

Final Author comments

Authors' response to Referee #1 and Referee #2 comments on "New continuous total ozone, UV, VIS and PAR measurements at Marambio 64°S, Antarctica" by Kaisa Lakkala et al.

The authors thank the Referees for constructive comments and reply to all comments here below. The answer is structured as follow: (1) comments from Referee, (2) author's response, (3) author's changes in the manuscript.

Referee #1

(1) What I mainly miss is some discussion (maybe a small paragraph) regarding the uncertainty in the GUV measurements performed at Marambio. Comparison with other instruments provides very strong evidence of the reliability of the measurements. Are however the calculated differences representative for the overall measurement uncertainties? I believe that some discussion regarding the magnitude of the overall uncertainties – not necessarily a precise determination of the uncertainty budget - and the main uncertainty factors would be useful for the readers, as well as for people interested for the data.

(2) A new paragraph (see below) has been added to Section 3.1. discussing the uncertainties of GUV data products (e.g., UV index, UVB, UVA, DNA damage).

(3) "The uncertainty of GUV data products is composed of (i) the uncertainty of SUV-100 measurements, (ii) the uncertainty of the transfer of the calibration from the SUV-100 to the GUV, (iii) the uncertainty of the conversion from response-weighted irradiance to data products D , and (iv) the drift of GUV calibrations. The uncertainty of SUV-100 measurements has been assessed by Bernhard et al. (2004). The expanded uncertainty (coverage factor $k = 2$, corresponding to a level of confidence of approximately 95 %) of the UV index and DNA-damaging irradiance varies between 5.8 and 6.4 %. The upper limit of errors in the SUV-to-GUV calibration transfer was estimated to $\pm 2\%$ from the reproducibility of the vicarious calibration method. The uncertainty in calculating the UV index from response-weighted irradiance using Eq. (1) was assessed by Dahlback (1996). For SZAs between 0 and 80° and total ozone between 200 and 500 DU, the approximations implied in using Eq. (1) agreed to within $\pm 5\%$ with exact radiative transfer calculations. However, larger errors were found for total ozone columns smaller than 200 DU or for SZAs larger than 80° (when absolute values of the UV Index are small). Bernhard et al. (2005) compared measurements of UV-B irradiance, UV-A irradiance, the UV index, and DNA-damaging performed at the South Pole with a SUV-100 and a GUV radiometer, which was calibrated with the method described in this paper. For SZAs smaller than 80°, measurements by the two instruments agreed on average to within $\pm 2.5\%$, and the standard deviation of the ratio of GUV/SUV data was smaller than 4.0% for the four data products. The magnitude of these variations is in good agreement with the theoretical calculations by Dahlback (1996). Lastly, the uncertainty attributed to drifts was calculated from the observed change of 5% in the responsivity of the GUV's 305 nm channel. The four uncertainty components were combined in quadrature and multiplied with a coverage factor of $k = 2$. For SZAs smaller than 80°, the expanded ($k = 2$) uncertainty is 9% for UV-B irradiance, the UV index, and DNA-damaging irradiance."

New reference added: Bernhard, G., Booth, C. R., and Ehramjian, J. C.: Version 2 data of the National Science Foundation's Ultraviolet Radiation Monitoring Network – South Pole, *J. Geophys. Res.*, 109, D21207, <https://doi.org/10.1029/2004JD004937>, 2004.

Specific comments:

(1) References are missing at several points in the introduction. I suggest adding references to the following phrases: P2, L4 – 5: “The international . . . UV levels” P2, L5 –6: “The . . . materials” P2, L17 – 19: “As . . . effects” P2, L26 – 27: “Both . . . Ocean” P2, L27 – 28: “Changes . . . series”

(2) References have been added.

(3) References have been added: “The international Montreal Protocol was signed in 1987 to restrict the use of ozone depleting substances and thus protect the whole ecosystem from excessive UV levels (e.g., Barnes et al. (2019); WMO (2018b)). The shortest UV wavelengths are known to cause skin cancer, sunburn and cataract, and they can also harm plants, animals and materials (e.g., EEAP (2019)).

As the stratospheric ozone depletion has influenced both stratospheric and surface climate, e.g., cooling of Antarctic stratosphere and southward shift of mid-latitude rain, and the Southern Ocean temperature and circulation, the recovery is expected to have the opposite effects (WMO, 2018a).

Both UV and PAR affect micro-organisms living in Antarctic ice and the Southern Ocean (Deppeler and Davidson, 2017; Häder et al., 2014).

Changes in the amount of aerosols and pollution as well as changes in sea ice extension or ground albedo are also reflected in both UV and VIS radiation time series (Fountoulakis et al., 2014; Wild, 2009).”

(1) P3, L6: Please replace “measurements were” with “was”

(2) Replaced as suggested.

(3) Replaced as suggested.

(1) P7, L7: in “ki”, i is an index

(2) We agree.

(3) Manuscript updated following the comment.

(1) P7, L8 – 9: “A sensitivity . . . time” is there any reference which can be used to support this statement?

(2) The statement was based on unpublished data, which at present cannot be supported with a reference. We deleted this sentence.

(3) Sentence was deleted.

(1) P8, L7: Was the sky clear in August 2016? Please specify since the conclusions from Figure 2 might be slightly different if the sky was cloudy.

(2) Yes, there was clear sky on 24 August 2016.

(3) The information was added to the manuscript and the following sentence of Section 3.2 was updated: “For clear sky and SZAs smaller than 60°, UV indices measured by both GUV radiometers agreed within 1% of each other and the Brewer spectroradiometer.”

(1) Figure 2: Could the apparent dependence of the ratio on SZA (or part of it) be a masked effect of temperature on the response of any of the instruments?

(2) The internal temperature of the GUV radiometer was monitored and kept constant at 40°C, which ensured that no drift of response was due to temperature change. The UV data of FMI’s spectroradiometer was temperature corrected using the method presented in Lakkala et al. 2008. The reference spectroradiometer QASUME was temperature stabilized during the measurement campaign. The small dependence of the ratio on SZA is likely due to a combination of several factors, including difference in the angular response of the instruments (including incomplete corrections of cosine

errors), the approximation in calculating the UV index with Eq. (1) and small time shifts between GUV and spectroradiometric measurements as the recording of a spectrum takes several minutes. (3) The information of Brewer temperature correction and temperature stabilization of the QASUME instrument was added to the text in section 3.2. Referring to the Brewer: “The measurements are cosine and temperature error corrected...” and to the QASUME “The measurements are temperature stabilized...”.

Reference: Lakkala, K., Arola, A., Heikkilä, A., Kaurola, J., Koskela, T., Kyrö, E., Lindfors, A., Meinander, O., Tanskanen, A., Gröbner, J., and Hülsen, G.: Quality assurance of the Brewer spectral UV measurements in Finland, *Atmos. Chem. Phys.*, 8, 3369–3383, <https://doi.org/10.5194/acp-8-3369-2008>, 2008.

(1) Section 3.3: Is this change in the response of the instrument (taking place within the 1 or 2 years between sequential calibrations) somehow taken into account, e.g. by interpolating the calibration factors?

(2) The change in the response of the instrument is not taken into account. The manuscript has been updated including the information.

(3) The manuscript has been updated including the information. The following sentence was added to section 3.3.: “The drift was considered to be within the uncertainties of the measurements and the time series was not corrected for. “

Referee #2

(1) Unfortunately, after analyzing the data sets and presentations in the manuscript, I found several flaws regarding use of measurement units, as well as suspecting anomalous results in final data. The manuscript is within the scope of the journal, but the data sets should be re-evaluated, and manuscript revised before publication.

(2) The authors thank the Referee for the careful review and agree with the problem in the units of the data. The mistake was found to be in the data-analyse phase, not in the original data set.

(3) The units and data have been corrected and the updated dataset is uploaded to <http://doi.org/10.5281/zenodo.3688700>.

Specific comments:

(1) Page 1 in Abstract: lines 9-11: A long list of final data products is stated being available – however in the internet link provided I found only a subset available at the given repository (e.g. 10 biologically effective dose rates and corresponding doses are missing). Please, be more specific. A screendump of files available (attached figure).

(2) We included in the available dataset the studied data sets, which are irradiances at five UV channels and one VIS channel, daily maximum UVB/UVA and PAR dose rates, noon UVB/UVA and PAR dose rates, noon total column ozone and UVB/UVA daily doses. The data including weighting using 10 different biologically active were not studied in this specific paper, even if it is routinely derived from the actual measurements. That data is available from the authors.

(3) The Abstract has been updated to be more specific: “The studied dataset, including daily maximum irradiances at five UV channels and one VIS channel, daily maximum UVB/UVA and PAR dose rates, noon UVB/UVA and PAR dose rates, noon total column ozone and UVB/UVA daily doses, is freely accessible at <http://doi.org/10.5281/zenodo.3688700>. (Lakkala et al., 2019).”

The Data availability -section has been updated to be more specific: “The studied datasets of daily maximum irradiances at five UV channels and one VIS channel, daily maximum UVB/UVA and PAR dose rates, noon UVB/UVA and PAR dose rates, noon total column ozone and UVB/UVA daily doses are freely available at Zenodo, <http://doi.org/10.5281/zenodo.3688700>. (Lakkala et al., 2019). The additional datasets including biologically weighted dose rates and daily doses are available from the authors.”

(1) Page 1 line 14-15 in Abstract: “Average daily maximum UVB dose rates 7.6

– 10.2 kW/m²”. Compare these numbers for the UVB with the Sun’s total integrated solar irradiance at the top of the atmosphere – 1.366 kW/m²

(<https://wattsupwiththat.com/2018/09/19/how-constant-is-the-solar-constant/>).

One might suspect a misprint, that the prefix k (kilo) should be omitted, but even in that case, 7.6-10.2 W/m² is almost a factor 10 above realistic natural surface irradiance levels, compared with quality controlled measurement data for an equivalent latitude and network station (e.g. mountain station Finse in the UV-monitoring network in Norway, spanning latitudes 58-78 N <https://github.com/uvnrpa>). The same applies to Table 4, UVB DMDR [kW/m²] and UVA DMDR [kW/m²], as well as Figure 5C and 5D, where units and numbers also look anomalous.

(2) The authors thank the Referee for the careful review and agree with the problem in the units of the data. The mistake was found to be in the data-analyse phase, not in the original data set. The units and data have been corrected and the updated dataset is uploaded in <http://doi.org/10.5281/zenodo.3688700>.

(3) The units have been corrected and the Table 4 and Figure 5 (now 7) have been updated. The sentence in the abstract has been updated to be “ Average daily maximum UVB dose rates 76 – 102 uW/cm²...”

(1) Page 3-4, section 2.1 Marambio station and section 2.2: The text relates much to the same content provided in section 2.1.1 and section 2.2.1 in another publication under discussions by the two first authors, where the same datasets are applied <https://www.atmos-chem-phys-discuss.net/acp-2019-896/acp-2019-896.pdf> .

Figure 4A and Figure 4B on page 10 in Lakkala et al. is almost identical with Figure 5 and Figure 6 in the second paper submitted by Aun et al. 2019. Although the two papers deal with different topics: one presenting the data sets, QA/QC and resources, and the second paper an analysis of UVI and erythemal doses in relation to the long-term series, there are redundancies between the two papers.

(2) As commented by the Referee, the two papers have different topics. This one, Lakkala et al., focuses on the set up of the new measurement system, quality assurance and it’s data set, and gives example of the use of the data by presenting products (UV, PAR and VIS) measured during the first two years of operation. Aun et al., 2019 analyse the characteristics of UV index and erythemally weighted UV time series during the first two years, compare them to earlier measurements (2000-2008) and to measurements from other Antarctic sites.

The text of sections 2.1. and 2.2. has been revised to avoid same content/phrasing. Naturally the sections include in some extent similar content than the sections 2.1.1 and 2.2.1. in Aun et al., as the sections are about the same measurements. The revised version of the section 2.2.1. of Aun et al., 2019 has been updated and includes now references to Lakkala et al.

We agree that the Figures 4A and 4B in Lakkala et al. are very similar to Figures 5 and 6 of Aun et al. However, we think that for a reader, the UV index and total column ozone are the most interesting and familiar, and we would prefer to keep the Figures in our manuscript. To avoid overlapping, we have not plotted the same quantities: In Aun et al. the plotted data in Figure 5 is **daily maximum** UV index and in Figure 6 **daily average** total ozone, while in Lakkala et al. the Figure 4a and b includes UV index

and total ozone measured **at noon**. At Marambio, total column ozone can change during the day so that the daily average is not the same as the one measured at noon. Regarding UV index, the noon value is the maximum for a clear sky day, but as Marambio has frequent changing cloudiness conditions, the daily maximum UV index can be measured either before or after noon, if the sky is cloudy during noon, but free from clouds later or earlier during the day.

The manuscript includes already the following sentences to clarify the differences between the two papers (First two sentences in Section 4.): “The erythemally weighted UV, maximum UV index and daily average total ozone time series were discussed in Aun et al.(2019) and the results were compared to measurements from other Antarctic measurement sites. In this paper, the time series of noon UV index, noon total ozone, UVB, UVA, VIS and PAR measurements for the period March 2017 – May 2019 are presented in Figs. 4 and 5 (Lakkala et al., 2019).”

(3) The text of sections 2.1. and 2.2. has been revised to avoid same phrasing.

(1) Page 5, Figure 1c and Figure 1 d: Units are given as uW/m^2 . The units should probably be given as uW/cm^2 (difference 10^4).

(2) The authors agree.

(3) The Figures have been updated.

(1) Pages 7-8. The section describes calibrations and comparisons of UVI against spectroradiometers SUV, Brewers and QASUME. I miss a comparison which includes other data products as well in order get an estimate of overall uncertainties.

(2) The comparison of UVB and total ozone have been added to the analysis. Comparisons of UVA, PAR and VIS products were not possible, as the spectroradiometer at Sodankylä doesn't measure these wavelengths. Bernhard et al., (2005) have validated several GUV radiometer products against a SUV spectroradiometer. The same methodology was used for the calibration of FMI's GUV radiometers, and the accuracy of the measurements can be assumed to be similar. A discussion about overall uncertainties has been added to the text in Section 3.1.

(3) Figures of UVB and total ozone column comparisons in Sodankylä have been added to the manuscript sections 3.2. and 3.2.1. The following discussion about the overall uncertainties has been added to the Section 3.1. :“The uncertainty of GUV data products is composed of (i) the uncertainty of SUV-100 measurements, (ii) the uncertainty of the transfer of the calibration from the SUV-100 to the GUV, (iii) the uncertainty of the conversion from response-weighted irradiance to data products D , and (iv) the drift of GUV calibrations. The uncertainty of SUV-100 measurements has been assessed by Bernhard et al. (2004). The expanded uncertainty (coverage factor $k = 2$, corresponding to a level of confidence of approximately 95 %) of the UV index and DNA-damaging irradiance varies between 5.8 and 6.4 %. The upper limit of errors in the SUV-to-GUV calibration transfer was estimated to $\pm 2\%$ from the reproducibility of the vicarious calibration method. The uncertainty in calculating the UV index from response-weighted irradiance using Eq. (1) was assessed by Dahlback (1996). For SZAs between 0 and 80° and total ozone between 200 and 500 DU, the approximations implied in using Eq. (1) agreed to within $\pm 5\%$ with exact radiative transfer calculations. However, larger errors were found for total ozone columns smaller than 200 DU or for SZAs larger than 80° (when absolute values of the UV Index are small). Bernhard et al. (2005) compared measurements of UV-B irradiance, UV-A irradiance, the UV index, and DNA-damaging performed at the South Pole with a SUV-100 and a GUV radiometer, which was calibrated with the method described in this paper. For SZAs smaller than 80° , measurements by the two instruments agreed on average to within $\pm 2.5\%$, and the standard deviation of the ratio of GUV/SUV data was smaller than 4.0% for the four data products. The magnitude of these variations is in good agreement with the theoretical calculations by Dahlback (1996). Lastly, the

uncertainty attributed to drifts was calculated from the observed change of 5% in the responsivity of the GUV's 305 nm channel. The four uncertainty components were combined in quadrature and multiplied with a coverage factor of $k = 2$. For SZAs smaller than 80° , the expanded ($k = 2$) uncertainty is 9% for UV-B irradiance, the UV index, and DNA-damaging irradiance."

New reference added: Bernhard, G., Booth, C. R., and Ehramjian, J. C.: Version 2 data of the National Science Foundation's Ultraviolet Radiation Monitoring Network – South Pole, *J. Geophys. Res.*, 109, D21207, <https://doi.org/10.1029/2004JD004937>, 2004.

(1) Table 5 and Table 6. Please, consider if this information is too detailed in this context.

(2) We agree and the tables were removed.

(3) The tables were removed.

(1) Page 14, Figure 5E: Units is given as $\text{mW}/\text{cm}^2/\text{nm}$ at 555 nm. Realistic units is $\text{uW}/\text{cm}^2/\text{nm}$ (difference 10^3). Page 14, Figure 5F: PAR, units given as $\text{E}/\text{m}^2/\text{s}$. Realistic units is $\text{E}/\text{cm}^2/\text{s}$ (difference 10^4).

(2) We agree and the units have been updated. Note that the PAR is $\text{uE}/\text{cm}^2/\text{s}$.

(3) Units have been updated.

(1) Page 16, Section 5 Data availability. The units of most datasets look correct. However, the six data files max305nm.dat etc are expressed in units $\text{W}/\text{cm}^2/\text{nm}$. The units should probably be $\text{uW}/\text{cm}^2/\text{nm}$ (difference 10^6). I have plotted the spectral irradiance data from these files, and changed observations to appropriate units, and made a model comparison, see attached figure. The model takes as input the total ozone given in file GUV_UVB_UVA_PAR_O3_noon.dat, the noon SZA for Marambio, assuming snowfree ground (albedo given as 0.2 to 0.3 for November/December in Aun et al. 2019) and assuming clear sky conditions.

(2) We agree with the unit problem. They have now been updated.

(3) Units have been updated to $\text{uW}/\text{cm}^2/\text{nm}$.

(1) Irradiance values of observations look anomalous for the 313 and 320 nm (factor almost 2x), but reasonable for 305 nm, 340 nm and 380 nm. You can see this by observing the differences in spectral irradiances for increasing wavelengths of observations and modelled data: Subset of data covering 2017/2018, wavelengths 305-313-320-340-380 nm: Observations: 4-26-16-54-74 $\text{uW}/\text{cm}^2/\text{nm}$ Model predications: 4-15-26-54-64 $\text{uW}/\text{cm}^2/\text{nm}$ (continuous increase, matching observations at the three wavelengths 305, 340 and 380 nm).

(2) We thank the Referee for the careful review of manuscript and for the model calculations. By downloading the data set from <http://doi.org/10.5281/zenodo.3553634> and plotting it, we didn't get the same features than the Referee. We think that the irradiances of 320 and 313 nm have accidentally been mixed in the analysis of the Referee. Please find here below (Fig.1) a plot of irradiances at 305, 313, 320, 340 and 380 nm. We think that these time series match well with the model calculations presented by the author.

(3) None.

(1) Page 16, line 26-27. "...crucial to obtain homogenized long-term series.". Here I miss references to the international intercomparison of multi-band filter radiometers, held in Oslo 2005: GAW report no. 179 / WMO/TD-No. 1454.

Geneve: World Meteorological Organization. 2008.

(2) We added the reference.

(3) Reference added.

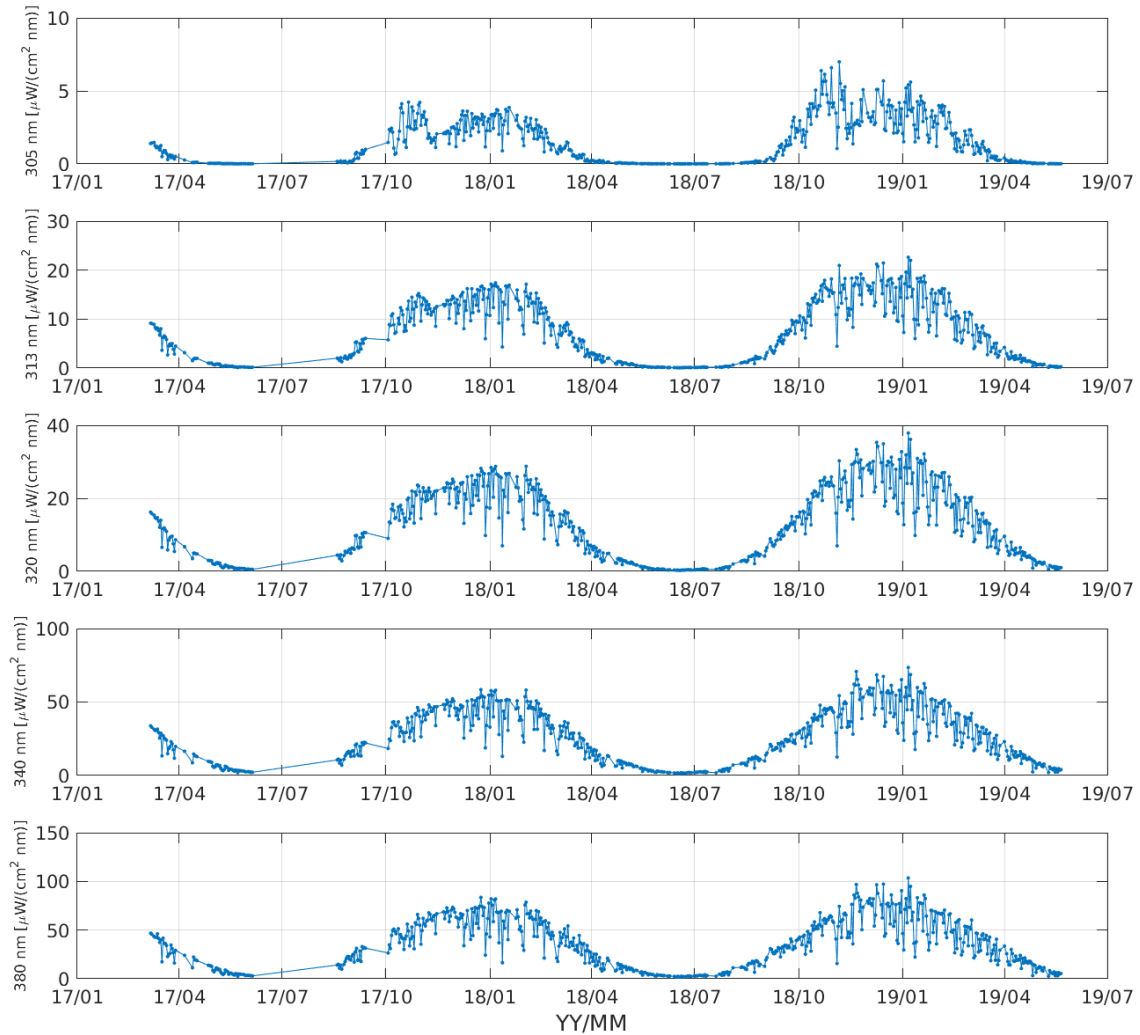


Figure 1. UV irradiances measured at Marambio, plotted from the dataset 10.5281/zenodo.3688700.

New continuous total ozone, UV, VIS and PAR measurements at Marambio 64°S, Antarctica

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Abstract. A GUV multifilter radiometer was set up at Marambio, 64°S 56°W, Antarctica, in 2017. The instrument measures continuously ultraviolet (UV) radiation, visible (VIS) radiation and photosynthetically active radiation (PAR). The measurements are designed for providing high quality long-term time series which can be used to assess the impact of global climate change in the Antarctic region. The quality assurance includes regular absolute calibrations and solar comparisons performed at the [site Marambio](#) and at Sodankylä, Finland. The ~~actual measurements continue the time series measured at Marambio~~ [measurements continue observations at Marambio that were performed](#) with NILU-UV radiometers ~~during 2000–2010 between 2000 and 2010~~ as part of the Antarctic NILU-UV network. ~~They are optimal. These measurements are ideally suited~~ for assessing the effects of the ongoing stratospheric ozone recovery on the ecosystem as the data products include information on radiation at various wavelengths ranging from UV to VIS so that changes on biologically effective radiation due to ozone can be separated from those due to other factors. ~~The final data products are~~ [Data products include](#) total ozone, ~~PAR, VIS~~ [photosynthetically active radiation \(PAR\), visible \(VIS\)](#) radiation at 555 nm, UV index, UV irradiance at 5 channels, UVB and UVA dose rate/daily dose, and biologically weighted UV dose rate/ daily dose, ~~including~~ [calculated with](#) 10 different action spectra. The data from the last five days and the daily maximum UV index time series are plotted and updated daily on the web page fmiarc.fmi.fi/sub_sites/GUVant/. The first two years of UV measurements were very different in terms of the results: ~~The~~ [For October, November and December the](#) monthly average of daily maximum UVB dose rates were clearly higher in 2018 than in ~~2017 during the period from October to December. 2017.~~ The largest difference was observed in October, when the average of daily maximum UVB dose rates was ~~7.6 kWm~~ [76 \$\mu\text{W cm}^{-2}\$](#) and ~~10.2 kWm~~ [102 \$\mu\text{W cm}^{-2}\$](#) in 2017 and 2018, respectively. ~~The monthly averages were close to each other for all~~ [Monthly averages of](#) the three months ~~were similar~~ in 2018, while in 2017 the monthly average of October was lower than those of November and December. VIS and PAR time series show that daily maxima in 2018–2019 exceed those in 2017–2018 during ~~the~~ late spring and ~~the~~ summer (mid-November – January). The studied dataset, [including daily maximum irradiances at five UV channels and one VIS channel, daily maximum](#)

[UVB/UVA and PAR dose rates, noon UVB/UVA and PAR dose rates, noon total column ozone and UVB/UVA daily doses](http://doi.org/10.5281/zenodo.3688700), is freely accessible at <http://doi.org/10.5281/zenodo.3688700> (Lakkala et al., 2019).

1 Introduction

- The Antarctic area suffers from the well-known springtime ozone hole since the late 1970s (e.g., Farman et al., 1985; WMO, 1990). Increase of harmful UV radiation reaching the surface was expected, as the stratospheric ozone layer protects the Earth from the most dangerous ultraviolet (UV) wavelengths. The international Montreal Protocol was signed in 1987 to restrict the use of ozone depleting substances and thus protect the whole ecosystem from excessive UV levels (e.g., Barnes et al. (2019); WMO (2018b)). The shortest UV wavelengths are known to cause skin cancer, sunburn and cataract, and they can also harm plants, animals and materials (e.g., EEAP (2019)).
- The Spanish-Finnish-Argentinean Antarctic NILU-UV network was established in 1999/2000 to monitor UV radiation, photosynthetically active radiation (PAR) and total ozone in the Antarctic area and to serve the multidisciplinary UV research community (Lakkala et al., 2008). The NILU-UV multifilter radiometer measurements stopped due to degradation of the instrument at Marambio and other stations in 2013, but they provide reference time series for the severe Antarctic ozone hole period. Between years 2000–2010 the UV index at Marambio reached a maximum of 12 when the station was inside the ozone hole (Lakkala et al., 2018b).
- The latest Scientific Assessment of Ozone Depletion (WMO, 2018b) suggests that the Montreal Protocol has been successfully adapted and the Antarctic ozone hole has started to recover even though it continues to occur each year. To detect the recovery and its effects on UV radiation and the ecosystem, the Finnish Meteorological Institute (FMI) and the Servicio Meteorológico Nacional Argentina (SMN) started UV and total ozone measurements with multifilter GUV radiometers at Marambio in 2017. As the stratospheric ozone depletion has influenced both stratospheric and surface climate, e.g., cooling of Antarctic stratosphere and southward shift of mid-latitude rain, and the Southern Ocean temperature and circulation, the recovery is expected to have the opposite effects (WMO, 2018a). However the influence of the ozone recovery on climate and ocean ~~can not~~ cannot easily be predicted as the influence will depend on the evolution of green house gas concentrations in the atmosphere which is the key driver of future southern hemisphere climate (WMO, 2018b).
- The total ozone, UV radiation, visible (VIS) radiation and PAR measurements of the GUV multifilter radiometer can be used in research assessing the effects of the ongoing climate change. For example, cloud optical depth can be retrieved using measured irradiances (~~Bernhard et al., 2005~~) (Dahlback, 1996) and, together with the irradiance of the visible channel, cloud optical depth can give information on changes in cloudiness. PAR measurements are directly applicable in studies of the effects of climate change on the photosynthesis of plants, algae and bacteria. Both UV and PAR affect micro-organisms living in Antarctic ice and the Southern Ocean (Deppeler and Davidson, 2017; Häder et al., 2014). Changes in the amount of aerosols and pollution as well as changes in sea ice extension or ground albedo are also reflected in both UV and VIS radiation time series (Fountoulakis et al., 2014; Wild, 2009).

In 2013, the FMI installed Solar Light 501A radiometers at Marambio to monitor both incoming and outgoing UV radiation. The measurements are used to determine local albedo and detect changes in it. They can be used to link the NILU-UV measurement time series to the new GUV measurements as they monitor the erythemally weighted UV irradiance, which was one product of the NILU-UV radiometer measurements and is now one of the GUV radiometer products. The Solar Light 501A data collection system consists of one upward and one downward radiometer, which measure the ~~incoming and outgoing~~ irradiance weighted with the action spectrum for UV induced erythema (McKinlay and Diffey, 1987), which in turn also has a contribution from the UVA. The radiometer pair is selected to represent as similar spectral and cosine responses as possible, as demonstrated in Fig. 1 of Meinander et al. (2008). The system measures in one minute intervals and saves data automatically.

In this paper, the new UV, VIS, PAR and total ozone measurements at Marambio are described and the quality assurance procedures are discussed. Time series of the first two years of measurements are shown.

2 GUV multifilter radiometer measurements at Marambio

GUV multifilter radiometer ~~measurements were~~ was installed at Marambio, Antarctica, in collaboration between the SMN and the FMI, in March 2017. The SMN is responsible for the operation and quality control of the measurements at the station, while the FMI is responsible for the quality assurance, data dissemination and data storage.

2.1 Marambio station

~~The station of Marambio, Marambio station in Antarctica is located on the highest part of the Seymour/Marambio ice-free Island, surrounded by Weddell sea, at 64° 14' S; 56° 37' W, was founded in 1969 and contributes to the World Meteorological Organization (WMO) Atmospheric Watch Programme. It is located on an island in the eastern on the north-east side of the Antarctic Peninsula. The altitude of the station is 198 m above the sea level. The temperature ranges from -30°C in winter to During the last decades the station is well-known for studies related to the ozone hole phenomena because it is located at the edge of the polar vortex. The temperatures at the site are between +10°C in summer and the mean monthly temperature in June-July is around -15 during the summer, and -30°C and in December-January around -2C during the winter, though strong winds can lower the apparent temperature down to -60°C (e.g., Asmi et al., 2018)C. During most part of the year, the soil is frozen and covered with snow. The Weddell Sea to the east of the Peninsula is frozen year round. During the winter months the ice may extend to latitudes of around 60°S, also covering the coast of the Marambio island. Snow cover is variable but typically exists throughout the winter and a major part of the summer Marambio Island. The prevailing wind directions are from the southwest and the northwest. The wind speed can reach values close to 100 km/h. Heavy cloudiness and fog are common during the summer months, and heavy winds in the winter can blow the snow away yielding to lifting of dust into the air. The prevailing winds at the station are from southwest and northwest station is part of the Global Atmospheric Watch (GAW) program of the World Meteorological Organization (WMO). More information on the site can be found from earlier publications (e.g., Karhu et al., 2003; Asmi et al., 2018).~~

~~Previously,-~~

UV measurements in Marambio have been conducted with a NILU-UV multichannel radiometer ~~during the years 2000–2013~~ between 2000 and 2013 (Lakkala et al., 2018b), and before that, measurements of erythral UV radiation were carried out by a biometer Solar Ligth 501 installed in 1996 by SMN. The instrument was calibrated by personnel from the World Radiation Center (WRC) in 2006 and 2010. In 2013 it was replaced by a similar instrument belonging to the FMI, when also a radiometer was installed to measure outgoing solar UV radiation (Meinander et al., 2014). Details about the calibration of the new radiometers can be found in Meinander et al. (2008).

Synoptic observations are performed hourly at the Marambio station and additional weather data is obtained from an automatic station. In addition to the synoptic visual cloud observations, cloud coverage and height in Marambio are monitored since the year 2016 with a ceilometer model CL51 (Vaisala), with 30 second resolution.

Marambio also reports to the World Ozone and UV Data Center (WOUDC) and has additionally programs to monitor greenhouse gas concentrations and various aerosol parameters. The number concentration of aerosol particles larger than 10 nm in radius is measured with a Condensation Particle Counter (CPC) model 3772 (TSI Inc.) installed as a part of the particle size distribution measurement system. Aerosol optical properties at the surface are measured with online instrumentation: scattering with an Aurora 3000 nephelometer and absorption with a Multi-Angle Absorption photometer (MAAP) model 5012 (Asmi et al., 2018). Aerosol chemistry is measured offline from the collected weekly filter samples (see e.g., Asmi et al., 2018). Aerosol optical depth in the atmospheric column is measured with a PFR sun photometer which is part of the GAW-PFR network (Tomasi et al., 2015).

2.2 GUV radiometer

The radiometer installed in Marambio is a GUV radiometer, model ~~GUV-2511, is a multifilter radiometer~~ GUV-511, manufactured by Biospherical Instrument Inc. (BSI), ~~US. The instrument was purchased to monitor~~ USA, which monitors UV, visible and PAR radiation ~~at Marambio. The radiometer includes five channels at wavelengths in the UV and one channel in the visible part of the electromagnetic spectrum. The~~. It includes six channels which central wavelengths are at 305, 313, 320, 340, 380 and 555 nm, and the full width at half maximum (FWHM) is around 10 nm. In addition there is a seventh channel which measures PAR in the 400–700 nm wavelength region. Using the combination of different channels offers the possibility to retrieve cloud optical thickness and total ozone column. The instrument is environmentally sealed and temperature-stabilized at 40°C. It has a Teflon-covered quartz cosine collector. The instrument specifications, as provided by the manufacturer, are shown in Table 1. Details of the GUV radiometer and its performance are described in Bernhard et al. (2005).

2.3 Data retrieval and products

The sampling rate of the GUV radiometer can be set to 1-20 Hz and one minute averages are automatically recorded using **LOGGER** data-acquisition software. The software converts raw data into irradiances, which are saved as ascii and Microsoft Access Database formats. The daily data is automatically transferred to an ftp server in the SMN, from which the data is transferred and stored into the FMI's data base. Measurements from the last five days and the daily maximum UV index time series are plotted and updated daily on the web page fmiaarc.fmi.fi/sub_sites/GUVant/. Examples of these web plots including

Table 1. GUV model 2511 irradiance array specifications.

Filter type	Custom low-fluorescence interference
Cosine collector	Teflon -covered quartz
Collector area	2.1 cm diameter
Out-of-band rejection	1×10^{-6}
Angular response	0-5% from 0° to 70°; $\pm 10\%$ from 71° to 85°
Typical Saturation	$10^5 \mu\text{Wcm}^{-2} \text{ nm}^{-1}$
Noise equivalent irradiance	$10^{-11} \text{ Wcm}^{-2} \text{ nm}^{-1}$
Temperature coefficient of the dark Signal	Less than $\pm 3 \mu\text{volts}/^\circ\text{C}$
Response Temperature Coefficient	Less than $\pm 0.15 \text{ \%}/^\circ\text{C}$

UV index, UVB, UVA and PAR radiation are shown in Fig. 1. In addition to routine checks of data transfer, the housekeeping includes cleaning of the diffusers at least once a week, or more frequently if needed, and checking of the leveling once a month.

Near real time data processing includes calculation of data products based on UV, VIS and PAR measurements. The main idea is that each UV product is calculated using a linear combination of calibrated irradiances measured at the five UV channels. The coefficients are determined from solar comparisons between the GUV radiometer and a spectroradiometer following the theory-method of Dahlback (1996) and Bernhard et al. (2008) as explained in Section 3.1. Bernhard et al. (2005) showed that GUV multifilter UV data was within 3% agree within 5% from data from a well calibrated spectroradiometer for solar zenith angles (SZAs) smaller than 80°. In addition to UV products, VIS radiation and PAR are measured using the 555 nm and 400–700 nm channels, respectively. Total ozone is calculated using lookup tables and the ratio of irradiance measured at 305 nm (a wavelength strongly attenuated by ozone) and 340 nm (a wavelength weakly attenuated by ozone). The complete list of products including several biologically weighted UV dose rates and doses and their references can be found in Table 4.4. of Bernhard et al. (2008). Table 2 summarizes the products calculated from the GUV measurements at Marambio.

3 Quality assurance of the measurements

The quality assurance of the GUV measurements is based on regular absolute calibrations and solar comparisons against high quality spectroradiometers. The FMI has purchased two GUV radiometers in order to avoid gaps in measurement time series. Those two instruments rotate so that one is measuring at Marambio, while the other one is calibrated and compared against reference instruments. Solar comparison is performed each time when the instruments are switched, to detect possible drift or jump in the calibration scale.

Solar comparisons against spectroradiometers are performed at Sodankylä 67°N, in Finland. As the-this site is located at high-latitudes a high latitude, similar to Marambio, the atmospheric path of UV radiation reaching the surface is similar at both

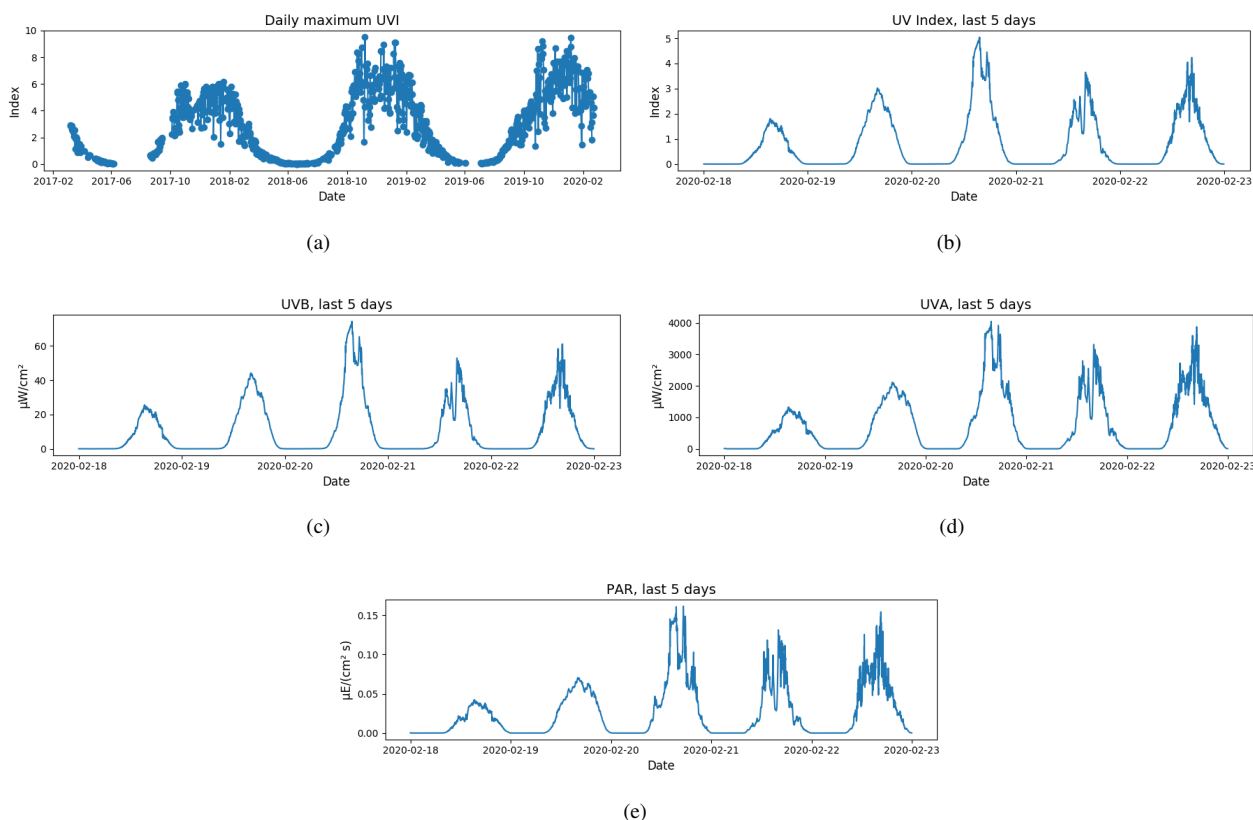


Figure 1. Near real time plots updated daily: A) Daily maximum UV index time series, b) UV index, c) UVB irradiance, d) UVA irradiance and e) PAR. Data in plots b–e are one minute averages.

locations: long path in the atmosphere due to high-large SZAs, which means more ozone absorption and scattering than at lower-smaller SZAs. Sodankylä can be classified as an Arctic site from the point of view of stratospheric meteorology. This means that, as Marambio, the station can be located inside, outside or at the edge of the stratospheric Arctic polar vortex, and severe spring time stratospheric ozone loss can occur.

- 5 Solar comparisons make it possible to use comparable irradiance scales in both Marambio and Sodankylä. For example, by comparing UV levels at the Arctic and Antarctic location, the effect of ozone loss events on UV radiation can be studied.

3.1 Absolute calibration

Both GUV-2511 radiometers (serial numbers #162 and #163) were vicariously calibrated at BSI in April 2016 against measurements of a SUV-100 spectroradiometer that is part of the NSF UV Monitoring Network (Booth et al., 1994). GUV and SUV instruments were installed side by side on the roof platform of BSI in San Diego, California. The SUV-100 measures spectra

Table 2. Data products of the GUV radiometer at Marambio.

Product	Wavelengths [nm] / Biological effect
UV Irradiance	305, 313, 320, 340, 380
UVB and UVA dose rate and daily dose	290–315, 290–320, 315–360, 320–360, 360–400, 315–400, 320–400
Visible irradiance	555
Photosynthetically active radiation (PAR)	400-700
UV Index	
Total ozone	
Biologically weighted UV dose rate and daily dose	Erythema, DNA damage, skin cancer in mice, skin cancer in mice corrected for human skin, generalized plant damage, plant growth, damage to anchovy, inhibition of phytoplankton carbon fixation, Inhibition of phytoplankton photosynthesis of phaeodactylum and prorocentrum, inhibition of photosynthesis in Antarctic phytoplankton

of global irradiance between 290 and 600 nm with a spectral resolution of 1 nm. The GUV radiometer with serial number 163 was recalibrated at BSI in April/May 2019 using the same approach.

The calibration method is identical to the method described in Section 4.3 of Bernhard et al. (2008). The procedure is based on the work by Dahlback (1996). In brief, uncalibrated, dark signal corrected measurements of each channel of the GUV radiometers are regressed against cosine error corrected measurements of the SUV-100 spectroradiometer. Measurements of the SUV-100 are weighted with the spectral response functions of the GUV prior to performing the regression. The procedure results in a calibration factor k_i for each channel i of the GUV. Calibration factor established for GUV #163 in 2016 and 2019 agree to within $\pm 2\%$ with the exception of the channel at 305 nm, which drifted by $\pm 5\%$. A sensitivity change of this magnitude is not uncommon for a new instrument and typically diminishes over time.

Calibrated measurements are finally calculated by dividing the GUVs' raw data with these calibration factors. Measurements of the GUVs' PAR channel are calibrated slightly differently (Bernhard et al., 2008).

Because the SUV-100 data were weighted with the GUV's response functions, calibrated measurements of the GUV are "response-weighted" irradiances (Seckmeyer et al., 2010). The conversion from response-weighted irradiance to useful data products D , such as spectral irradiance for a given wavelength, erythema irradiance, or the UV index, is performed with the method suggested by Dahlback (1996). In brief, D is approximated by a linear combination of the dark signal corrected signals of the GUVs' UV channels V_i :

$$D = \sum_{i=1}^5 a_i V_i, \tag{1}$$

where the coefficients a_i depend on the calibration factors k_i and the action spectrum of the biological effect of interests (e.g., erythema response). The coefficients are determined by solving a system of linear equations as described by Bernhard

et al. (2008), taking into account the conditions at the deployment site (e.g., range of total ozone and surface albedo). ~~A comprehensive validation~~

The uncertainty of GUV data products is composed of (i) the uncertainty of SUV-100 measurements, (ii) the uncertainty of the transfer of the calibration from the SUV-100 to the GUV, (iii) the uncertainty of the conversion ~~method can be found in the paper by Bernhard et al. (2005)~~ from response-weighted irradiance to data products D, and (iv) the drift of GUV calibrations. The uncertainty of SUV-100 measurements has been assessed by Bernhard et al. (2004). The expanded uncertainty (coverage factor $k = 2$, corresponding to a level of confidence of approximately 95 %) of the UV index and DNA-damaging irradiance varies between 5.8 and 6.4 %. The upper limit of errors in the SUV-to-GUV calibration transfer was estimated to ± 2 % from the reproducibility of the vicarious calibration method. The uncertainty in calculating the UV index from response-weighted irradiance using Eq. (1) was assessed by Dahlback (1996). For SZAs between 0 and 80° and total ozone between 200 and 500 DU, the approximations implied in using Eq. (1) agreed to within ± 5 % with exact radiative transfer calculations. However, larger errors were found for total ozone columns smaller than 200 DU or for SZAs larger than 80° (when absolute values of the UV Index are small). Bernhard et al. (2005) compared measurements of UV-B irradiance, UV-A irradiance, the UV index, and DNA-damaging performed at the South Pole with a SUV-100 and a GUV radiometer, which was calibrated with the method described in this paper. For SZAs smaller than 80° , measurements by the two instruments agreed on average to within ± 2.5 %, and the standard deviation of the ratio of GUV/SUV data was smaller than 4.0 % for the four data products. The magnitude of these variations is in good agreement with the theoretical calculations by Dahlback (1996). Lastly, the uncertainty attributed to drifts was calculated from the observed change of 5 % in the responsivity of the GUV's 305 nm channel. The four uncertainty components were combined in quadrature and multiplied with a coverage factor of $k = 2$. For SZAs smaller than 80° , the expanded ($k = 2$) uncertainty is 9 % for UV-B irradiance, the UV index, and DNA-damaging irradiance.

Total ozone column is calculated from GUV measurements with lookup tables, which relate total column ozone to SZA and the ratio of GUV measurements at 305 and 340 nm. The retrieval algorithm is similar to the method described by Stamnes et al. (1991). Lookup tables are calculated with the radiative transfer model UVSPEC/libRadtran (Mayer and Kylling, 2005) and resulting model spectra are weighted with the GUV response functions at 305 and 340 nm. On average, GUV total ozone data agree with OMI data to within $\pm 5\%$ for SZA smaller than 75° . At larger SZA, differences become greater due to the dependence of the retrieval on the vertical distribution of ozone in the atmosphere. Also noise in the GUV 305 nm channel affects the ability to calculate total ozone with good precision at large SZAs. A systematic comparison of GUV and TOMS total ozone measurements at several sites was performed by Bernhard et al. (2005). For SZAs $< 80^\circ$, the bias between GUV and TOMS measurements is less than 5 %.

3.2 Comparisons against spectroradiometers at Sodankylä

The ~~GUV radiometer measurements~~ UV radiation and ozone measurements of the GUV radiometers were compared against high quality spectroradiometer measurements during each Northern Hemisphere summer since 2016 (Table 3). The comparisons were performed at the measurement platform next to the sounding station of the FMI Arctic Space Centre in Sodankylä. The FMI has two Brewer spectroradiometers, which routinely measure UV spectra about every half hour (Mäkelä et al., 2016).

The measurements are cosine [and temperature](#) error corrected (Lakkala et al., 2018a) and the irradiance scale is traceable via the National Standard Laboratory MIKES, Aalto University, Finland, to the scale maintained by the National Research and Testing Institute (SP) (Heikkilä et al., 2016; Lakkala et al., 2008). As the measurement of a Brewer UV spectrum takes around 3 minutes while the GUV radiometer records irradiances averaged over one minute, the three minutes averages of GUV measurements were used in the comparison [of UV measurements](#). UV indices [and UVB dose rates](#) calculated from Brewer #37 spectra and from GUV irradiances were compared and results are shown as function of SZA in [FigFigs. 2 – 3](#). [As the Brewer #37 measures only until wavelengths up to 325 nm, only the comparison of the UVB part of the UV spectrum is possible. For the calculation of Brewer UV indices, as explained in Mäkelä et al. \(2016\), a predefined UVA spectrum is scaled to the last wavelength to complement the measurements and thus take into account the whole effective wavelength range. The erythral weighting function, which is used in the calculation of UV indices, approaches zero towards longer UVA wavelengths and the uncertainty caused by this UVA extension is only of the order of \$10^{-1}\$ \(Mäkelä et al., 2016\).](#)

The first comparison was made in August 2016, [right shortly](#) after the radiometers were purchased and before the first one was sent to Marambio. For [SZAs lower-clear sky and SZAs smaller](#) than 60° , UV indices measured by both GUV radiometers agreed within 1% of each other and the Brewer spectroradiometer. [The differences for UVB dose rates ranged from 3 to 5 %.](#)

After the comparison, the GUV radiometer #163 was sent to Antarctica and continuous measurements started there in March 2017.

The GUV radiometer #162 stayed in Sodankylä and was compared with the Brewer UV indices during the next years, in June 2017 and July 2018. The sky was clear from clouds during the comparisons. [UV indices-Both UV indices and UVB dose rates](#) agreed within 3% for SZAs smaller than 60° in 2017. In 2018, the differences [in UV indices and UVB dose rates](#) ranged from 2 to 6 % [and from 0 to 3 %, correspondingly](#), depending on the SZAs. The SZA dependency was due to the non-corrected angular response of the GUV radiometer. Non-perfect leveling might have caused small azimuth dependency in 2018.

The portable UV world reference spectroradiometer QASUME from the World Calibration Center for UV (WCC-UV) at the Physikalisch-Meteorologisches Observatorium Davos, World Radiation Center (PMOD/WRC) visited Sodankylä for a site audit in June 2018. QASUME is a double monochromator spectroradiometer, whose solar UV irradiance measurements are traceable to the primary spectral irradiance standard of the Physikalisch-Technische Bundesanstalt (PTB), Germany, (Gröbner and Sperfeld, 2005). The [measurements are temperature stabilized and the](#) expanded relative uncertainty (coverage factor $k=2$) of solar UV irradiance measurements is 3.1% for SZAs smaller than 75° (Hülse et al., 2016). The GUV #162 was measuring on the site during the audit and the ratios of UV indices derived from it's measurements and QASUME's spectra were calculated. The results are shown in Fig. 2. The weather was unstable during the whole week when the comparisons were made, with changing cloudiness and rain. The day with the thinnest clouds was chosen and spectra measured under rapidly changing cloudiness were excluded from the analysis. The results show that the GUV radiometer UV indices were within $\pm 5\%$ from the QASUME UV indices for SZA [lower-smaller](#) than 60° . Although measurements performed during rapidly changing cloudiness were excluded from the analysis, the influence of changing cloudiness during the Brewer scan can be identified as the outlier in Fig. 2, at around SZA 56° on 5 June 2018.

Table 3. Solar comparisons at Sodankylä, Finland.

Date	GUV radiometer	Spectroradiometers
August 2016	GUV162, GUV163	Brewer037, Brewer214
June 2017	GUV162	Brewer037, Brewer214
June 2018	GUV162	Brewer037, Brewer214, QASUME Bentham
July 2018	GUV162	Brewer037, Brewer214

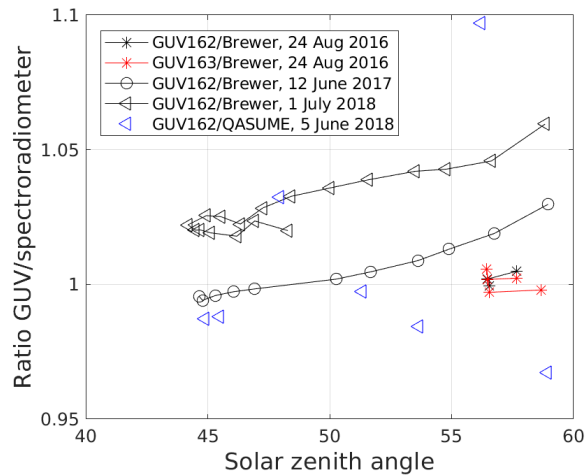


Figure 2. Ratio of the GUVs' and spectroradiometers' UV index at Sodankylä.

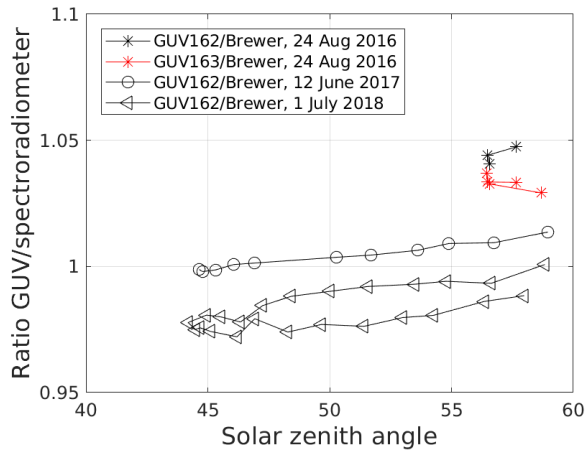


Figure 3. Ratio of the GUVs' and Brewer spectroradiometer's UVB dose rates at Sodankylä.

3.2.1 Total ozone column

The total ozone column measured by the GUV radiometer was compared to Brewer measurements during the same days than the UV comparisons were performed in Sodankylä. The Brewer #214 was used as a reference instrument as it has proven good performance during international calibrations and comparison campaigns. The Brewer participated in the international comparison campaign held in Huelva, Spain in May 2017. The results showed that the instrument agreed to the reference within 1% (Redondas et al., 2019) after proper calibration. The calibration is maintained by regular, every second year, visits of the International Ozone Service (IOS) to perform the maintenance and calibration at the measurement site.

The Fig. 4 shows the results of the comparisons in 2016, 2017 and 2018. For the Brewer radiometer, only direct sun measurements with a standard deviation less than 2.5 DU were included in the study. The nearest, which was within one minute, GUV radiometer measurement was chosen for the study. The total ozone column measurements of the GUV and the Brewer agreed to within $\pm 2\%$ for SZAs smaller than 60° .

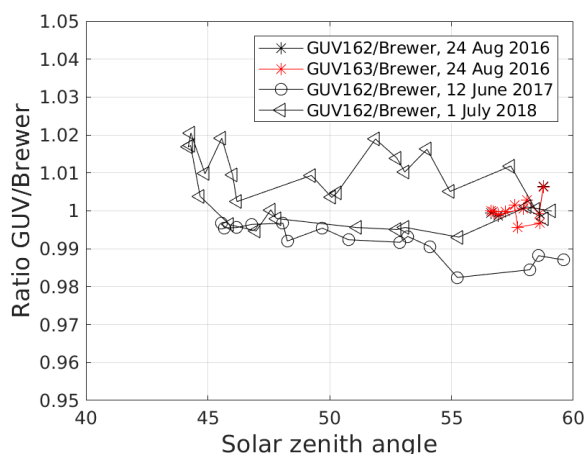


Figure 4. Ratio of the GUVs' and Brewer total ozone column at Sodankylä.

3.3 Solar comparison at Marambio

At Marambio, the first ~~switch~~ exchange of the GUV radiometers occurred in November 2018. The GUV #162 was set up next to the GUV #163, which had measured continuously for nearly two years at Marambio. The two instruments measured simultaneously for two weeks before the GUV #163 was packed and shipped for recalibration. The results of the comparison show a difference of 4 – 6% for SZAs smaller than 60° (Fig. 5). Such differences were expected, as it is common for filter radiometers that the wearing-in of Teflon diffuser introduces small changes in the response of channels during the first years of measurements. The drift was considered to be within the uncertainties of the measurements and the time series was not corrected for.

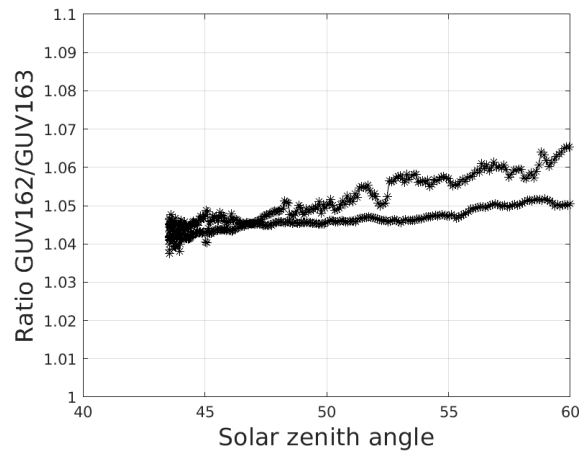


Figure 5. Ratio of the GUVs' UV index at Marambio on 25 November 2018.

4 First two years of measurements

The erythemally weighted UV, maximum UV index and daily average total ozone time series were discussed in Aun et al. (2019) and the results were compared to measurements from other Antarctic measurement sites. In this paper, the time series of the noon UV index, noon total ozone, UVB, UVA, VIS and PAR measurements for the period March 2017 – May 2019 are presented in Figs. 6 and 7 (~~Lakkala et al., 2019~~). Noon time UV index refers to the UV index measured at minimum SZA of the day. The maximum noon UV indices of the spring–autumn seasons were 6 and 9 in 2017–2018 and 2018–2019, respectively. ~~Noon-time UV index refers to the UV index measured at minimum SZA of the day.~~ In 2017–2018, the maximum noon UV index was measured both in the spring (October) and in the summer (January). In 2018–2019 the maximum was measured in the spring when the station was inside the polar vortex (Aun et al., 2019) and exceeded the summer noon maximum UV index which was 8. For the months October – February total ozone was ~~most of the days~~ lower in 2018–2019 than ~~in~~ 2017–2018 for most of the days (Fig. 6). This resulted in higher UV indices and UVB levels in 2018–2019 compared to ~~the period~~ 2017–2018. The effect of total ozone changes is more pronounced at shorter UVB wavelengths than at longer UVA wavelengths (Fig. 7), following the ozone absorption spectrum. Elevated UVB dose rates are detected in the spring and in the summer (Fig. 7c). The peak in October–November is due to the ozone hole, when the station is located below the stratospheric polar vortex. The reasons for the observed UV features are analyzed and discussed in more detail in Aun et al. (2019).

Variations in VIS and PAR are dominated by variations in cloudiness. Rapidly changing cloudiness is typical in Marambio (Aun et al., 2019), which is seen as high daily variation in all radiation measurements: UV, VIS and PAR. Surface albedo is higher during the spring months than during the summer months, which increases measured radiation in the spring at both shorter UV wavelengths and longer VIS and PAR wavelengths. This results in non-symmetric distribution of daily maxima; with higher spring time VIS radiation and PAR than in the autumn at similar SZAs. The time series of VIS and PAR radiation

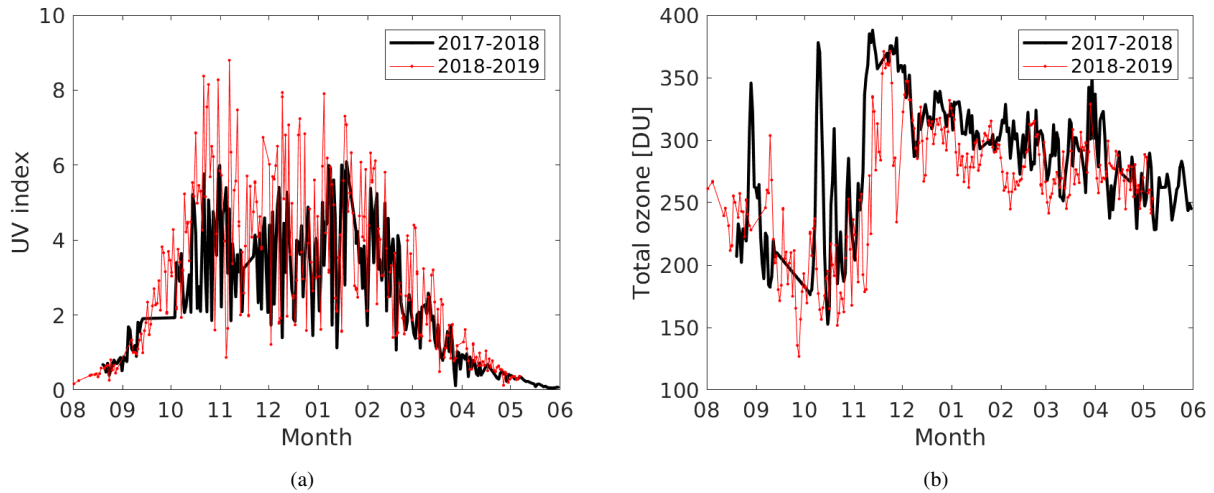


Figure 6. UV index (a) and total ozone (b) at noon calculated from the Marambio GUV radiometer measurements for the months August – May in 2017-2018 and 2018-2019.

suggest that there were more cloud-free ~~moments~~ periods around noon during mid-November – January in 2018–2019 than during the same months in 2017 – 2018, as peaks values were higher during the second year.

Table 4 summarizes the monthly means of the UVB and UVA daily doses, the UVB and UVA daily maximum dose rates, and the VIS and PAR daily maxima for September–March in 2017-2019. ~~The number of measurement days available for daily dose and daily maxima calculations are shown in Tables. ?? and ??, respectively.~~

~~The number of measurement days during 2017-2019 which were available for daily dose calculations. The months from January to December are numbered from 1 to 12. Year-Month-Number of days 2017-3-22 2017-4-8 2017-8-10 2017-9-13 2017-10-28 2017-11-25 2017-12-29 2018-1-26 2018-2-26 2018-3-31 2018-4-15 2018-8-9 2018-9-25 2018-10-30 2018-11-26 2018-12-31 2019-1-31 2019-2-28 2019-3-27 2019-4-25~~

~~The number of measurement days during 2017-2019 which were available for daily maximum calculations. The months from January to December are numbered from 1 to 12. Year-Month-Number of days 2017-3-22 2017-4-10 2017-5-29 2017-6-5 2017-8-12 2017-9-13 2017-10-28 2017-11-25 2017-12-29 2018-1-26 2018-2-26 2018-3-31 2018-4-23 2018-5-25 2018-6-21 2018-7-21 2018-8-18 2018-9-25 2018-10-30 2018-11-26 2018-12-31 2019-1-31 2019-2-28 2019-3-27 2019-4-30~~

Figure 8 shows ratios of UV irradiances at 305, 313, 320 and 380 nm compared to VIS irradiance at 555 nm in 2017–2018 and 2018–2019. For wavelengths which are affected by ozone absorption (mainly 305 and 313 nm), the spring time low ozone episodes are clearly identified as increased ratios (October and November). In the autumn, the radiation path in the atmosphere becomes longer day by day, so that the attenuation of UV radiation is stronger than that of VIS radiation. This is seen at short wavelengths as decrease of the UV/VIS ratio from summer to late autumn. The longer the atmospheric path is, the higher is

Table 4. Monthly means of GUV radiometer products measured at Marambio during the period March 2017 – April 2019. Months with more than 15 measurement days are included. One standard deviation is showed in the parenthesis. DD=daily dose, DMDR = daily maximum dose rate, DM=daily maximum. VIS is irradinace measured at 555 nm.

Jan	Feb	Mar	Sept	Oct	Nov
NaN	NaN	16.19(7.06)	NaN	33.55(11.03)	41.38(8.00)
42.62(14.72)	34.60(11.19)	14.09(5.73)	17.74(7.92)	44.04(12.55)	44.89(14.28)
45.86(15.25)	31.62(11.59)	16.36(9.14)	NaN	NaN	NaN
NaN	NaN	0.50(0.20)	NaN	0.78(0.25)	1.18(0.26)
1.04(0.37)	0.91(0.28)	0.45(0.14)	0.45(0.13)	0.89(0.22)	1.04(0.38)
1.08(0.36)	0.79(0.26)	0.51(0.24)	NaN	NaN	NaN
NaN	NaN	3.7637.64(1.3713.68)	NaN	7.6075.96(2.2522.54)	8.1681.56(1.2612.57)
9.2992.91(2.4724.68)	7.7577.46(2.0620.57)	3.3433.40(1.3613.61)	4.3243.24(1.8218.21)	10.17101.67(2.8027.97)	10.38103.80(2.9629.60)
9.7597.55(3.3433.39)	7.3773.70(2.6026.00)	3.6636.56(1.8418.39)	NaN	NaN	NaN
NaN	NaN	238.902.39(64.150.64)	NaN	317.743.18(69.280.69)	431.814.32(59.490.59)
412.454.12(117.061.17)	397.573.98(89.160.89)	235.242.35(65.990.66)	219.362.19(33.010.33)	357.973.58(77.560.78)	437.814.38(134.021.34)
430.914.31(141.641.42)	359.563.60(96.350.96)	240.692.41(81.820.82)	NaN	NaN	NaN
NaN	NaN	75.61(18.85)	NaN	96.16(22.23)	132.00(21.03)
120.14(39.14)	120.72(28.75)	76.18(23.40)	68.39(15.43)	104.92(24.77)	123.58(41.62)
123.36(47.74)	103.05(33.12)	72.73(26.88)	NaN	NaN	NaN
NaN	NaN	10.310.10(2.570.03)	NaN	13.130.13(3.040.03)	18.080.18(2.870.03)
16.470.16(5.380.05)	16.450.16(3.800.04)	10.430.10(3.170.03)	9.870.10(1.440.01)	14.360.14(3.380.03)	17.010.17(5.730.06)
16.490.16(6.260.06)	14.060.14(4.410.04)	9.720.10(3.610.04)	NaN	NaN	NaN

the path-integrated total ozone concentration which can absorb UV radiation. The opposite is seen in the spring ratios. UV-A radiation, irradiances at 320 and 380 nm, is only partially absorbed by ozone, and the effect is smaller.

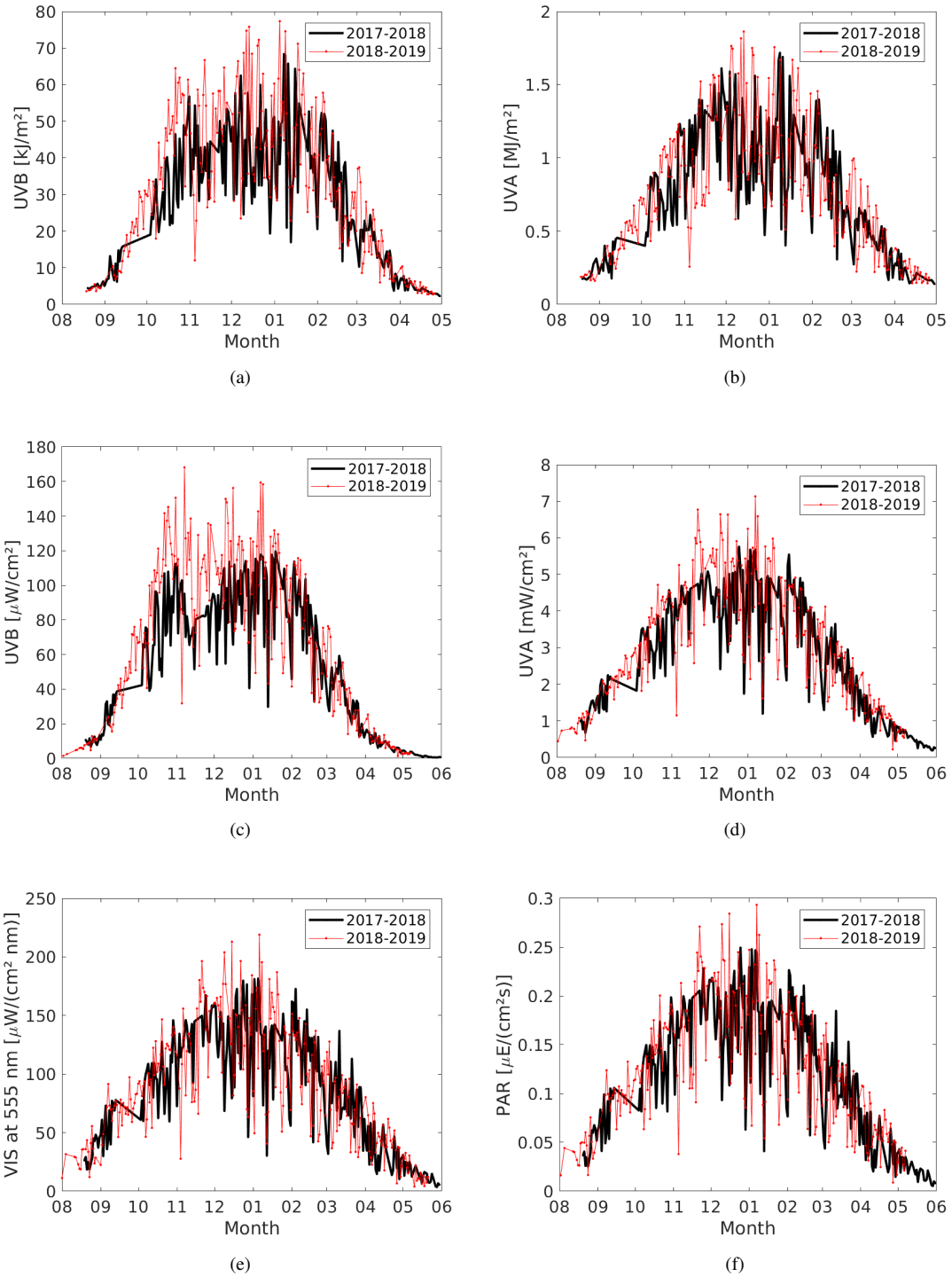


Figure 7. UV and visible radiation time series for the months August – May during periods 2017-2018 and 2018-2019. A) UVB daily dose, b) UVA daily dose, c) UVB daily maximum, d) UVA daily maximum, e) daily maximum VIS at 555 nm and f) PAR daily maximum.

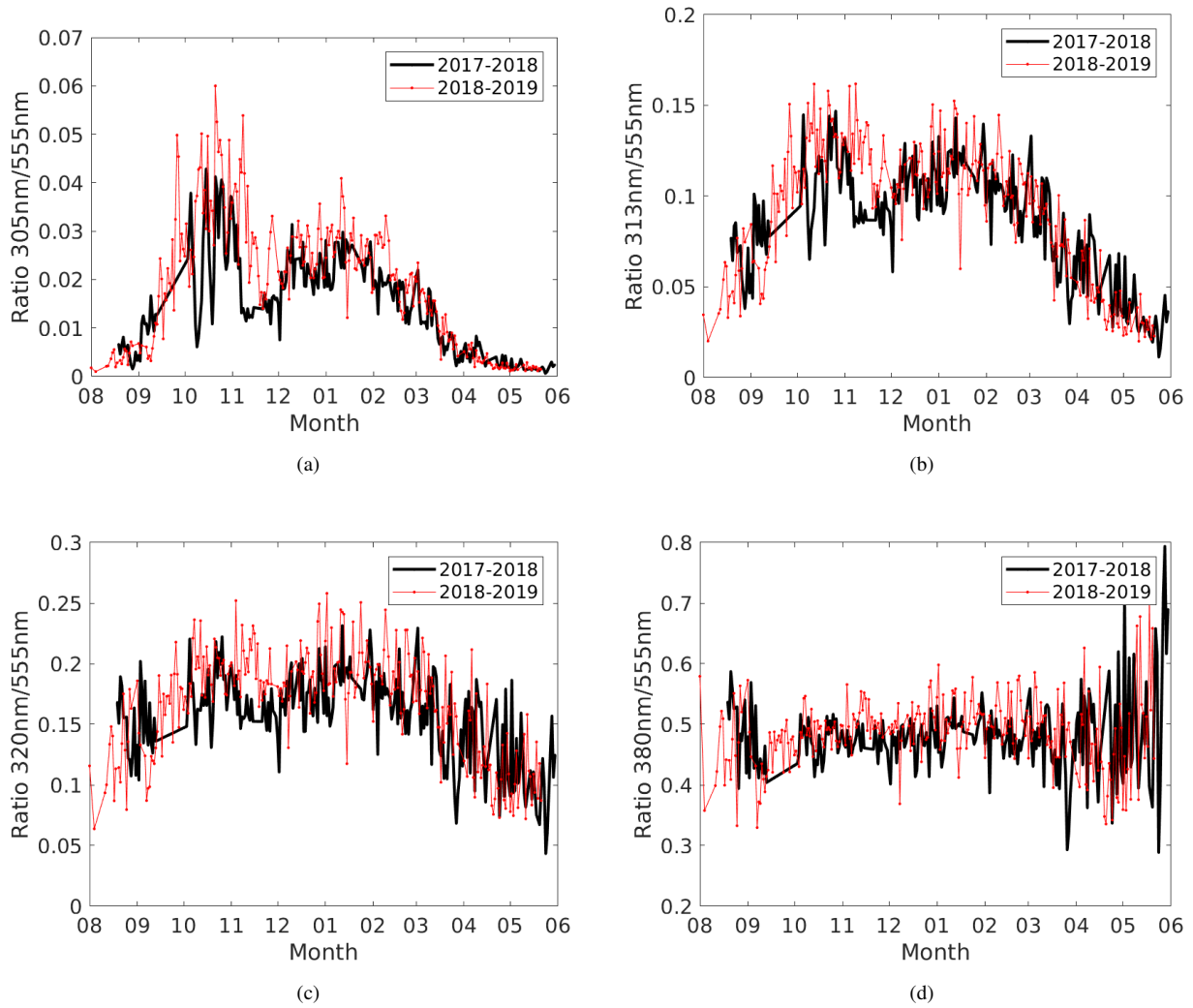


Figure 8. Ratio of daily maximum UV irradiances and VIS irradiances for the months August – May during periods 2017-2018 and 2018-2019. UV irradiances at A) 305 nm, b) 313 nm, c)320 nm and d) 380 nm.

5 Data availability

The studied datasets of ~~UV irradiances, visible and photosynthetically active radiation~~ daily maximum irradiances at five UV channels and one VIS channel, daily maximum UVB/UVA and PAR dose rates, noon UVB/UVA and PAR dose rates, noon total column ozone and UVB/UVA daily doses are freely available at Zenodo, (~~Lakkala et al., 2019~~), <http://doi.org/10.5281/zenodo.3688700> (Lakkala et al., 2019). The additional datasets including biologically weighted dose rates and daily doses are available from the authors.

6 Discussion and Conclusions

Marambio's new GUV multifilter radiometer measurements cover the spectral range from the UV to the VIS, and thus can be used for assessing the spectral effects of changes in factors affecting the amount of solar UV and VIS radiation at the surface: aerosols, albedo, cloudiness and total ozone. Those factors are also influenced by climate change, and thus the measurements serve as important tools for assessing effects of ~~climate change~~ a changing climate on the Antarctic environment.

The recovery of the stratospheric ozone layer has started (WMO, 2018b) and the GUV total ozone measurements will be used to assess features of the recovery in Antarctica. The signs of the recovery can be detected by the UV channels of the radiometer, as actual measurements can be compared with the time series of the Antarctic-NILU-UV network and other UV measurements performed in Antarctica (Aun et al., 2019). The time series of the first two years of GUV measurements show that the recovery is not seen as linear increase in UV levels, and clear differences exist between 2017–2018 and 2018–2019: the effect of the spring time ozone depletion was more pronounced in 2018–2019 than in 2017–2018 with maximum noon UV indices of 9 and 6, respectively. This is however far from the maximum UV index measured within the Antarctic NILU-UV network during the years 2000–2010, when ~~maximum UV index~~ a UV index of 12 was measured at Marambio in the spring (November) 2007 (Lakkala et al., 2018b). For comparison: at Arctic sites the spring time stratospheric ozone loss doesn't increase UV levels to reach the yearly maxima, even during years with severe ozone depletion (Bernhard et al., 2013). In the Arctic, UV index maxima are observed in the summer. For example in Sodankylä, Finland, which is a site affected by stratospheric ozone loss, a maximum UV index of 6 has been measured in the summers of the years 2011 and 2013 (Lakkala et al., 2016).

The Antarctic UV measurements of Marambio can be directly compared to the Arctic UV measurements using the results from solar comparisons between Marambio's GUV radiometer, PMOD-WRC and FMI's spectroradiometers in Sodankylä. The results show that the measurements agreed within $\pm 6\%$ for SZA < 60°. These yearly comparisons with Arctic measurements are possible because two GUVs have been purchased for Marambio's measurements: one is measuring at Marambio, while the other is participating in solar comparisons. In addition, regular absolute calibration against high-quality spectroradiometers is part of the quality assurance. The first re-calibration occurred in the spring of 2019. Both solar comparisons and regular absolute calibrations are crucial to obtain homogenized long-term time series (Johnsen et al., 2008), which can serve multidisciplinary research communities. In addition, these high-quality measurements can be used for validation of satellite data, e.g., validation

of UV products based on TROPOMI measurements on board the Sentinel-5 Precursor satellite (Lindfors et al., 2018). They are highly valuable, as the ground based measurement network is very sparse at high latitudes.

Ultraviolet radiation products calculated from Marambio's GUV measurements include UV dose rates and doses calculated using 10 different biological action spectra related to UV effects on, e.g., skin, plant, anchovy and phytoplankton. All these UV products can be used when assessing the effects of climate change on the whole Antarctic ecosystem. The effects are not self-evident even though the melting of ice sheet and snow (e.g., Shepherd et al., 2018) due to temperature increase (Steig et al., 2009) is well recognized. The uncertainty is mostly due to complicated atmosphere-sea-land feedback mechanisms, (e.g., Wang et al., 2019). In future, GUV's VIS radiation time series can be used together with ceilometer measurements and synoptic cloud observations to assess the impact of climate change on cloudiness. Together with the ongoing albedo and aerosol measurements in Marambio they form a complex measurement system for atmospheric radiation studies.

Author contributions. K. Lakkala: Primarily responsible for the QA of the UV data, analyzed the data and led the manuscript preparation.

M. Aun: Programmed GUV data processing for the Marambio station, participated in data [analyze-analysis](#) and contributed to the writing of the manuscript.

R. Sanchez: Responsible in SMN for Marambio UV measurements. Contributed to the writing of the manuscript.

15 G. Bernhard: Performed the absolute calibration of the GUV radiometers, programmed UV data processing program and contributed to the writing of the manuscript.

E. Asmi: Data [analyze-analysis](#) of Marambio cloudiness data. Contributed to the writing of the manuscript.

O. Meinander: Responsible for the UV albedo measurements at Marambio. Contributed to the writing of the manuscript.

F. Nollas: Data [analyze-analysis](#) of Marambio cloudiness data. Contributed to the writing of the manuscript.

20 G. Hülsen: Performed the QASUME site audit in Sodankylä in 2018. Processed QASUME reference UV data.

[T. Karppinen: Sodankylä Brewer ozone processing and QC/QA. Contributed to the writing of the manuscript.](#)

V. Aaltonen: Responsible for the PFR measurements at Marambio. Contributed to the writing of the manuscript.

A. Arola: Overseeing the work in the group and contributed to the writing of the manuscript.

G. de Leeuw: Group leader during the set up of Marambio UV measurements. Contributed to the writing of the manuscript.

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