

Wind, waves, and surface currents in the Southern Ocean: Observations from the Antarctic Circumnavigation Expedition

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Abstract.

The Southern Ocean has a profound impact on the Earth's climate system. Its strong winds, intense currents, and fierce waves are critical components of the ~~air-sea~~air-sea interface and contribute to absorbing, storing, and releasing heat, moisture, gasses, and momentum. Owing to its remoteness and harsh environment, this region is significantly under sampled, hampering the validation of prediction models and large scale observations from satellite sensors. Here, an unprecedented data set of simultaneous observations of winds, surface currents, and ocean waves is presented, to address the scarcity of in situ observations in the region—<http://dx.doi.org/10.26179/5ed0a30aaf764> (Alberello et al., 2020c), and <http://dx.doi.org/10.26179/5e9d038c396f2> (Derkani et al., 2020). Records were acquired underway during the Antarctic Circumnavigation Expedition (ACE), which went around the Southern Ocean from December 2016 to March 2017 (Austral summer). Observations were obtained with the wave and surface current monitoring system WaMoS-II, which scanned the ocean surface around the vessel using marine radars. Measurements were assessed for quality control and compared against available satellite observations. The data set is the most extensive and comprehensive collection of observations of surface ~~process~~processes for the Southern Ocean and ~~has the potential to support further~~ is intended to underpin improvements of wave prediction models around Antarctica and research of air-sea interaction processes, including gas exchange and dynamics of sea spray aerosol particles. The data set has further potentials to support theoretical and numerical research on lower atmosphere, ~~air-sea~~air-sea interface and upper ocean processes.

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1 Introduction

The Southern Ocean comprises an uninterrupted band of water around Antarctica south of the 60th parallel. More broadly, it refers to the body of water south of the main landmasses of Africa, Australia, and South America, with a northern limit at approximately 40° South (see, for example, Young et al., 2020). This region is dominated by strong westerly winds, the notorious roaring forties, furious fifties and screaming sixties (Lundy, 2010). They fuel the Antarctic Circumpolar Current (the world's largest ocean current, e.g. Park et al., 2019), which mixes warm waters descending from the Atlantic, Indian and Pacific Oceans with northward cold streams from the Antarctic. Above all, intense winds give ~~raise~~ rise to some of the fiercest waves on the planet (e.g. Barbariol et al., 2019; Vichi et al., 2019; Young and Ribal, 2019; Young et al., 2020). Acting as an interface between the lower atmosphere and the upper ocean, waves entrain and release momentum, heat, moisture and gasses through breaking (Melville, 1996; Csanady, 2001; Veron, 2015) and drive ~~air-sea~~ air-sea fluxes (e.g. Humphries et al., 2016; Schmale et al., 2019; Thurnherr et al., 2020). Due to almost unlimited fetches (the distance of open water over which the wind blows), Southern Ocean waves are normally long and fast moving, allowing them to inject turbulent motion throughout the water column down to depths of 100–150 m, i.e. approximately half wavelength, and contributing to ocean mixing (Babanin, 2006; Qiao et al., 2016; Toffoli et al., 2012; Alberello et al., 2019b). The combined effect of the Antarctic Circumpolar Current, which cools and sinks near-surface water, and waves, which regulate fluxes and stir the upper ocean, produces a well mixed layer that extends from about 100 m in the summer months to ~~about~~ approximately 500 m in the winter months (e.g. Dong et al., 2008). This deep mixed layer gives the Southern Ocean capacity to store more heat and gasses than any other latitude band on the planet, making this remote ocean a major driver of the Earth's climate system (see, for example, Dong et al., 2007).

South of the 60th parallel, a strong sea ice seasonal cycle of advance and retreat (Eayrs et al., 2019) forms an integral part of a coupled ~~atmosphere-sea-ice-ocean~~ atmosphere-sea ice-ocean system and influences Southern Ocean dynamics. Sea ice extent around Antarctica impacts albedo, atmospheric and thermohaline circulation, and ocean productivity (Perovich et al., 2008; Massom and Stammerjohn, 2010; Notz, 2012), contributing to the heat balance. Further, it attenuates waves, modulating ~~air-sea~~ air-sea fluxes and mixing (Thomas et al., 2019). In turn, waves (in combination with wind) have a significant feedback to the Antarctic sea ice state, extent, and thickness (e.g. Wadhams, 1986; Bennetts et al., 2017; Alberello et al., 2019a; Vichi et al., 2019; Alberello et al., 2020a).

In situ observations of atmospheric and oceanographic properties are scarce due to the remoteness of the region. Although there have been many expeditions crossing the Southern Ocean (see a general overview in Schmale et al., 2019), measurements have primarily been taken en-route to Antarctic stations, leaving entire sectors undersampled. Further, measurements normally concentrate on the lower atmosphere and/or the upper ocean (not necessarily concomitantly), while waves are generally not monitored. Only a handful of buoys have operated in the region: (i) the Southern Ocean Flux Station (Schulz et al., 2011, 2012), a meteorological buoy first deployed in 2010 at approximately 350 nautical miles south-west of Tasmania (Australia) that provides observations of meteorological parameters, including the directional wave spectrum, downwelling radiation, and sea water temperature and salinity; (ii) the Southern Ocean wave buoy network, which comprises of one directional wave buoy deployed south of Campbell Island (New Zealand) and five drifting buoys (Barbariol et al., 2019); and (iii) the Global Southern

Ocean Array and the Global Argentine Basin Array, which are networks of fixed and moored platforms and mobile profilers (gliders) deployed South-West of Chile and in the Argentinian basin, respectively, to monitor waves, ~~air-sea~~ ~~air-sea~~ fluxes of heat, moisture and momentum, and physical, biological and chemical properties throughout the water column (Trowbridge et al., 2019). Buoys have also been deployed in the Antarctic marginal ice zone to monitor waves-in-ice and sea ice drift (e.g. Meiners et al., 2016; Ackley et al., 2020; Meylan et al., 2014; Vichi et al., 2019; Alberello et al., 2020a).

A database covering the Southern Ocean more uniformly is provided by polar orbiting microwave radar satellites such as Altimeters, Scatterometers, and Synthetic Aperture Radar (SAR). Nevertheless, sea state observations are scattered in both space and time due to the nature of polar satellite orbits, and normally limited to average wind and current speeds, and wave heights. SAR technology provides images that can be converted into directional wave energy spectra ([Khan et al., 2020a](#)). However, SAR only detects swell systems, i.e. long-wave systems no longer under the effect of local winds, and with wavelengths longer than 115 m (Collard et al., 2009). It does not resolve the ~~most-relevant~~ wind sea, i.e. the short-wave components directly generated by local winds. This limitation is partly addressed by the recently launched Chinese French Oceanography Satellite (CFOSAT) mission, which detects wave systems with wavelengths longer than 70 m (~~Hauser et al., 2020~~) ([Hauser et al., 2020](#); [Aouf et al., 2020](#)).

The scarcity of in situ observations has a negative feedback on the satellite network, which cannot rely on sufficient ground truth to be validated with high confidence. In turn, this drawback impacts prediction models, which are impaired by notable biases in the Southern Ocean (see, for example, Yuan, 2004; Li et al., 2013; Zieger et al., 2015). To address the lack of in situ observations and support calibration and validation of satellite sensors and prediction models, an international initiative, organised by the Swiss Polar Institute, led an unprecedented circumnavigation of the Antarctic continent (the Antarctic Circumnavigation Expedition, ACE; Walton and Thomas, 2018) during the Austral summer of 2016–2017. The objective of the expedition was to sample concomitant processes in the lower atmosphere, at the ocean surface and in the upper ocean layers all around the Antarctic continent in a single season (Rodríguez-Ros et al., 2020; Schmale et al., 2019; Smart et al., 2020; Thurnherr et al., 2020). Here we present ~~the a~~ database of underway sea state observations, comprising of concurrent records of winds, surface currents and waves. In Sections 2 and 3, details of the expedition, the instrumentation, and its calibration are presented. An overview of the database is presented in Section 4. A comparison against available satellite data and an assessment of uncertainties is given in Section 5. Concluding remarks are made in the last section.

2 The Antarctic Circumnavigation Expedition

ACE took place from 20 December 2016 to 19 March 2017, encompassing the entire Austral summer. It consisted of a voyage around the Southern Ocean between 34° and 74° South aboard the Russian research icebreaker Akademik Tryoshnikov (see technical details in Walton and Thomas, 2018).

The voyage was divided into three legs. ~~The first one (leg Leg 1)~~ was along the Indian Ocean from Cape Town, South Africa, to Hobart, Australia, with stops at Marion Island, Iles de Crozet et Kerguelen and Heard Island. Leg 2 went across the Pacific Ocean to Punta Arenas, Chile, with stations at Mertz glacier, Balleny Islands, Scott Island, Mount Siple (the southernmost station), Peter I Island and Diego Ramirez. ~~The final one (leg Leg 3)~~ crossed the Atlantic Ocean back to Cape Town via South

85 Georgia, South Sandwich Islands and Bouvetøya. In addition, scientific observations were carried out during transit across the Atlantic on the way to/from South Africa (Legs 0 and 4, respectively).

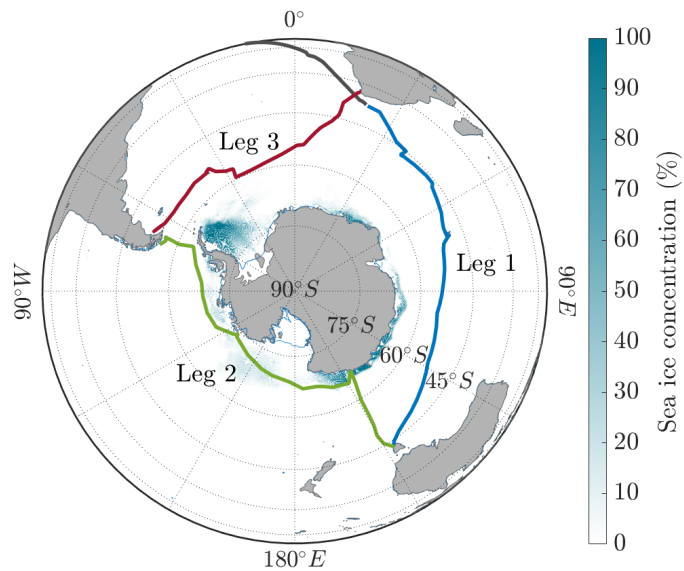


Figure 1. Map of the ACE voyage divided by legs. Average sea ice concentration during the expedition is also shown.

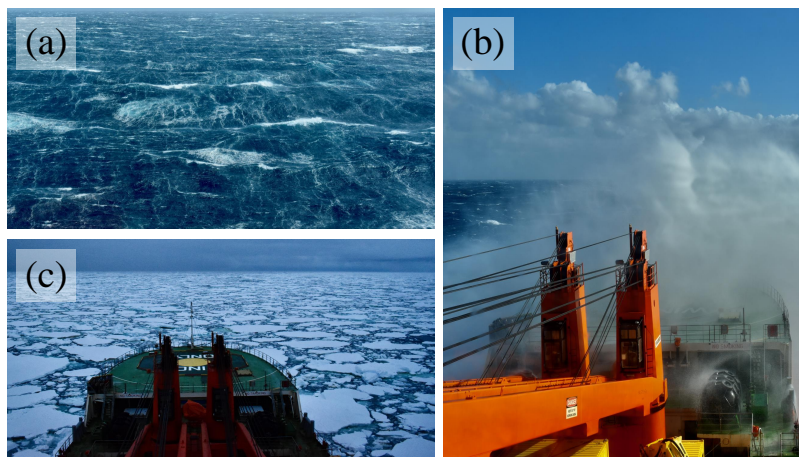


Figure 2. Examples of sea states conditions: ocean surface during storm conditions (a); sailing through a storm (b); the marginal ice zone (c).

A schematic of the expedition is presented in Fig. 1 and photos of the environmental conditions [are reported](#) in Fig. 2. Legs 1 and 3 mostly covered the open ocean north of the 60th parallel, roughly between the sub-Antarctic and Polar fronts

delimiting the Antarctic Circumpolar Current (Figs. 2a and b). Leg 2 primarily concentrated on the Antarctic marginal ice zone (see Fig. 2c) south of the 60th parallel, with two transects across the Western Pacific Ocean sector south of Tasmania and the Drake Passage at the beginning and at the end of the leg. Average sea ice extent during ACE as detected by the Advanced Microwave Scanning Radiometer 2 sensor ([AMSR2](https://seaice.uni-bremen.de/sea-ice-concentration/amsre-amsr2/)—<https://seaice.uni-bremen.de/sea-ice-concentration/amsre-amsr2/>; [Spren et al., 2008](#)) is shown in Fig. 1.

3 The sea state monitoring system

95 3.1 Instrumentation and technical configuration

Sea state observations were recorded with the wave and surface current monitoring system WaMoS-II (details on software, hardware and measurement principles can be found in Reichert et al., 1999). The instrument uses the marine X-band radar (9.41 GHz)—a standard equipment on any vessels—to acquire high-definition radar images of the surrounding ocean surface and derive the directional wave energy spectrum, related integral parameters such as the significant wave height and mean wave period, and surface current speed and direction. [Performance of WaMoS-II and its limitations are discussed in Hessner et al. \(2002, 2008, 2019\); Lund et al. \(2015a, b\). A summary of the range and accuracy of measured parameters are reported in Appendix A.](#)

The overall system consists of an A/D converter, a PC, and a processing software connected to the X-band radar (a schematic of WaMoS-II is presented in Fig. 3a). ~~In addition, true wind velocity and direction from the ship's meteorological station, water depth from the echo-sounder and ship's positions, speed, and course from a Global Positioning System (GPS) receiver are fed into the system.~~ The basic configuration for the X-band radar requires an antenna with rotation speed of 24 rpm, horizontal opening angle of 0.9°, and radar pulse width of 80 ns. In addition, the radar has to be operated in the near range, i.e. 1.5 nautical miles (≈ 2.8 km). This allows WaMoS-II to acquire a radar image with a spatial resolution of 12 m and an angular resolution of 0.9° for every radar rotation (a sample image is reported in Figure 3b). [Further, water depth from the echo-sounder, ship's positions, speed, and course from a Global Positioning System \(GPS\) receiver and true wind velocity and direction from two two-dimensional sonic anemometers operating as part of an automated weather station \(AWS\) and mounted at 31.5 m above mean sea level \(see Schmale et al., 2019; Landwehr et al., 2020a; Thurnherr et al., 2020\) are fed into the system. Wind measurements were acquired at a rate of 1 Hz, averaged over 175 s and converted from the measurement height to a neutral 10-metre wind speed \(\$U_{10}\$ \) by assuming a logarithmic profile \(see Holthuijsen, 2007\) before being passed on to the WaMoS-II.](#)

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3.2 Measurement principles

The marine radar forms images of the surrounding area based on the backscatter of radar beams. The short wavelets (ripples) on the ocean surface contribute notably to reflection, while long wave components modulate the returning signal. This results in stripe-like patterns in the radar images that correspond to the waves ([these patterns are known as sea clutters](#)). The system

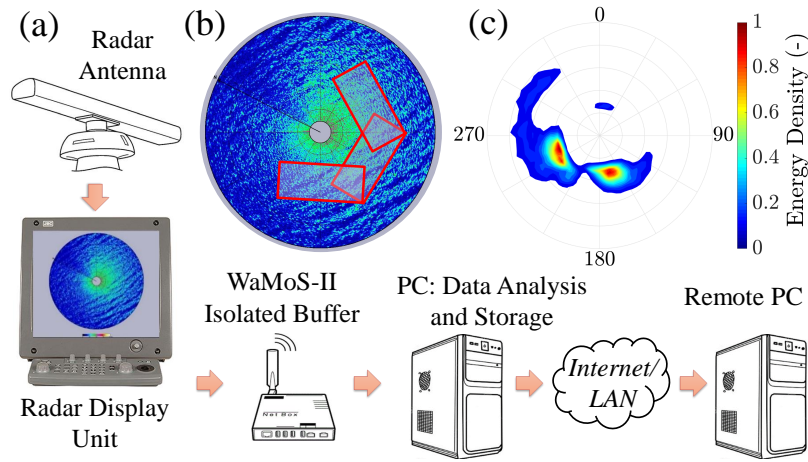


Figure 3. Schematic of the wave and surface current monitoring system WaMoS-II (a); example of radar imagery and sub-areas (not in scale) for post-processing (b); and a sample derived directional wave energy spectrum (c).

120 can ~~only operate with wind speeds~~ detect sea clutters reliably only if wind speed is greater than 3 m/s, ~~which ensure~~ s^{-1} , ~~which ensures~~ the ocean surface is rough enough (i.e. ripples are developed) to backscatter the signal efficiently (Hatten et al., 1998). ~~A small portion of the observations during ACE (approximately 9%) were taken during low wind speed and hence removed from the data set.~~

The basic input for extracting sea state features is a sequence of 64 consecutive images, ~~which correspond to a time period~~ 125 ~~of 175 s (one complete image is acquired for every full turn of the antenna).~~ Post processing is carried out on sub-areas of $600\text{ m} \times 1200\text{ m}$ normally taken in front of the vessel, at port and at starboard to avoid contamination due to the ship's wake (an example of sub-areas is presented in Fig. 3b). The temporal sequence for each sub-area, $I(x, y, t)I_s(x, y, t)$, is transformed with a ~~three dimensional~~ Discrete Fourier Transform into its spectral domain counterpart, i.e. a three-dimensional image spectrum $I^{(3)}(k_x, k_y, \omega)$, where $\vec{k} = (k_x, k_y)$ is the ~~two-dimensional~~ ~~two-dimensional~~ wavenumber vector and ω is the angular 130 frequency. Assuming linear wave theory (Holthuijsen, 2007), spectral components in $I^{(3)}(k_x, k_y, \omega)$ that correspond to ocean waves have to satisfy the linear dispersion relation

$$\omega = \sqrt{g|\vec{k}|\tanh(|\vec{k}|d) + \vec{k}\vec{u}}, \quad (1)$$

where g is the acceleration due to gravity, d is the water depth, \vec{u} is the surface current and $|\vec{k}| = \sqrt{k_x^2 + k_y^2}$ is the ~~two dimensional~~-wavenumber. Spectral components that do not obey equation (1) are assumed to be noise and hence removed. 135 The remaining ~~(filtered)~~ three-dimensional image spectrum is integrated over the positive frequency domain ~~and corrected to remove shadowing and tilting effects, resulting in the directional wave energy spectrum $E_r(k_x, k_y)$ to obtain a wavenumber image spectrum $I(k_x, k_y)$.~~ The latter is expressed as a function of wave frequency $f = \omega/2\pi$ and direction ϑ in WaMoS-II's standard output; ~~the change of variable from wavenumbers to frequency-directions satisfies the dispersion relation in equation (1). An ensemble average over the sub-areas is computed.~~

140 ~~The primary wave spectrum $E_r(f, \vartheta)$ is not scaled correctly, however, does not coincide with the wave energy spectrum,~~
because it ~~depends on~~ represents the intensity of the radar backscatter rather than the amplitude of the water surface elevation
(Nieto Borge et al., 1999; Hessner et al., 2002). Therefore, its zero-th order moment ~~represents the (m_0)~~ represents a signal-
to-noise-ratio (SNR) ~~;~~ instead of the ~~variance of the surface elevation (m_0) .~~ ~~To re-scale the wave spectrum,~~ significant wave
145 height, i.e. a measure of average wave height that is defined as $H_s = 4\sqrt{m_0}$. Consequently, the image spectrum requires a
re-scaling to convert ~~SNR is converted into m_0 with a~~ into the corresponding wave height. This is achieved with the linear
regression equation ~~(see Young et al., 1985, for details);~~ (see Nieto Borge et al., 1999, 2004)

$$\underline{m_0} H_s = A + B\sqrt{SNR}, \quad (2)$$

where A and B are empirical constants that have to be calibrated following installation. ~~Energy at each spectra modes is~~
~~corrected accordingly~~ Re-scaling m_0 enables correction of energy at each spectral mode and derivation of the wave energy
150 spectrum $E_r(k_x, k_y)$. Radar imaging effects like tilt modulation, which refers to changes of the effective incidence angle along
the long wave slope, and shadowing, which is caused by the highest waves in the image, contribute to an inaccurate form
of the resulting spectral density function, shifting energy towards high wavenumbers (Nieto Borge et al., 2004). These effects
depend on the view geometry (height and range of the antenna). Consequently, tilting and shadowing can be assumed to be
homogeneous in the relatively small sub-areas used for post processing and can be minimised with a single modulational
155 transfer function (MTF, Nieto Borge et al., 2004). As the imaging effects depend on the wavelength, the MTF is a function of
the wavenumber that corrects the spectral density at each mode. An ensemble average over all sub-areas is computed to derive
the final wave ~~energy spectrum~~ spectrum $E(k_x, k_y)$ from the input 64 images.

In its standard output format, WaMoS-II archives the wave spectrum as a function of wave frequency, $f = \omega/2\pi$, and
direction, ϑ ~~$E(f, \vartheta)$~~ (see an example; the change of variables from wavenumbers to frequency-direction satisfies the dispersion
160 relation in equation (1). An example of directional wave spectrum is shown in Fig. 3c). The resolution of the wave energy
spectrum is dictated by the size of the sub-areas, which are used to derive the wavenumber counterpart in the first instance,
and not by the temporal window. Considering the resolution of the image (12 m) and the minimum dimension of the sub-area
(600 m), WaMoS-II can detect wavelengths between 15 m and 600 m, which correspond to wave periods from 3 s to ≈ 16 s.

~~Wave-Specific wave~~ parameters are derived by integrating $E(f, \vartheta)$. These include the significant wave height, dominant
165 and mean wave periods, associated wavelengths, directional width and mean wave direction (see Appendix A for a full list of
parameters and their definitions). WaMoS-II also partitions the directional wave energy spectrum to derive wave heights and
periods for wind sea and the first three swell systems. The partitioning of the wave spectrum is performed using the “path of
steepest ascent” technique (Hanson and Phillips, 2001), which is a specific implementation of the inverse catchment scheme
introduced by Hasselmann et al. (1996). The spectral peak that satisfies the condition

$$1.2 \frac{U}{c_p} \cos(\theta - \psi) > 1, \quad (3)$$

where U is the wind speed, c_p is the phase velocity, θ is the wave direction and ψ is the wind direction, is assumed to be associated to the wind sea. All other systems are swell and are ranked based on their energy contents as primary, secondary and tertiary swell.

175 Ocean currents induce a Doppler shift to the wave field. Both current speed and direction can be quantified by minimising the distance between the position of the spectral energy in $I^{(3)}(k_x, k_y, \omega)$ and the theoretical position given by equation (1) using least squares techniques (Young et al., 1985). The vessel's forward speed and heading are used to derive the true current.

Rain, snow and sea ice produce an excess signal backscatter, which ~~affects the quality of the resulting radar image and thus accuracy results in low quality images and consequently inaccurate post processing products.~~ WaMoS-II automatically assesses the reliability of images ~~with an Internal Quality index (IQi)(Hessner et al., 2019). Low quality images are automatically~~
180 ~~excluded from post processing through an internal quality control protocol (see Hessner et al., 2019), which evaluate backscatter intensity, number of sea clutters and stability of ship motion among other parameters (we remark that tilting and shadowing effects are compensated independently using the MTF and do not contribute to quality control). Images that are deemed of low quality are excluded. The majority of low quality images were acquired in the marginal ice zone (i.e. south of the 60th parallel) during Leg 2. As a consequence, observations of waves-in-ice are not available in the present database.~~

185 3.3 Underway observations and file types

WaMoS-II operated continuously to record observations of the sea state during ACE. The vessel was equipped with one X-band radar, which was shared between science (requiring short range settings) and navigational aid (operating at medium and long range). Therefore, data acquisition was interrupted anytime the radar was needed for navigation, resulting in gaps in the observations. This was most common during ~~leg~~ Leg 2, as the radar was often switched to long range to detect icebergs.

190 The wave spectrum $E(f, \vartheta)$ was sampled at ~~160~~ 175 s, assuming no gaps or corrupted images occur during the sampling of 64 consecutive radar images. Output files consist of: (i) the directional wave-energy spectrum in the wavenumber domain ($E(k_x, k_y)$, file extension D2S) and frequency-directional domain ($E(f, \vartheta)$, file extension FTH); and (ii) the (single) one-dimensional frequency-energy spectrum $S(f)$ obtained by integrating the directional spectrum $E(f, \vartheta)$ over ϑ (file extension D1S). Each file also includes a header that provides metadata such as geographical references (latitude and longitudes), time,
195 wind speed and direction, ship speed and heading, current speed and direction, plus additional integrated parameters.

WaMos-II also performed a running average over 20 minutes to minimize the effects of natural variability. Output files consist of: (i) the mean directional wave-energy spectrum in both wavenumber and frequency-directional domain (file extensions D2M and FTM, respectively); and (ii) the mean one-dimensional frequency-energy spectrum (file extension D1M). These files are sampled every ~~160~~ 175 s, with the first one 20 minutes after starting the equipment.

200 In addition, time series of wind, current and wave variables from mean directional wave spectra are archived in monthly summary files.

3.4 Calibration

The calibration of coefficients A and B in equation (2) was performed by forcing the ~~signal-to-noise-ratio- SNR~~ to match independent ~~(benchmark-) observations of m_0~~ . The benchmark ~~observations of H_s~~ . The reference values were reconstructed from records of ship motion, which were measured throughout the expedition with an Inertial Measurement Unit (IMU) at a sampling rate of 1 Hz ~~(Alberello et al., 2020b)~~ (Alberello et al., 2020b; Landwehr et al., 2020b).

An overview of ship motion to sea state conversion ~~(the wave-buoy-analogy)~~ can be found in e.g. ~~Milne and Zed (2018)~~ Nielsen (2017). The method relies on the principle that the vessel is a rigid body with six degrees of freedom (three translations: heave, surge and sway; and three rotations: pitch, roll and yaw) that moves in response to the incident wave field expressed as the frequency spectrum $S(f) = \int E(f, \vartheta) d\vartheta$ and restoring forces ~~expresssed~~ as a function of its mass, geometry, loading conditions, and forward speed, among other parameters (Newman, 2018). The relation between the ship motion and the wave field is evaluated via the response amplitude operator ~~$R(f)$~~ ($R(f)$, see Newman, 2018), i.e. a ship-specific function that translates the motion spectrum $S_{ship}(f)$ into the wave spectrum: $S(f) = S_{ship}(f)/R(f)^2$. Motion spectra were evaluated by applying a ~~Discrete~~ Fourier Transform to ~~160 s time series-5 minute long time series of heave motion~~. An approximation of $R(f)$ for the Akademik Tryoshnikov was calculated solving the equation of motion with a model based on ~~a the~~ boundary element method (NEMOH, Babarit and Delhommeau, 2015) ~~. The variance of the surface elevation is then calculated from the derived $S(f)$ as $m_0 = \int S(f) df$ and taking into account the ship's heading, forward speed and loading conditions. The model is based on a linear approach and, thus, nonlinearities were excluded. The significant wave height was validated against freely available satellite altimeter data (Ribal and Young, 2019); the scatter plot of satellite observations versus reconstructed H_s is presented in Fig. B1 of Appendix B.~~

Coefficients A and B in equation (2) were estimated using a maximum likelihood method for the period 9–11 December 2016 (~~leg-Leg 0~~). The root mean square error ($RMSE$) of the fit is 0.21 m with correlation coefficient $R \approx 0.90$, and scatter index $SI \approx 0.1$. Time series of ~~m_0~~ H_s derived from the IMU sensor and calibrated ~~m_0~~ H_s from WaMoS-II are shown in Fig. 4. Calibrated A and B coefficients were subsequently used to re-scale individual modes of the energy spectrum.

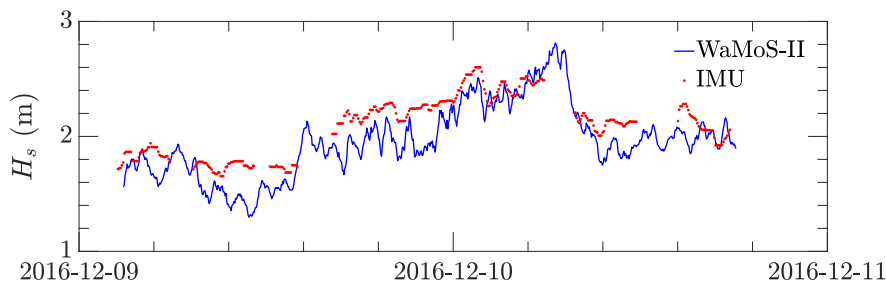


Figure 4. Time series of ~~spectral variance- m_0~~ significant wave height H_s : benchmark observations derived from ship motion data (red dots) and calibrated records from WaMoS-II (blue solid line).

4.1 Sea state climate during ACE

Excluding the regions south of the 60th parallel, which ~~undergoes~~undergo a strong seasonal sea ice cycle (Eayrs et al., 2019), the Southern Ocean is normally characterised by weak seasonal variability (Young et al., 2020). Therefore, extreme sea states remain likely even during summer. As a reference, wind, current and wave climate statistics in the form of the 50th and 90th percentiles (hereafter P50 and P90, respectively) aggregated in $2^\circ \times 2^\circ$ regions for the summer months (December, January, February) are reported in Fig. 5. Data of wind speed and wave height are from all satellite missions mounting altimeter sensors that are available from 1985 to 2019 ~~for wind speed and wave height (Ribal and Young, 2019) and (Ribal and Young, 2019).~~ Data of current speed are from the COPERNICUS-GLOBCURRENT database—<https://marine.copernicus.eu>—that combines the velocity field of geostrophic surface currents from satellite sensors recorded from 1993 to 2019 for surface current speed (Rio et al., 2014)(Rio et al., 2014) and modelled Ekman currents, which include components from wind stress forcing obtained from atmospheric system and drifters data.

The wind speed is represented by its value at 10-metre above sea level (U_{10}). Apart from a region ~~East~~east of Argentina, where the South American continent induces a shadowing effect, wind speed is fairly uniform throughout the ocean: P50 varies between 10 and 12 m/s s^{-1} , while P90 ranges between 15 and 18 m/s s^{-1} (see Figs. 5a and b). Close to the Antarctic continent and outside the belt of the strong westerly winds (south of the 70th parallel), wind speed weakens with P50 reducing to $\approx 3 \text{ m/s s}^{-1}$ and P90 to $\approx 10 \text{ m/s s}^{-1}$. There are also low wind speeds ($U_{10} < 3 \text{ m/s s}^{-1}$ for both P50 and P90) in the lee of the Antarctic ~~peninsula~~Peninsula, although this may relate to uncertainties ~~related~~due to a high concentration of sea ice (see Fig. 1) and/or the increased drag over sea ice compared to open water (Martinson and Wamser, 1990). Note that, excluding the station at Mount Siple, the expedition remained within the belt of westerly winds.

Significant wave height H_s follows the wind pattern, underpinning the dominance of wind seas on swell systems. Between 40° and 60° South, the belt where most of ~~legs~~Legs 1 and 3 took place, ~~the~~P50 ≈ 3.5 m and P90 ranges between 5 and 6 m (Figs. 5c and d). There is an evident shadowing effect east of the Drake passage, due to a combination of lower wind speed and a reduction of fetches. South of the 60th parallel (Leg 2), H_s drops notably with the P50 decreasing to ≈ 2 m and P90 reducing to ≈ 4 m, despite strong westerly winds being active down to 70° South. The attenuation is induced by sea ice (Bennetts et al., 2015; Toffoli et al., 2015; Montiel et al., 2016), which has high concentration close to the Antarctic coastline, even in the summer months, especially in the West Pacific, Ross sea and Amundsen sea sectors of Antarctica (see Fig. 1).

The speed of the Antarctic Circumpolar Current has P50 $\approx 0.5 \text{ m/s s}^{-1}$. The P90 shows velocity with maxima in excess of 0.75 m/s s^{-1} , especially in the Indian Ocean sector (first half of ~~leg~~Leg 1). Besides the Antarctic Circumpolar Current, ACE crossed two regions with strong currents: the Agulhas region, where P90 excess 1.5 m/s s^{-1} ; and the Argentine basin (beginning of ~~leg~~Leg 3), where a northward extension of the Antarctic Circumpolar Current forces water flow to speeds of approximately 1 m/s s^{-1} . South of the polar front (~~leg-2~~latitudes higher than 60° South), there ~~are~~is no significant circulation ~~patter~~pattern, with maximum current speed less than 0.3 m/s s^{-1} (see Fig. 5 f).

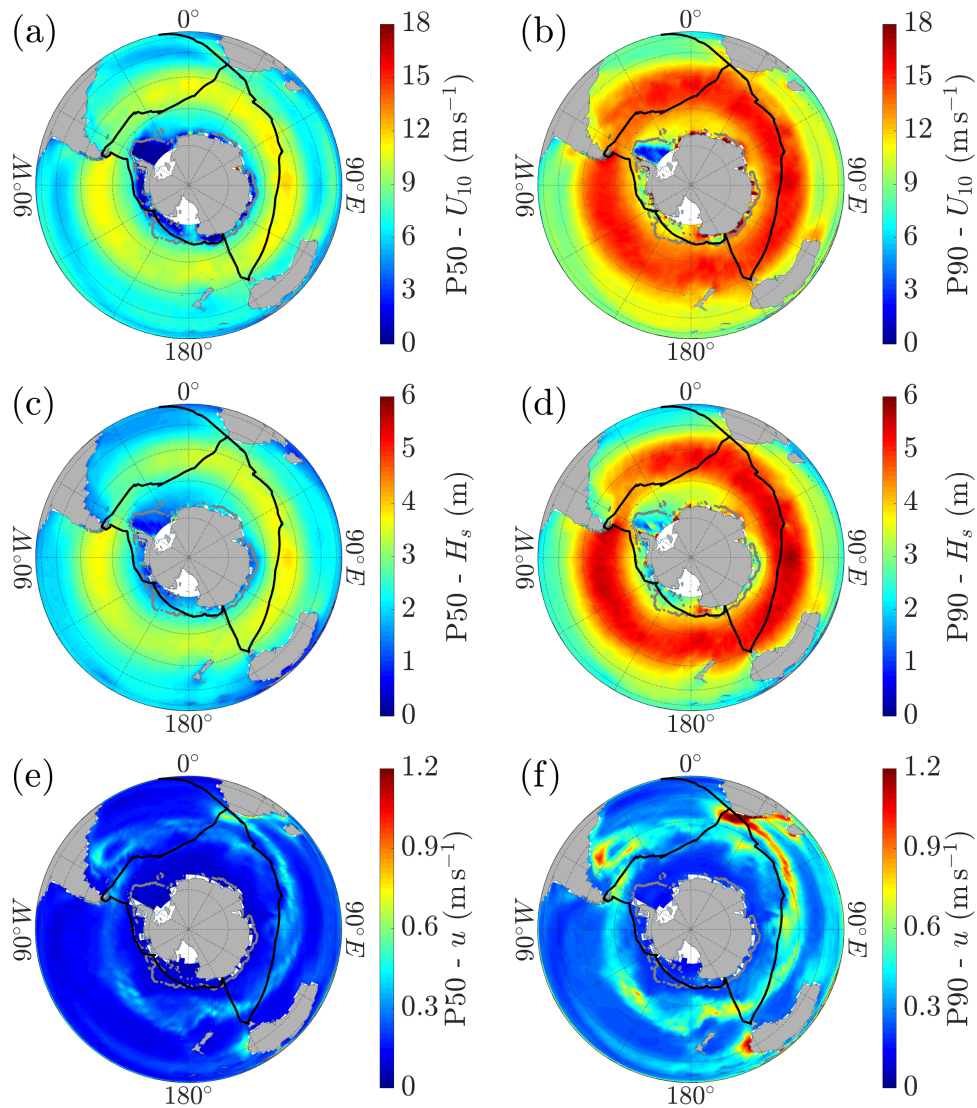


Figure 5. Wind speed (U_{10}), significant wave height (H_s), and surface current speed (u) climatology in austral-Austral summer: (a) 50th percentile (median) wind speed, (b) 90th percentile wind speed, (c) 50th percentile (median) significant wave height, (d) 90th percentile significant wave height, (e) 50th percentile (median) surface current speed, (f) 90th percentile surface current speed. Latitudes are shown every 15° (from 15°S to 90°S) by thin lines; the route of the ACE voyage is reported as a black solid line); and the sea ice edge, defined by the 10% sea ice concentration, is shown as a grey solid line.

4.2 Observed sea states during ACE

Figure ?? Figures 6 shows time series of the observed sea-state conditions—10for the entire expedition of 10 m true wind speed (U_{10}), significant wave height (H_s), and current speed (u). As a benchmark, collocated values of P50, their interquartile range

260

(IQR) and P90 are reported in the figure. Time series of mean wave period (T_{m-10}), wave directional spreading (σ_θ), inverse wave age (μ), and wave steepness (ε) and current speed (u)—throughout the expedition are presented in Figure 7. Definitions of the variables are reported in Appendix A. As a benchmark, collocated values of P50 and P90 for wind speeds, significant wave heights and current speeds, are also reported in figure ??.

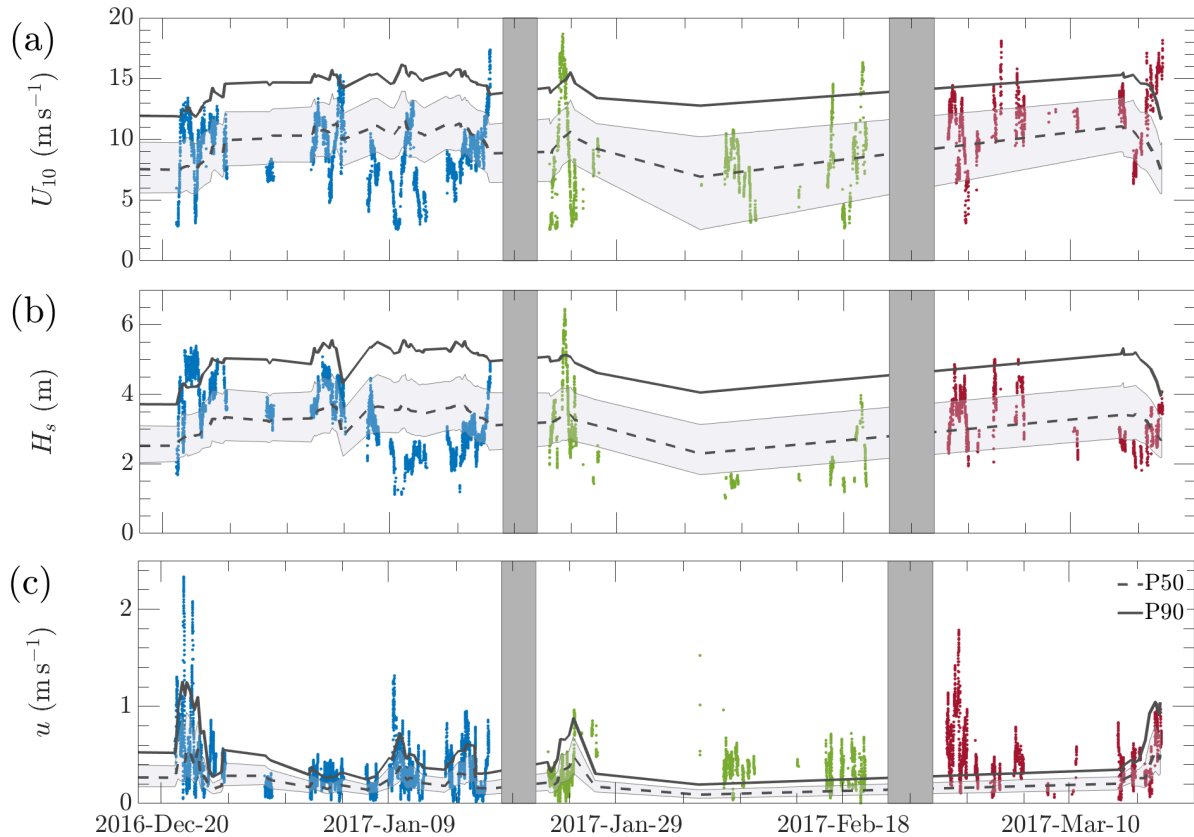


Figure 6. Time series of sea state variables in **leg-Leg 1** (blue), **leg-Leg 2** (green), and **leg-Leg 3** (red): **wind speed** (a) **wind speed from the automated weather station**; **significant wave height** (b) **;** **mean wave period** (c); **mean directional spread** (d); **inverse significant wave age** (e) **height**; **wave steepness** and (f) **;** and **current speed** (g). Details of variables are reported in Appendix A. Climate statistics derived from satellite observation for wind speed For each variable, significant wave height and current speed are shown: 50th (the dashed line) and 90th shading represent 50th (percentile and its interquartile range IQR, respectively, based on climate statistics from satellite observations; the solid line) indicates the 90th percentile.

265 Overall, the observed median wind speed was 7.25 m/s s^{-1} with an interquartile (IQ) of 5.1 m/s s^{-1} . During leg s^{-1} . During Leg 1, the expedition went through six storm events with wind speeds reported in excess of 12 m/s s^{-1} ($P50 \approx 10 \text{ m/s s}^{-1}$). Two of these events were equal to or greater than the P90 for the season ($\approx 15 \text{ m/s s}^{-1}$). Leg 2 started with the most extreme storm during ACE; wind winds reached speeds close to 20 m/s s^{-1} , which is well above P90. The remainder of leg Leg 2 was

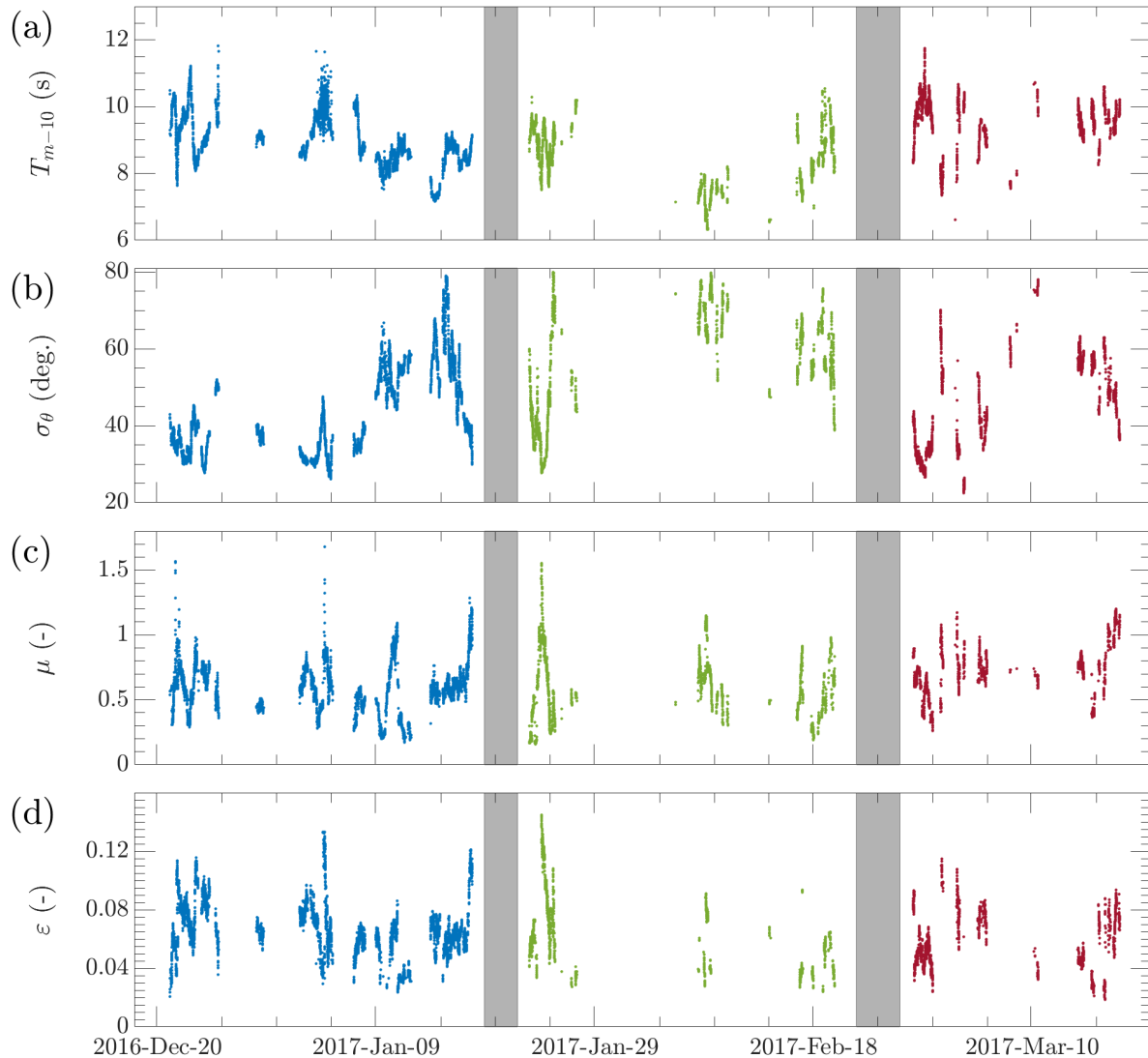


Figure 7. Time series of sea state variables in Leg 1 (blue), Leg 2 (green), and Leg 3 (red): (a) mean wave period; (b) mean directional spread; (c) inverse wave age; (d) wave steepness. Details of variables are reported in Appendix A.

characterised by relatively low wind speeds, consistent with P50. Two more storms with wind speeds in excess of P90 were encountered at the end of ~~leg~~ Leg 2, while approaching and crossing the Drake Passage. The final leg was also characterised by intense storms with wind speeds notably above P50 for almost the entire leg. Three significant storm events with wind speeds above P90 ($U_{10} \approx 18 \text{ m/s}$) were reported.

The median significant wave height during the expedition was 2.61 m and $1Q \approx 1.6$, $3Q \approx 1.6$ m. To avoid the most energetic waves, the ship's course was continuously adapted to bypass storms. Despite this, intense wave conditions were encountered

275 with H_s reaching the P90 (≈ 5 m) during almost all storm events, especially during [leg-Leg 1](#) and [leg-Leg 3](#). The largest waves ($H_s > 6$ m) were encountered at the beginning of [leg-Leg 2](#) (see photos of the sea state in Figs. 2a and b). Thereafter, H_s was less than 2 m as a result of the interaction with sea ice (see Figs. 2c). The crossing of the Drake passage at the end of [leg-Leg 2](#) did not record significantly large waves, with $H_s \approx 4$ m at most. Wave periods were generally long and normally in excess of 8 s ($\approx 100 > 100$ m wavelengths), substantiating the extensive (almost infinite) fetches for wave development. Concomitantly
280 with almost all storms, T_{m-10} increased and reached maximum values of 11–12 s (wavelengths ≈ 200 m).

The majority of Legs 1 and 3 followed the Antarctic Circumpolar Current with records of surface current speeds oscillating around 0.5 m s^{-1} . Interestingly, observations were notably higher than P90 for the majority of the expedition. Despite being primarily south of the polar front, currents faster than P90 were also recorded in Leg 2, primarily in the marginal ice zone. ACE crossed two regions characterised by strong surface currents: East of South Africa at the southernmost edge of the
285 Agulhas current (beginning of Leg 1), with speeds up to 2 m s^{-1} ; and East of South America where the Antarctic Circumpolar Current has a northward extension, with surface speeds recorded up to 1.8 m s^{-1} . Observations at both locations exceeded P90 notably. We remark that P50 and P90 include the contribution of geostrophic surface currents and wind stresses. However, additional components such as inertial oscillations (Treguier and Klein 1994) are not taken into account due to the course resolution of satellite observations and Ekman components. To some extent, the absence of inertial oscillations in climate
290 statistics substantiate the significant current speeds recorded by the WaMoS-II. Further, Ekman components remain uncertain in the Southern Ocean due to inaccuracies in estimating wind stress from the atmospheric system, adding inconsistencies to benchmark statistics.

An intrinsic feature of oceanic sea states is the directional distribution of the spectral density function ([Toffoli et al., 2017; Fadaeiazar et al.](#)) which is summarised in the form of a mean directional spreading (i.e. the circular standard deviation of the directional wave energy spectrum). Sea states dominated by strong winds are normally characterised by a broad spreading with $\sigma_\vartheta > 40^\circ$ (Donelan et al., 1985). These conditions were reported consistently throughout the expedition, with maximum mean directional spreading reaching values as large as 80° . Narrow directional distributions ($\sigma_\vartheta \leq 30^\circ$) were also common ~~though~~ and primarily recorded in between storms, where the sea state was dominated by swells.

The inverse wave age μ and the wave steepness ε are parameters that estimate the stage of growth of the wave field. ~~To~~
300 ~~some extent, both~~ Both variables are associated to nonlinear mechanisms that lead to large (extreme) waves (Onorato et al., 2009; Toffoli et al., 2017), and wave-breaking (Toffoli et al., 2010), and, thus, ocean–atmosphere fluxes (Schmale et al., 2019; Thurnherr et al., 2020). The inverse wave age is the ratio of wind speed to wave phase velocity (i.e. the ratio of wavelength to period). Following their generation, waves grow in height and length until they move faster than winds (Holthuijsen, 2007). For $\mu > 0.8$, waves are “young” as they are in a growing phase. This condition is normally characterised by a steep profile,
305 which leads to breaking. Young waves were recorded during all storm events with steepness generally in excess of 0.1. For $\mu < 0.8$, the waves no longer receive energy from wind as they have reached full development. The shape of waves is gently sloping (i.e. the wave steepness is small) and breaking is unlikely (the ocean is dominated by swell). During the most extreme events at the beginning of [leg-Leg 2](#), steepness reached a maximum of about 0.13. This is an exceptionally high value for ocean waves and is normally associated to the formation of rogue waves (Onorato et al., 2009; Toffoli et al., 2010).

310 The majority of legs 1 and 3 followed the Antarctic Circumpolar Current with records of surface current speeds oscillating
around 0.5 m/s. Interestingly, observations were notably higher than P90 for the majority of the expedition. Despite being
primarily south of the polar front, currents faster than P90 were also recorded in leg 2, primarily in the marginal ice zone. ACE
crossed two regions characterised by strong surface currents: East of South Africa at the southernmost edge of the Agulhas
315 current (beginning of leg 1), with speeds up to 2 m/s; east of South America where the Antarctic Circumpolar Current has a
northward extension, with surface speeds recorded up to 1.8 m/s.

5 Comparison against satellite observations

5.1 Wind speed and significant wave height

Wind speeds and significant wave heights are compared against collocated satellite observations from altimeter sensors (same
data source discussed in the previous session). Due to the scattered nature of satellite data, average values are computed for
320 clusters with spatial resolution of $\Delta X = \Delta Y = 0.5^\circ$ and temporal resolution of $\Delta t = 3$ h. Although satellite observations have
been quality controlled and calibrated against available in-situ sensors (see details in Ribal and Young, 2019), the scarcity of
in-situ data in the Southern Ocean leaves uncertainties in the data set.

Fig. 8 shows scatter diagrams of matching averages at collocated clusters (panels a and b) and geographical distribu-
tions of biases (difference between WaMoS-II and satellite observations, panels c and d). Overall, in-situ measurements
325 of wind speeds during ACE are consistent with concurrent satellite observations, with data laying along the 1:1 correla-
tion line (wind observations are also in agreement with reanalysis data from the ERA-5 archive; Landwehr et al., 2020a). Nev-
ertheless, there is a notable $RMSE \approx 3.2$ m/s s^{-1} , with $R \approx 0.70$ and $SI \approx 0.360$. Biases show both over (especially at the
beginning of leg-Leg 1 and leg-Leg 3) and under (at the end of leg-Leg 1 and leg-Leg 2) estimations of satellite observa-
tions, varying between -6 to 6 m/s s^{-1} . The most substantial positive biases are reported in the marginal ice zone, where
330 Antarctic sea ice affects wind speed detection.

Significant wave heights match better with satellite observations than wind speeds: with $RMSE \approx 0.42$ m, $R \approx 0.93$, and
 $SI \approx 0.155$. Most of collocated observations were found in leg-Leg 1. Overall, the bias is positive, indicating a slight under
estimation of the sea state from satellite sensors. The largest biases (ranging between 0.4 and 1.2 m) were linked to storm events
and are $\approx 10\%$ of the in-situ measured values.

335 5.2 Directional wave spectrum

Altimeter sensors only measure specific variables, namely the significant wave height and the wind speed, whereas SAR im-
agery can be converted into a directional wave energy spectrum (e.g. Collard et al., 2009) (e.g. Collard et al., 2009). Collocated
SAR spectra from Sentinel-1A/1B missions within area of $0.5^\circ \times 0.5^\circ$ and maximum temporal difference of 6 h were retrieved
from the Australian Ocean Data Network (AODN) portal (Khan et al., 2020b). Overall, ten SAR spectra were found during
340 ACE, with $\approx 70\%$ of them in the Indian Ocean during leg-Leg 1.

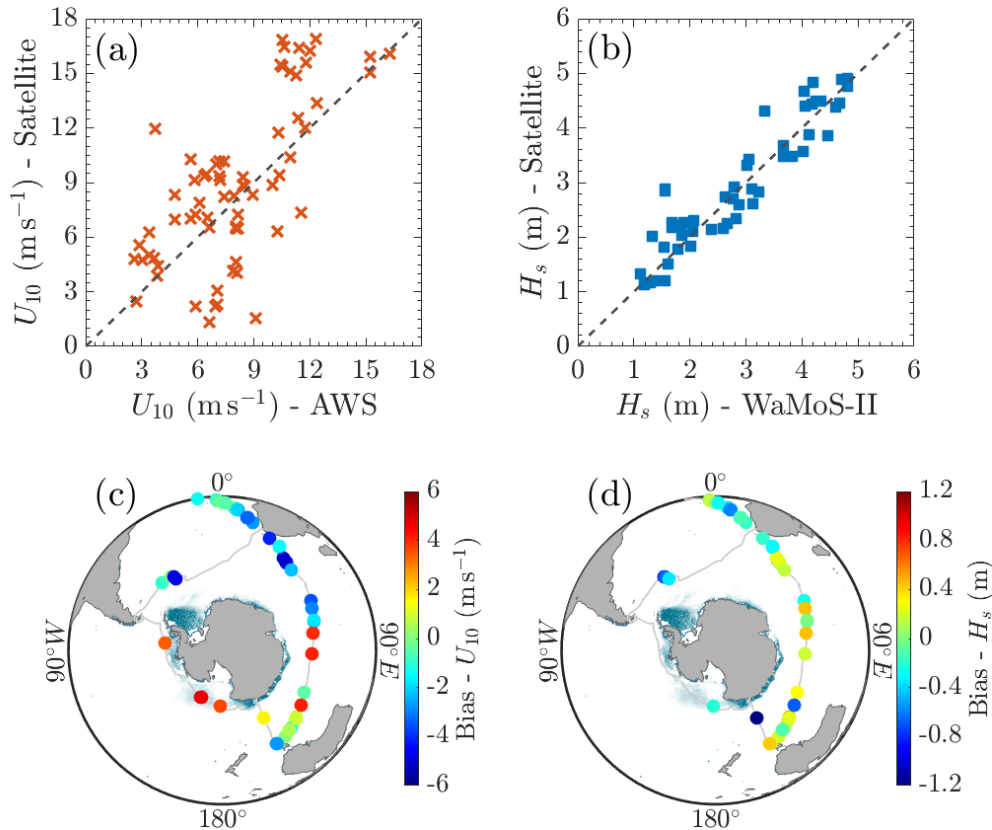


Figure 8. Wind [from the automated weather station](#) and significant wave height from WaMoS-II versus satellite observations: scatter diagrams (panels a and b); and geographical distribution of biases (panels c and d). [Average sea ice concentration during the expedition is overlaid in panels c and d.](#)

~~Example of corresponding~~ [Examples of collocated](#) wave spectra from WaMoS-II and SAR are presented in Fig. 9. ~~The overall spectral shape~~; [mean wind direction is also reported.](#) We remark that SAR detects wavelength longer than 115 m (approximately, wave periods exceeding 8 s or frequencies below 0.1 Hz) and represented swell systems primarily. WaMoS-II, on the contrary, captures the full spectrum, including the short wavelengths of the wind sea. Within the operational range of SAR ($f < 0.1 \text{ Hz}$ in the figure), the spectral shape from both sensors agrees well, ~~particularly especially~~ for the portion around the ~~most energetic peak~~ [primary \(most energetic\) swell](#). Notable discrepancies, however, are evident for less energetic secondary peaks, for which the relative uncertainty grows. [High frequency components \(\$f > 0.1 \text{ Hz}\$ \) are not resolved in SAR, but appear in the WaMoS-II spectra.](#) Note that the misalignment of high frequency components with the wind direction in the upper two panels is due to recent wind change.

To provide a more robust comparison, scatter diagrams for H_s , T_{m-10} and σ_θ are presented in Fig. 10. For consistency, wave spectra from WaMoS-II have been filtered to eliminate high frequency modes that are not detected by SAR ($f > 0.117 \text{ Hz}$ or

wavelength $L < 115$ m). SAR and WaMoS-II observations agree well, with $RMSE \approx 0.36$ m, $R \approx 0.92$, and $SI \approx 0.20$ for H_s ; $RMSE \approx 0.42$ s, $R \approx 0.92$, and $SI \approx 0.038$ for T_{m-10} , noting wave periods from SAR are consistently (slightly) higher than WaMoS-II's; and $RMSE \approx 13.41^\circ$, $R \approx 0.56$, and $SI \approx 0.295$ for σ_θ , despite two outliers.

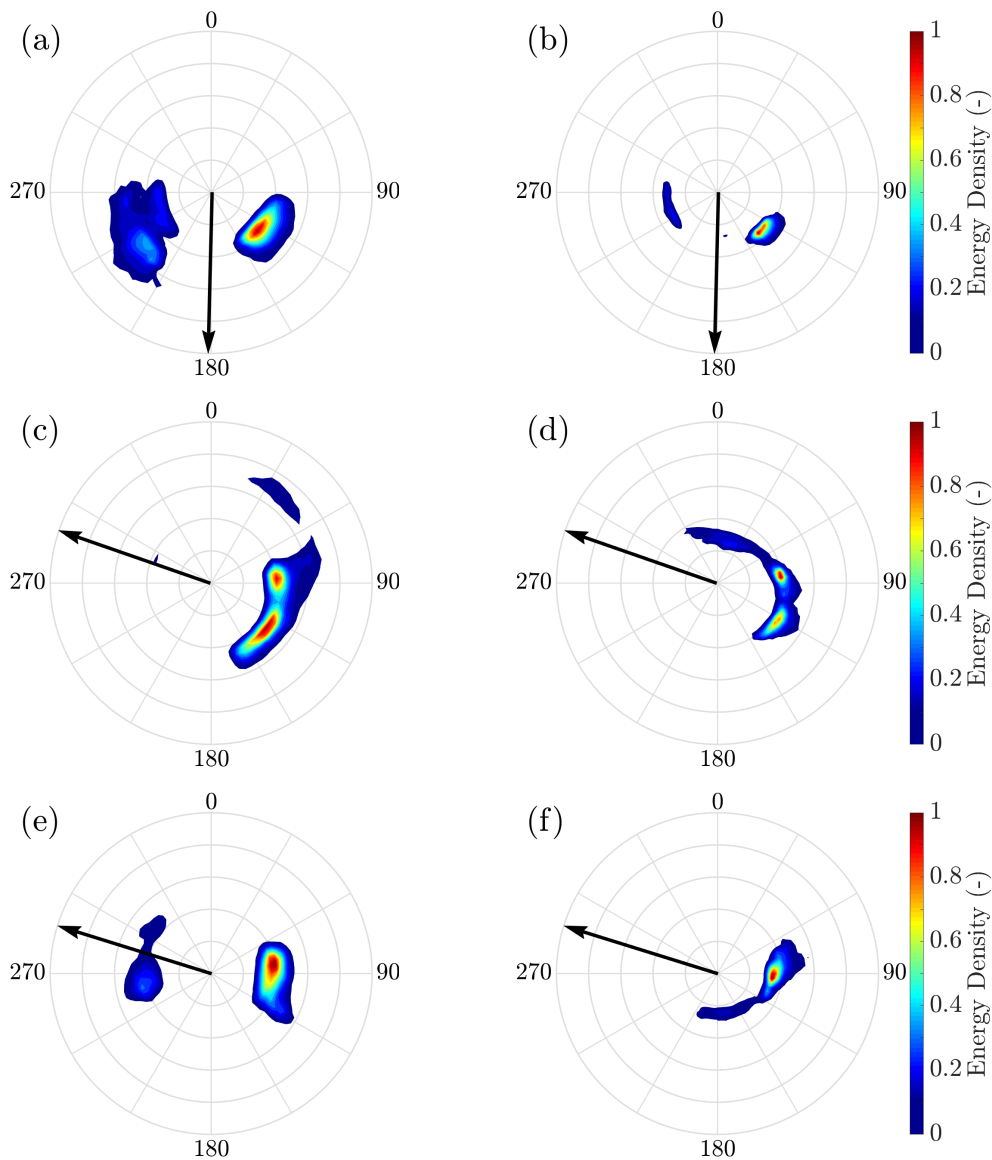


Figure 9. Example directional wave energy spectra recorded during ACE (panels a, c, and e) and collocated SAR spectra (panels b, d, and f). Wind direction recorded during ACE is shown as black arrows. Both the wave spectra and wind direction follow the “coming from” convention. Circles in the polar plot indicates frequencies from 0.05 Hz (innermost) to 0.25 Hz (outermost) with a step of 0.05 Hz; radiant lines indicate direction with a 30° step.

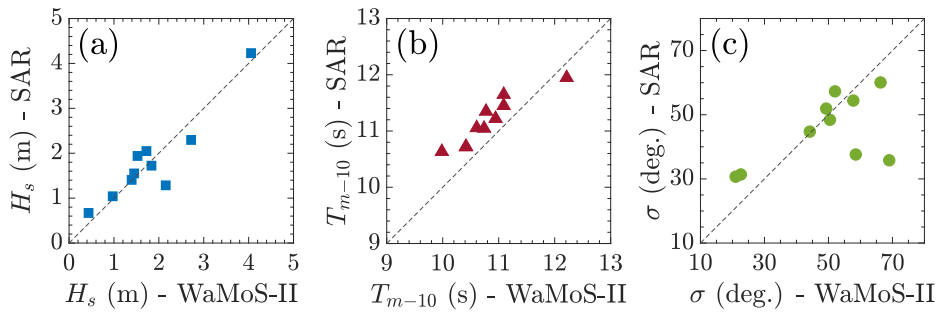


Figure 10. Comparison of integrated parameters (WaMoS-II versus SAR): significant wave height (H_s , panel a); [energy](#) wave period (T_{m-10} , panel b); and wave directional spreading (σ_θ , panel c)

355 5.3 Surface current

[FigureFig. 11](#) shows the scatter diagram comparing surface current speeds from WaMoS-II and collocated observations derived from altimeter sensors (Rio et al., 2014). The geographical distributions of current speeds, directions and concurrent differences between WaMoS-II and altimeter sensors are presented in [FigureFig. 12](#). Note that values in both figures represent averages of observations falling in clusters of $0.5^\circ \times 0.5^\circ$ with temporal resolution of 0.5 h.

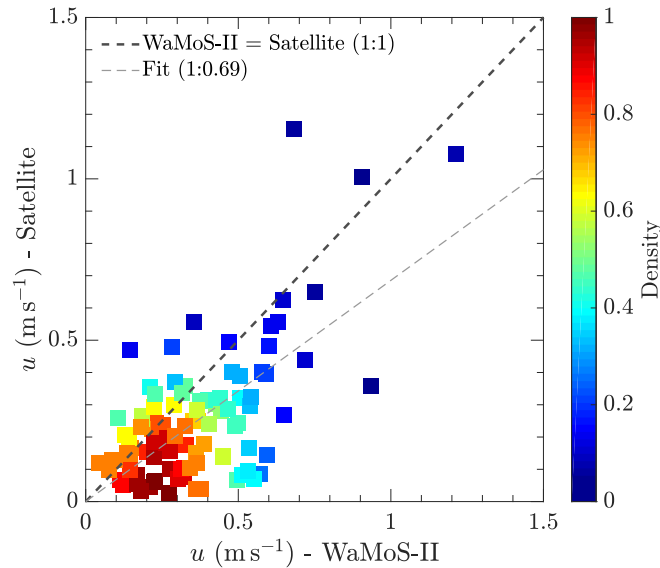


Figure 11. Scatter plot of WaMoS-II surface current speeds against observations derived from satellite sensors.

360 Contrary to wind and wave parameters, current speeds from WaMoS-II show notable differences from satellite observations. The former produces current speeds that are about 30% larger than the latter. Other basic metrics of the scatter diagram

are: $RMSE \approx 0.2 \text{ m/s s}^{-1}$; $R \approx 0.63$; and $SI \approx 0.80$. Bias associated to current speed are uniformly distributed across the expedition. Relatively small bias was detected in [leg-Leg 1](#), when sailing along the Antarctic Circumpolar Current. The largest biases (about 0.5 m/s s^{-1} , the same order of magnitude of the current speed itself) were detected primarily at the beginning of [leg-Leg 1](#) and at the end of [leg-Leg 3](#), where the ship crossed the Agulhas current, and in [leg-Leg 2](#), when crossing the Antarctic marginal ice zone. [The reported differences are linked to inconsistencies between WaMoS-II and benchmark data due to inertial oscillations \(Treguier and Klein, 1994\), which are not detected by satellite observations, and inaccuracy of wind stresses in the Ekman components.](#)

Current direction is generally in better agreement with satellite observations than speed (see Fig. 12a and b). Differences between WaMoS-II altimeter sensors are normally small throughout the expedition with common values of about 10° . The only substantial differences were recorded at the beginning of [leg-Leg 3](#), east of South America.

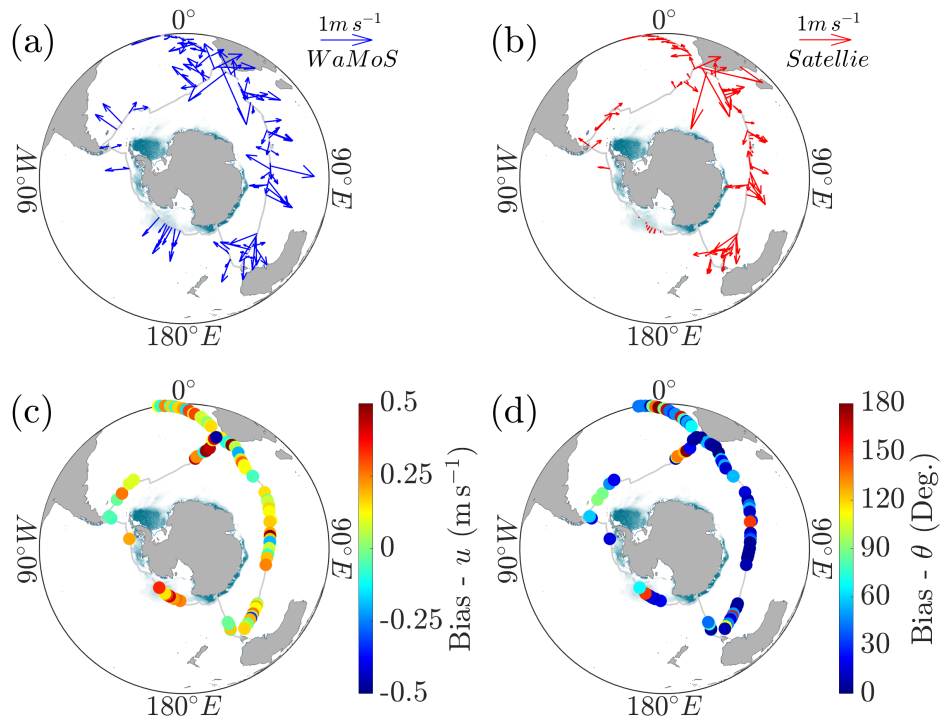


Figure 12. Surface current velocities along the ACE voyage: WaMoS-II (a); and space-borne altimeter sensors (b). Surface current biases between WaMoS-II and satellite observations for current speed (c) and direction (d). [Average sea ice concentration during the expedition is overlaid in all panels.](#)

6 Conclusions

The scarcity of field observations in the Southern Ocean hampers the accuracy of satellite sensors and prediction models. In response to this issue, a unique data set of sea state parameters, comprising concomitant observations of winds, waves and surface currents, was recorded in the Southern Ocean during the Antarctic Circumnavigation Expedition, from December 2016 to March 2017. Measurements were obtained using a radar-based wave and surface current monitoring system (WaMoS-II) and complemented with records of winds from the meteorological station on board of the research icebreaker Akademik Tryoshnikov. Despite some gaps, observations of wind speeds and directions, directional wave energy spectra, integrated parameters such as wave heights, mean wave periods and wavelengths, and current speeds and directions were collected underway during the entire expedition with outputs every 160175 s. The sea state monitoring system was calibrated with benchmark sea state records, which were reconstructed from the ship motion. Measurements were also compared against available observations from satellite-borne sensors to verify the ~~robustness~~ robustness of the database.

The data set includes observations around the Southern Ocean from latitude 34° to 74° South. This ~~includes~~ comprises records in the open ocean across the Antarctic Circumpolar Current and in the Antarctic marginal ice zone. Due to its exposure to strong westerly winds, the Southern Ocean is subjected to harsh sea state conditions all year round. Although the expedition took place during the Austral summer, the data set contains records of severe sea states, in excess of the 90th percentile expected for the season.

The expedition was conceived to bring together a broad range of Earth Science disciplines ~~to explore~~ with the aim of exploring the interplay of processes in the lower atmosphere, ocean surface, subsurface and land ~~;~~ simultaneously with simultaneous observations. Additional data sets of ~~processes affected by sea state such as air-sea fluxes, aerosols, ocean chemistry and micro-fibres~~ atmospheric and oceanic variables that relate to sea state can be found in Schmale et al. (2019); Rodríguez-Ros et al. (2020); Smart et al. (2020); Thurnherr et al. (2020) ~~;~~ and Suaria et al. (2020). ~~The data set presented herein is the most extensive collection of in situ observations of surface processes that is available for~~ These include, but are not limited to, air-sea fluxes (mass, gas, heat and momentum), aerosol concentrations, stable water isotopologues and micro-fibres. Collocated observations have been the foundation for research on moist diabatic processes (Thurnherr et al., 2020) and sea spray aerosols dynamics (Landwehr et al., 2020b), demonstrating capacities for surface waves to modulate water isotopologues concentrations and marine aerosols emissions, settling velocity and lifetime in the marine boundary layer up to the cloud condensation height. The presented database has further potentials to support research enhancing wave model performances in the Southern Ocean. On its own or in combination with other concurrent data sets from the Antarctic Circumnavigation Expedition, it can support further theoretical and numerical research on lower atmosphere, air-sea interface and upper ocean processes, contributing to understanding this remote, and yet fundamental for Earth climate, region Southern Ocean, wave dynamics, including occurrence of rogue waves, wave dissipation mechanisms as well as other coupled processes, including those interconnecting waves with the upper ocean and sea ice in the Antarctic marginal ice zone.

7 Data availability

405 Data are available through the Australian Antarctic Data Centre: (i) *Alberello et al. (2020c)* contains data sets of wave spectra including files D1S, D2S, D1M, D2M and FTH (<http://dx.doi.org/10.26179/5ed0a30aaf764>); and (ii) *Derkani et al. (2020)* contains time series of wind speed and direction, current speed and direction, sea state parameters including wave height, period, wavelength and mean direction for total sea, wind sea and swell systems, ship course, position, and speed for each month of the expedition (<http://dx.doi.org/10.26179/5e9d038c396f2>). Day and time of available measurements can be found
410 in the file *Available_Measurements_List.txt*, which is included in *Alberello et al. (2020c)*.

Appendix A: WaMoS-II sea state parameters

Details of sea state related variables from WaMoS-II output files as well as integrated parameters are described in Table A1. The n^{th} order moment of the spectral density function, m_n , referred to in the table is defined as $m_n = \int \int f^n E(f, \vartheta) df d\vartheta$. Directional Fourier coefficients a and b used to compute the wave directional spreading are as follows:

$$415 \quad a = \int \int \cos(\vartheta) S(f, \vartheta) df d\vartheta, \quad (\text{A1})$$

$$b = \int \int \sin(\vartheta) S(f, \vartheta) df d\vartheta. \quad (\text{A2})$$

Appendix B: Validation of ship motion to sea state conversion

Significant wave heights reconstructed from ship motion data were validated against freely available satellite observations Ribal and Young (2019) for the entire ACE voyage (see scatter plot in Figure B1). Due to the coarse resolution of satellite data,
420 average values are computed for clusters with spatial resolution of $0.5^\circ \times 0.5^\circ$ and temporal resolution of 3 h. Overall, there is a good agreement between reconstructed and observed sea state. The root-mean squared error (*RMSE*) is 0.4 m, the correlation coefficient (*R*) is 0.94, and the scatter index (*SI*) is 0.17. Similar error metrics is obtained by comparing the reconstructed sea state against parameters from the European Centre for Medium-Range Weather Forecasts (ECMWF—<https://www.ecmwf.int/en/forecasts/data>) ERA-5 database.

425 *Author contributions.* KR, KM and AT participated to the expedition and acquired the data. MHD, AA, LB and AT conceived the manuscript. KGH provided on-shore technical support and LA provided marine forecast during the expedition. MHD and KGH calibrated and analysed the data. FN provided benchmark observation for calibration. LA and SSK provided SAR-spectra. All authors contributed to the data interpretation and to the writing of the manuscript.

Table A1. WaMoS-II output and integrated sea state related parameters and their symbol, definition, range, and accuracy.

Sea state related parameter	Symbol	Definition	Range	Accuracy
10m true wind speed and direction (m/s , $^{\circ}$)	U_{10}, α	–	–	–
2-D two-dimensional wave number spectrum (m^4)	$E(k_x, k_y)$	Refer to Sec. (3.2)	–	n/a
2-D frequency-direction spectrum ($m^2/(Hz \times rad)$) two-dimensional frequency-direction spectrum ($m^2 (Hz \times rad)^{-1}$)	$E(f, \vartheta)$	$ \vec{k} \left \frac{\partial}{\partial \omega} \vec{k} \right E(k_x, k_y)$	0.0078-0.5000 Hz, 0-360 $^{\circ}$	n/a
1-D one-dimensional frequency spectrum (m^2/Hz)	$S(f)$	$\int_0^{360^{\circ}} E(f, \vartheta) d\vartheta$	0.0078-0.5000 Hz	n/a
Significant wave height (m)	H_S	Refer to Eq. (2)	1-20 m	$\pm 0.5 m$
Energy wave period (s)	T_{m-10}	$T_{m-10} = \frac{m-1}{m_0}$	3.5-55 s	$\pm 0.5 s$
Peak wave period (s)	T_p	$\frac{1}{f_p}$	3.5-55 s	$\pm 0.5 s$
Mean wave direction ($^{\circ}$)	β_m	$\arctan(b/a)$	0-360 $^{\circ}$	$\pm 2^{\circ}$
Peak wave direction ($^{\circ}$)	β_p	$\vartheta(f_p); f_p = \frac{1}{T_p}$	0-360 $^{\circ}$	$\pm 2^{\circ}$
Peak wave length (m)	λ_p	$\lambda_p = \frac{gT_p^2}{2\pi} \sqrt{\tanh\left(\frac{4\pi^2 d}{T_p^2 g}\right)}$	19-600 m	–
1 st , 2 nd and 3 rd significant wave height for swell systems 1, 2, and 3 (m)	$H_{s_{1,2,or3}}$	$4\sqrt{m_{0_{1,2,or3}}}$	1-20 m	$\pm 0.5 m$
1 st , 2 nd and 3 rd wave peak period for swell systems 1, 2, and 3 (s)	$T_{p_{1,2,or3}}$	$\frac{1}{f_{p_{1,2,or3}}}$	3.5-55 s	$\pm 0.5 s$
1 st , 2 nd and 3 rd wave length for swell systems 1, 2, and 3 (m)	$\lambda_{1,2,or3}$	$\frac{2\pi}{ \vec{k}_{p_{1,2,or3}} }$	19-600 m	–
1 st , 2 nd and 3 rd wave direction for swell systems 1, 2, and 3 ($^{\circ}$)	$\beta_{1,2,or3}$	$\vartheta(f_{p_{1,2,or3}})$	0-360 $^{\circ}$	$\pm 2^{\circ}$
Wave directional spreading (–)	σ_{θ}	$\sqrt{2[1 - \sqrt{\frac{a^2+b^2}{m_0^2}}]}$	0 – 90 $^{\circ}$	n/a
Inverse wave age (–)	μ	$\frac{U_{10}}{c_p}, c_p: \text{Wave phase velocity}$	–	–
Wave steepness (–)	ε	$\frac{H_S}{k} \frac{g}{2}$	–	–
Surface current speed (m/s)	u	Refer to Eq. (1)	0-20 $m s^{-1}$	$\pm 0.2 m s^{-1}$
Surface current direction ($^{\circ}$)	θ	Refer to Eq. (1)	0-360 $^{\circ}$	$\pm 2^{\circ}$

Competing interests. The authors declare that no competing interests are present.

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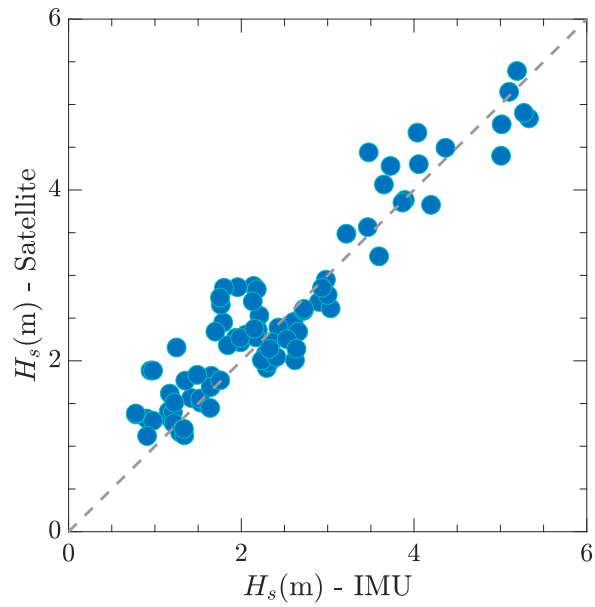


Figure B1. Satellite observations versus significant wave height reconstructed from ship motion.

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