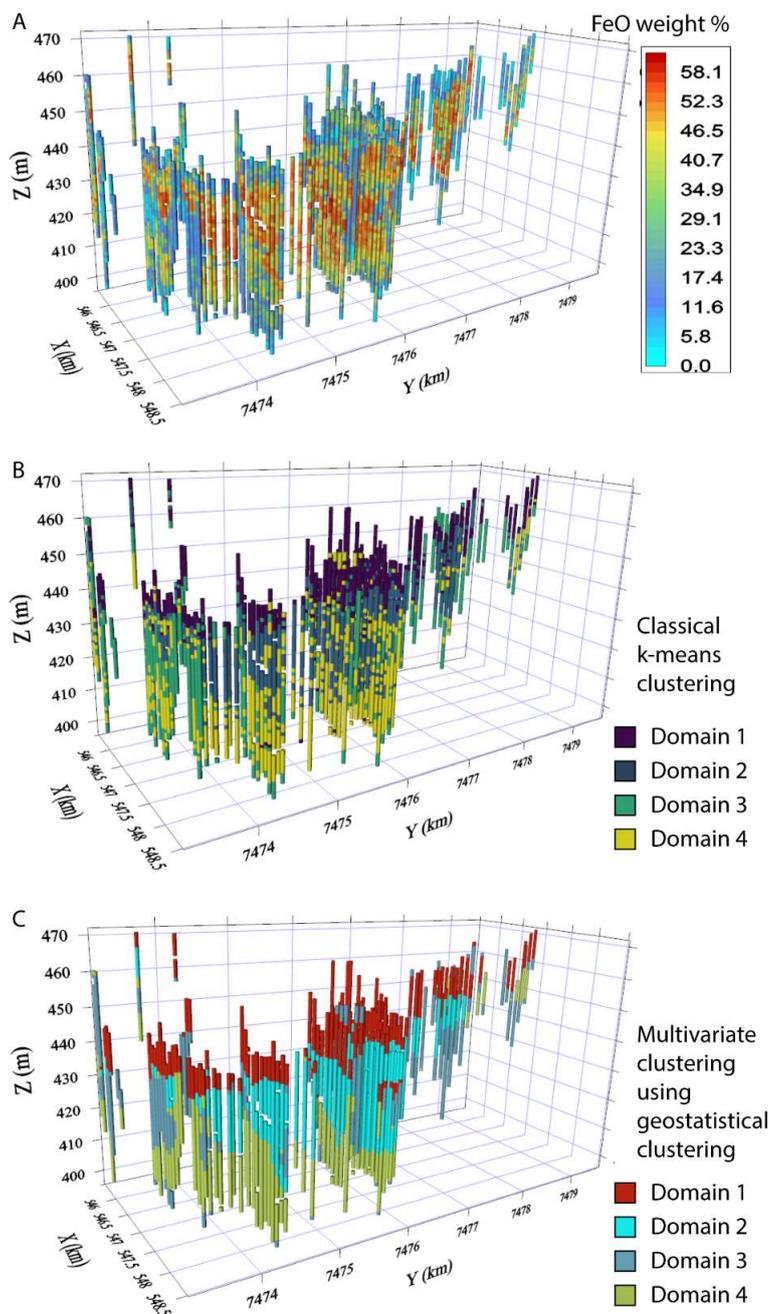


387

388 *5.4. Resource estimation*

389 Resource estimation of base and precious metal deposits requires the grouping of drill
390 hole data into domains that represent zones of homogenous properties for accurate grade
391 estimation and practical exploitation purposes. In practice, this is more than often performed
392 through a subjective time-consuming manual interpretation of sample analytical data.
393 Traditional automated clustering techniques, such as multivariate clustering and k-means, tend
394 to show poor spatial contiguity of domains in a mineral deposit. Fouedjio et al. (2017) used the
395 Rocklea Dome drill core data set to showcase how geostatistical clustering methods can take
396 spatial dependency into account (Figure 5). By integrating whole rock geochemistry and
397 hyperspectral drill core data, Fouedjio et al. (2017) revealed two distinct domains in the
398 Rocklea Dome Channel Iron Ore Deposit that are mainly characterised by four geochemical
399 variables (FeO, Al₂O₃, SiO₂ and TiO₂) and two mineralogical variables derived from
400 hyperspectral data (ferric oxide abundance and kaolinite abundance). Ore body domaining
401 through geostatistical clustering represents a method for objective samples clustering that
402 introduces scientific rigour to a traditionally subjective procedure. The robust domaining is
403 based in genuine multivariate geostatistics combining all available data. The flexible and
404 reproducible automatic domaining technique saves time, improves the understanding of
405 domains critical for exploitation of the ore and allows an easy integration of new data sets.

406



407

408 Figure 5: a) Spatial plot of FeO distribution in the Rocklea Dome Chanel Iron Deposit. (b)
409 Classical k-means clustering method using 4 domains. (c) Geostatistical spectral clustering
410 using 4 domains. Z-axis was scaled to ease visualisation. Modified from Fouedjio et al.
411 (2017).



412

413

414

415 5.5. *Teaching material*

416 The publicly available Rocklea Dome data set provided an opportunity to compile
417 training and teaching material about the application of hyperspectral drill core and chips data
418 for iron ore resource characterisation using The Spectral Geologist Software (TSG™,
419 <https://research.csiro.au/thespectralgeologist/>). Student exercises and example answers, as well
420 as a ppt for teaching are part of the data package:

421 Exercise: StudentExercises_Rocklea.docx

422 Answers: Answers_CIDexercises.docx

423 PPT for teaching: MinSpec_Workshop_7RockleaDomeTSG_HandsOn.pptx

424

425 6. CONCLUSIONS AND OUTLOOK

426 We have established an open-access dataset comprising drill core, surface and airborne
427 hyperspectral data of the Rocklea Dome area in the Hamersley Basin of Western Australia,
428 which features a wide variety of lithologies and morphologies and is prospective for channel-
429 hosted iron ore resources. The proximal and remote sensing data, together with associated
430 whole rock geochemistry are ideal for researching the geology of this economically significant
431 area and allow a thorough comparison of different geoanalytical techniques and their
432 effectiveness for resource characterisation. Combining the surface and subsurface data into 3D
433 mineral maps provides a better visual understanding of the geological environment.

434 In addition to the already published surface and subsurface mineral mapping products,
435 many more Geoscience Products can be generated to better understand this geologically
436 complex area. The here newly presented white mica and chlorite abundance maps clearly



437 highlight the potential for mapping out different sections of the Archaean monzogranitic
438 basement as well as different generations of mafic intrusives. Of particular interest are the
439 contact zones between the mafic dykes and their host rocks, as they could help to better
440 understand the intensity of alteration within the dyke and within the host granite as well as the
441 associated fluid-rock interaction processes.

442 The teaching material provided together with this open-access dataset aims to support
443 training of geoscience graduates and post-graduates in the potential applications of
444 hyperspectral proximal and remote sensing data for mineral exploration and resource
445 characterisation.

446 All analytical technologies used for collection of the geoscience data, as well as
447 software packages used for processing the data, are commercially available. However, it should
448 be noted that the HyChips™ system is now superseded by HyLogger3, which collects thermal
449 infrared wavelengths (TIR; 6000 to 14500 nm) in addition to the VNIR-SWIR data. The
450 collection of the TIR wavelength range enables the characterisation of major rock forming
451 minerals such as quartz, which are of major importance for characterisation of iron ore
452 resources, but were not detectable with HyChips™. The HyLogger3 technology is in operation
453 at the six nodes of the Australian National Virtual Core Library
454 (<https://www.auscope.org.au/nvcl>), which provides online open access to more than 3,500 drill
455 cores from the Australian continent.

456

457 7. DATA AVAILABILITY

458 The supplement related to this article is available online at:
459 <https://doi.org/10.25919/5ed83bf55be6a> (Laukamp, 2020). A 3D model of the Rocklea Dome
460 data set is also available from the Geological Survey of Western Australia:
461 <https://dasc.dmp.wa.gov.au/DASC?productAlias=Rocklea3D>.



462

463 **8. AUTHOR CONTRIBUTIONS**

464 CL, TC and MH contributed equally to the manuscript preparation. CL is the custodian
465 of the Rocklea Dome data set stored on the CSIRO's Data Access Portal.

466

467 **9. COMPETING INTERESTS**

468 The authors declare that they have no conflict of interest.

469

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481

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