



# 1 A high-quality hourly, daily and monthly solar irradiance 2 dataset in China during 1981-2014 based on MERRA-2 3 Reanalysis products

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10 **Abstract.** Solar irradiance (SI) is the main driving factor contributing to climate change and energy  
 11 balance between the land and atmosphere. High-quality records of global solar irradiance (GHI), direct  
 12 normal irradiance (DNI) and diffuse solar irradiance (DIF) are of vital importance for solar applications,  
 13 but the solar radiation observations are sparse around the world. As an alternative, numerous SI reanalysis  
 14 data in grid format have been developed in regional and global scales. Among them, the MERRA-2  
 15 (Modern-Era Retrospective Analysis for Research and Applications, version 2) products could provide  
 16 high quality SI records with acceptable accuracy and long temporal ranges. This study attempted to  
 17 improve the accuracy of GHI records derived from MERRA-2 products, and to generate grid DNI and  
 18 DIF datasets for all-sky conditions over mainland China during 1981-2014, based on the REST2 model  
 19 and cloud transmittance estimates combining sunshine observations. The results indicate that the  
 20 estimated GHI values (GHI<sub>new</sub>) show higher agreements with GHI measurement at 17 CMA (China  
 21 meteorological administrations) stations than that for the GHI records derived from MERRA-2 products  
 22 (MERRA-2 GHI). Then, grid GHI, DNI and DIF datasets (0.50° (lat) \* 0.625° (lon)) throughout China  
 23 were constructed. The results indicated that the MERRA-2 GHI records may overestimate the GHI values  
 24 over mainland China. Generally, the GHI and DNI values gradually decreased during 1981-2014,  
 25 however, DIF values gradually increased from 1981 to 2014, especially in 1992 (DIF = 90.914 Wm<sup>-2</sup>,  
 26 anomaly DIF value = 15.544 Wm<sup>-2</sup>). The Qinghai Tibetan Plateau has always been an area with the  
 27 highest GHI, the highest DNI and the lowest DIF values, whereas the Sichuan Basin has always been an  
 28 area with the lowest GHI, the lowest DNI and the highest DIF values. The grid GHI, DNI and DIF dataset  
 29 generated in this study can assist in numerous solar studies and applications. We provide these solar



30 irradiance data in publicly available repository: <https://doi.org/10.6084/m9.figshare.10026563> (Qin, W.  
31 et al., 2019).

32 **Keywords:** global horizontal solar radiation; direct normal irradiance; diffuse horizontal solar radiation;  
33 MERRA-2; China

## 34 1 Introduction

35 Solar energy is a clean, renewable and sustainable energy source for solar energy applications such  
36 as photovoltaic energy utilization (Besharat et al. 2013; Purohit and Purohit 2015). China has the largest  
37 thermal power generation of any country around the world, making it the largest emitter of greenhouse  
38 gases (Amadei et al. 2013). The large demand for electricity and energy consumption has caused the  
39 Chinese government to vigorously develop the concentrated solar thermal (CST) CSP industry (Li et al.  
40 2014). China is also in the leading position regarding the construction and planned installed capacity of  
41 CSP power generation around the world (Zhao et al. 2017). Therefore, accurate measurement of solar  
42 irradiance (SI) is the basis and prerequisite for effective utilization of solar radiation resources (Qin et al.  
43 2019).

44 Many observation networks have been constructed for providing SI records (in point format) in  
45 China. The Baseline Surface Radiation Network (BSRN, Zhang et al. 2015), the World Radiation Data  
46 Center (WRDC, Zhang et al. 2017), and the Global Energy Balance Archive (GEBA, Wild et al. 2017)  
47 can provide SI records covering more than 2,000 observation stations around the world. In China,  
48 according to the statistics of China Meteorological Data Network, there have been 122 solar radiation  
49 measurement stations installed since 1957. In the 1990s, there were only 96 stations for solar radiation  
50 measurement throughout China (Zou et al. 2017). However, because of the high cost of the site  
51 construction and observation instruments, especially in remote areas with poor natural conditions, these  
52 SI observation networks are still too scarce to support solar energy research and applications in China  
53 (Qin et al. 2018).

54 In contrast with solar radiation observation stations, there are thousands of meteorological stations  
55 covering mainland China. Thus, many studies have been conducted to construct GHI, DNI and DIF  
56 datasets (in point format) in China using meteorological measurements at CMA stations. Tang et al. (2018)  
57 first constructed a direct solar radiation data set in China with acceptable accuracy and high point density



58 (2474 CMA stations). Chen et al. (2019) applied five artificial intelligence models and one broadband  
59 model for estimating direct solar radiation. The direct solar radiation results could be converted to DNI  
60 by multiplying by the solar zenith angle. Feng et al. (2018) evaluated 15 empirical models for predicting  
61 DIF values at 17 CMA radiation stations. However, the drawback to these solar radiation estimations is  
62 that there are few DNI estimation results in Western China, especially in the Qinghai Tibetan Plateau due  
63 to the extremely sparse meteorological measurements in Western China. Thus, SI records in grid format  
64 covering mainland China with high spatial and temporal resolutions are urgently needed for solar  
65 research and solar energy applications in China.

66 Numerous SI products in grid format have been created providing grid GHI, DNI and DIF records  
67 with high spatial and temporal continuity covering mainland China; for example, the Global Energy and  
68 Water Exchanges-Surface Radiation Budget Project (GEWEX-SRB, Raschke et al., 2006) and the  
69 International Satellite Cloud Climatology Project-flux data (ISCCP-FD, Lohmann et al., 2006) can  
70 provide solar radiation records throughout China with spatial resolution of  $1^{\circ} \times 1^{\circ}$ . SI can also be derived  
71 from the GEDEX (Greenhouse Effect Detection Experiment) products developed by the NCAS British  
72 Atmospheric Data Centre (NCAS BADC) (Sinha and Shine 1995). The Climate Data Record (CDR)  
73 generated by NOAA can provide GHI records in China with long temporal ranges (1882-2019)  
74 (Coddington et al. 2016). However, these products still cannot meet the requirements of solar energy  
75 research in China needing GHI, DNI and DIF records with high accuracy and spatial resolution (Qin et  
76 al. 2015).

77 Remote sensing is an alternative method to obtain GHI, DNI and DIF values in China with high  
78 spatial resolution. GHI, DNI and DIF records could be derived from HelioClim (Blanc et al. 2011), MSG,  
79 Meteosat (Möser and Raschke 1984), GOES (Gautier et al. 1980), MODIS (Qin et al. 2011), Himawari  
80 (Bessho et al. 2016), and CM-SAF SARA (Riihelä et al. 2015) observations. However, the accuracy of  
81 these GHI, DNI and DIF records needs to be improved. Shi et al. (2018) evaluated the accuracy of the  
82 estimated GHI values derived from the Advanced Himawari Imager (AHI) aboard Himawari-8 at 36  
83 CERN (Chinese Ecosystem Research Network) stations. The results show that the GHI estimations did  
84 not show good agreement with GHI measurements. Thus, many scientists have developed efficient  
85 algorithms to improve the quality of solar radiation estimations in China using satellite images. Wei et  
86 al. Wei et al. (2019) compared the accuracy of the estimated GHI values over China based on four



87 different AI models using AVHRR data, and they analyzed the spatial and temporal variations in GHI  
 88 over mainland China. Qin et al. (2015) developed an efficient physical parameterization (EPP) for  
 89 estimating GHI values using MODIS land and atmospheric products and evaluated the EPP model at 91  
 90 CMA stations in China. However, the spatial resolution ( $1^{\circ}\times 1^{\circ}$ ) and spatial continuity of the estimation  
 91 results by the EPP model could not meet the requirements of solar energy research, which requires SI  
 92 records with high spatial resolution. Tang et al. (2016a) further improved the EPP model (EPP-TANG) by  
 93 combining the MODIS and MTSAT products in China. The spatial resolution of GHI estimations have  
 94 been improved to a 5 km spacing, but the spatial continuity of GHI estimations still restrict the  
 95 applicability of the EPP-TANG model in China. Liang et al. (2006) developed an efficient model based  
 96 on the look-up table method and the atmospheric radiation transfer model for incident GHI using MODIS  
 97 products. Zhang et al. (2014) further generated GHI, DNI and DIF products called GLASS (Global Land  
 98 Surface Satellite) covering China. Nevertheless, EPP, EPP-TANG and GLASS could only generate  
 99 instantaneous solar radiation values, thus they could not provide accurate daily GHI, DNI and DIF  
 100 records over mainland China (Tang et al. 2016b).

101 Reanalysis data is an alternative SI data source with acceptable accuracy and high spatiotemporal  
 102 continuity covering mainland China (Rienecker et al. 2011). ERA5 is the fifth generation of ECMWF  
 103 atmospheric reanalysis global climate data providing hourly and daily surface downward solar radiation  
 104 records from 1979 to present (Babar et al. 2019). SI values could also be derived from NCEP-DOE  
 105 AMIP-II reanalysis (Kanamitsu et al. 2002). The CRU JRA V2.0 dataset is also a data source with an  
 106 hourly downward solar radiation flux (Beck et al. 2017). The Climate Forecast System Reanalysis (CFSR)  
 107 developed by the National Oceanic and Atmospheric Administration (NOAA) could provide solar  
 108 radiation records from 1979 to present (Fuka et al. 2014). Using the GEOS-5 atmospheric general  
 109 circulation model (AGCM), the Modern-Era Retrospective Analysis for Research and Application  
 110 (MERRA) was stimulated by the National Aeronautics and Space Administration (NASA) Global  
 111 Modeling and Assimilation Office (GMAO), which could provide hourly, daily and monthly GHI records  
 112 during 1980-2019 at global scales (Bosilovich et al. 2011). The GHI values derived from MERRA-2  
 113 products were demonstrated to have good agreement with GHI measurements (Hodges et al. 2011;  
 114 Kennedy et al. 2011). Thus, the updated version (MERRA-2) was developed with numerous  
 115 improvements (Randles et al. 2017). In this study, it was supposed that the accuracy of GHI records in



116 MERRA-2 could be improved by integrating the effects of cloud transmittances. Moreover, the DNI and  
 117 DIF records are missing in previous MERRA-2 reanalysis data.

118 In what follows, GHI, DNI and DIF measurements at Wuhan station, Xianghe station, 40 CERN  
 119 stations and 17 first-class CMA meteorological stations throughout mainland China are used to evaluate  
 120 the performance of the estimated solar irradiance (GHI, DNI and DIF) values and GHI records derived  
 121 from MERRA-2 products during 1993–2014. In a subsequent step, the GHI, DNI and DIF databases  
 122 throughout mainland China are constructed using MERRA-2 products and sunshine duration  
 123 measurements at 2474 CMA stations. Finally, the spatiotemporal variations and possible influencing  
 124 factors on GHI, DNI and DIF over different climate zones and terrains in mainland China are investigated.  
 125 Overall, this study should prove helpful in solar resource and energy applications that need long-term  
 126 grid GHI, DNI and DIF data with moderate spatiotemporal resolution and acceptable accuracy. We  
 127 provide these solar irradiance data in publicly available repository (Qin, W. et al., 2019).

## 128 **2 Materials and methods**

### 129 **2.1 Sites and data processing**

130 Hourly GHI, DNI and DIF measurements at Xianghe station (BSRN) in China were used for  
 131 calculated the cloud transmittances for surface global horizontal solar radiation (GHI), direct normal  
 132 irradiance (DNI) and diffuse horizontal solar radiation (DIF). Hourly GHI measurements from 40 CERN  
 133 stations, hourly DNI and DIF measurements at Wuhan station (in Wuhan university), and Daily GHI,  
 134 DNI and DIF measurements during 1993-2014 at 17 CMA stations in China were used for evaluating the  
 135 model accuracy of the estimated hourly, daily and month GHI, DNI and DIF values generated in this  
 136 study. Meanwhile, the sunshine duration measurements during 1981-2014 that were routinely measured  
 137 at 2474 CMA stations over mainland China were also used to calculate the cloud transmittance of the  
 138 GHI, DNI and DIF. These meteorological data have been checked for data quality using various control  
 139 methods.

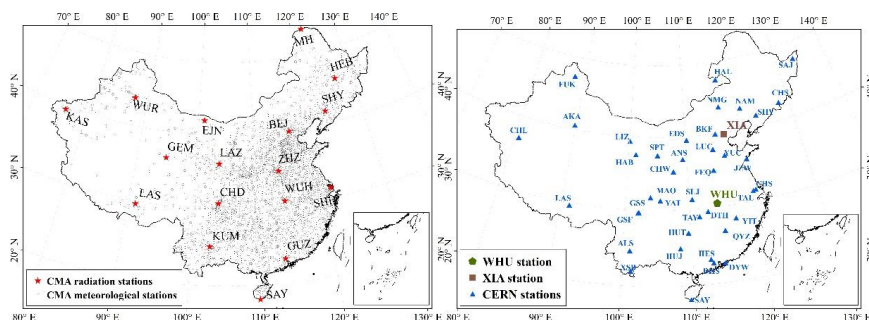


Figure 1. Spatial distributions of the CMA stations that are used in this study.

Figure 1 show the spatial distributions of the Wuhan station, Xianghe station and CMA meteorological stations that were used in this study. These stations covered most areas of China with distinct climatic and terrain features.

## 2.2 REST2 Model

REST2 is a physically based model for predicting hourly and daily broadband GHI, DNI and DIF values in clear sky conditions, which was first developed by (Gueymard 2003), then corrected and modified by (Gueymard 2012) (REST to REST2). REST2 has been validated as one of the best broadband solar radiation estimation models and has been widely used in numerous solar radiation research (Gueymard 2003). The REST2 model has corrected the diffusion calculation under low-AOD, near-Rayleigh conditions in the model. The GHI, DNI and DIF values in REST2 can be obtained using following equations:

$$DNI = \tau_R \tau_g \tau_o \tau_n \tau_w \tau_a E_0 \quad (1)$$

$$DIF = 0.5 \tau_g \tau_o \tau_n \tau_w (1 - \tau_a \tau_R) E_0 \quad (2)$$

$$GHI = DNI * \cos(\theta) + DIF \quad (3)$$

where  $\tau_R, \tau_g, \tau_o, \tau_n, \tau_w$  and  $\tau_a$  are the transmittances for Rayleigh scattering, uniformly mixed gases absorption, ozone absorption, nitrogen dioxide absorption, water vapor absorption and aerosol extinction, respectively.  $E_0$  is the extraterrestrial solar radiation.  $\theta$  is the solar zenith angle. These transmittances have been obtained accurately by fitting a large number of parametric runs of the SMARTS code to computationally efficient polynomial ratios (Gueymard 2012).



158        Considering the data availability, the hourly reanalysis meteorological records derived from the  
159        MERRA-2 dataset during 1981-2014, including aerosol optical depth in band 550 ( $AOD_{550}$ ), regional  
160        ground albedo ( $\rho_{\text{g}}$ ), air pressure ( $p$ ) and precipitable water vapor ( $w$ ) were used as model inputs for  
161        REST2. The spatial resolution of the MERRA-2 dataset that was used in this study is  $0.50^\circ$  (lat)  $\times$   $0.625^\circ$  (lon).  
162        More detailed descriptions and resulting equations of the REST2 model can be found in Ref (Gueymard  
163        2008, 2012).

### 164        **2.3 Anusplin**

165        The sunshine duration measurements during 1981-2014 at 2474 CMA stations throughout mainland  
166        China were used to calculate the cloud transmittances of GHI, DNI and DIF values. However, these CMA  
167        stations are still too sparse to support the solar radiation estimations in this study. Therefore, using the  
168        sunshine durations measurements at CMA stations, grid sunshine duration data ( $0.50^\circ$  (lat)  $\times$   $0.625^\circ$  (lon))  
169        during 1981-2014 over mainland China were generated based on the Anusplin tool. The ANUSPLIN  
170        package provides a facility for transparent analysis and interpolation of noisy multivariate data using  
171        thin-plate smoothing splines, comprehensive statistical analyses, data diagnostics and spatially  
172        distributed standard errors (Xu and Hutchinson 2013). The flowchart in the Anusplin tool was shown in  
173        Figure 2. A detailed description of the Anusplin tool could be found in Ref (Hutchinson and Xu 2004).  
174

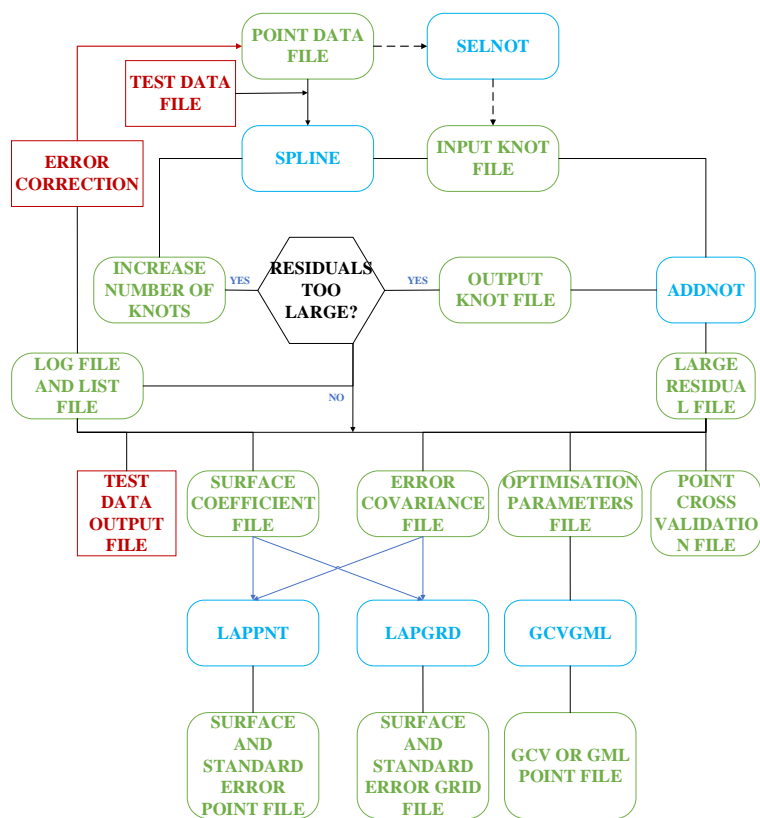


Figure 2. The flowchart of the processes of Anusplin tool.

## 2.4 Comparisons of measures of fit

In this study, 16 indicators were used to evaluate the model accuracy (Gueymard 2014).  $N$  and  $\bar{\cdot}$ , respectively, indicate the number of data and mean of the variables;  $e_i$  and  $o_i$  are the modeled and observed GHI, DNI and DIF values. These indicators are divided into four classes: Class A-indicators of dispersion, Class B-indicators of overall performance, Class C-indicators of distribution similitude and Class D-a global performance indicator.

### 2.4.1 Class A – indicators of dispersion

The Class A indicators are the root mean square error (RMSE), the mean absolute bias error (MAE), the relatively root mean square error (RMSD), the relatively mean absolute bias error (MAD), the correlation coefficient  $R$ , the standard deviation (SD), the slope of best-fit line (SBF), the uncertainty at 95% (U95), and the t-statistic (TS), which can be expressed as:





$$RMSE = \sqrt{\frac{1}{N} \sum_{i=1}^N (e_i - o_i)^2} \quad (4)$$

$$MAE = \frac{1}{N} \sum_{i=1}^N |e_i - o_i| \quad (5)$$

$$RMSD = \frac{100}{\bar{o}_i} \times \sqrt{\frac{1}{N} \sum_{i=1}^N (e_i - o_i)^2} \quad (6)$$

$$MAD = \frac{100}{\bar{o}_i} \times \frac{1}{N} \sum_{i=1}^N |e_i - o_i| \quad (7)$$

$$R = \frac{\sum_{i=1}^N (e_i - \bar{o})(o_i - \bar{o})}{\sqrt{\sum_{i=1}^N (e_i - \bar{o})^2} \sqrt{\sum_{i=1}^N (o_i - \bar{o})^2}} \quad (8)$$

$$SD = \frac{100}{\bar{o}} \times \left( \frac{\sum_{i=1}^N N(e_i - \bar{o})^2 - \left( \sum_{i=1}^N (e_i - \bar{o}) \right)^2}{N} \right) \quad (9)$$

$$SBF = \frac{\sum_{i=1}^N (e_i - \bar{o})(o_i - \bar{o})}{\sum_{i=1}^N (o_i - \bar{o})^2} \quad (10)$$

$$U95 = 1.96 \sqrt{(SD^2 + RMSD^2)} \quad (11)$$

$$TS = \sqrt{(N-1)MBD^2 / (RMSD^2 - MBD^2)} \quad (12)$$

## 188 2.4.2 Class B – indicators of overall performance

189 The Class B indicators are the Nash–Sutcliffe’s efficiency (NSE), the Willmotts’s index of agreement  
 190 (WIA) and the Legates’s coefficient of efficiency (LCE), which can be expressed as:

$$NSE = 1 - \frac{\sum_{i=1}^N (e_i - o_i)^2}{\sum_{i=1}^N (o_i - \bar{o})^2} \quad (13)$$

$$WIA = 1 - \frac{\sum_{i=1}^N (e_i - o_i)^2}{\sum_{i=1}^N (|e_i - o_i| + |o_i - \bar{o}|)^2} \quad (14)$$

$$LCE = 1 - \frac{\sum_{i=1}^N |e_i - o_i|}{\sum_{i=1}^N |o_i - \bar{o}|} \quad (15)$$

## 191 2.4.3 Class C – indicators of distribution similitude

192 The Class C indicators are the Kolmo-gorov–Smirnovtest Integral (KSI), the relative frequency of  
 193 exceedance (OVER) and combined performance index (CPI), which can be expressed as:

$$KSI = \frac{100}{A_c} \int_{x_{min}}^{x_{max}} D_n dx \quad (16)$$

$$A_c = D_c (X_{max} - X_{min}) \quad (17)$$

$$D_c = \phi(N) / \sqrt{N} \quad (18)$$

$$OVER = \frac{100}{A_c} \int_{x_{min}}^{x_{max}} Max(D_n - D_c, 0) dx \quad (19)$$

$$CPI = (KSI + OVER + 2RMSE) / 4 \quad (20)$$

194 where  $D_n$  is the absolute difference between the two normalized distributions within irradiance interval  
 195  $n$ ,  $x_{min}$  and  $x_{max}$  are the minimum and maximum values of the binned reduced irradiance,  $x$ , and  $A_c$  is a  
 196 characteristic quantity of the distribution. Detail descriptions of KSI, OVER and CPI indicators could be



197 found in Ref (Gueymard 2014).

#### 198 **2.4.4 Class D –a global performance indicator (GPI)**

199 Although 16 indicators are introduced to reveal the model accuracy of GHI, DNI and DIF values,  
 200 too many indicators cannot reflect the overall accuracy of the estimated GHI, DNI and DIF values. In  
 201 this study, a global performance indicator (GPI) is used to represent the global performance of the  
 202 estimated GHI, DNI and DIF values (Despotovic et al. 2015). The GPI can be described by following  
 203 equation:

$$GPI_i = \sum_{j=1}^n a_j (\tilde{y} - y_{ij}) \quad (21)$$

204 where  $\tilde{y}$  is the median of the scaled values of indicator  $j$ ,  $y_{ij}$  is the scaled value of indicator  $j$  for model  $i$ ,  
 205 and  $n$  is the number (16) of indicators.  $a_j$  equals -1 for R, SBF, NSE, WIA and LCE, and equals 1 for  
 206 other indicators. The greater the accuracy of the model, the higher the value of the GPI.

### 207 **3 Result**

#### 208 **3.1 Cloud transmittance for surface solar irradiance**

209 Due to the shape, type and phase variability in clouds, they have been considered to be the most  
 210 uncertain factor in estimating SI. In this study, the relative sunshine duration, defined as the ratio between  
 211 the measured sunshine duration and the maximum possible sunshine duration, ( $N$ ) was introduced to  
 212 correct the cloud effect on hourly GHI, DNI and DIF values. Following the example of the Ångström-  
 213 Prescott equation, we parameterized the cloud transmittance ( $\tau_c$ ) as a function of the relative sunshine  
 214 duration ( $n/N$ ), and the formula form was a quadratic polynomial formulation as follows:

$$\tau_c = \frac{R}{R_{ctr}} = a + b \left(\frac{n}{N}\right) + c \left(\frac{n}{N}\right)^2 \quad (22)$$

215 where  $R$  is the hourly and daily all-sky GHI, DNI and DIF;  $R_{ctr}$  is the hourly and daily clear-sky GHI,  
 216 DNI and DIF; and  $n$  and  $N$  are the sunshine duration and the maximum possible sunshine duration,  
 217 respectively. The calibrated cloud transmittance for hourly GHI/DNI/DIF values are shown as the  
 218 following equations:

$$\tau_{c1} = 0.368 + 0.628 \left(\frac{n}{N}\right) - 0.005 \left(\frac{n}{N}\right)^2 \quad (23)$$

$$\tau_{c2} = 0.035 + 0.331 \left(\frac{n}{N}\right) + 0.298 \left(\frac{n}{N}\right)^2 \quad (24)$$

$$\tau_{c3} = 0.752 + 2.396 \left(\frac{n}{N}\right) - 2.029 \left(\frac{n}{N}\right)^2 \quad (25)$$

219 where  $\tau_{c1}$ ,  $\tau_{c2}$  and  $\tau_{c3}$  are the cloud transmittance formula for hourly GHI, DNI and DIF values,



220 respectively.

221 The calibrated cloud transmittance for daily GHI/DNI/DIF values are shown as the following

222 equations:

$$\tau_{c1} = 0.280 + 0.954 \left( \frac{n}{N} \right) - 0.299 \left( \frac{n}{N} \right)^2 \quad (26)$$

$$\tau_{c2} = 0.024 + 0.227 \left( \frac{n}{N} \right) + 0.619 \left( \frac{n}{N} \right)^2 \quad (27)$$

$$\tau_{c3} = 0.959 + 4.115 \left( \frac{n}{N} \right) - 4.232 \left( \frac{n}{N} \right)^2 \quad (28)$$

223 where  $\tau_{c1}$ ,  $\tau_{c2}$  and  $\tau_{c3}$  are the cloud transmittance formula for daily GHI, DNI and DIF values,

224 respectively.

### 225 3.2 Validation of the estimated GHI, DNI and DIF at CMA stations

226 Hourly GHI, DNI and DIF measurements at Wuhan stations, Xianghe stations and 40 CERN stations

227 were used to validate the accuracy of the estimated hourly GHI, DNI and DIF values. Daily GHI, DNI

228 and DIF measurements during 1993-2014 at 17 CMA meteorological stations are used for evaluating the

229 model accuracy of the daily estimated GHI, DNI and DIF values. The GHI records derived from

230 MERRA-2 products are also compared with the estimated GHI values in this study.



Table 1. Validation results of the estimated hourly mean GHI values at Xianghe and Wuhan station

Stations	Solar irradiance without cloud transmittances						Solar irradiance with cloud transmittances					
	WHU						XIA					
Value	GHI	DNI	DIF	GHI	DNI	DIF	GHI	DNI	DIF	GHI	DNI	DIF
RMSE	280.52	329.55	208.81	140.90	283.13	70.46	129.30	177.60	197.13	57.64	125.54	58.17
MAE	204.56	288.19	170.70	64.34	162.46	33.04	97.33	133.82	163.57	27.11	55.20	28.14
RMSD	56.45	68.38	113.05	62.80	107.12	94.12	41.16	83.00	91.89	36.22	85.27	64.01
MAD	41.16	59.79	92.42	28.67	61.47	44.13	30.98	62.53	76.25	17.04	37.49	30.96
MBD	-32.08	-59.60	-22.68	-28.34	-60.50	18.57	7.43	-9.01	-33.43	1.05	-29.09	-2.32
SD	56.44	68.37	113.04	62.80	107.12	94.11	41.15	82.99	91.88	36.22	85.26	64.01
R	0.62	0.63	0.08	0.91	0.69	0.85	0.88	0.47	0.36	0.97	0.86	0.89
U95	156.46	189.52	313.34	174.07	296.92	260.87	114.08	230.05	254.69	100.40	236.34	177.42
TS	69.18	178.14	20.51	81.26	109.97	32.34	18.39	10.93	39.11	4.64	58.32	5.83
NSE	-0.08	-1.53	-4.79	0.77	0.22	0.51	0.61	0.20	-2.39	0.93	0.70	0.78
WIA	0.90	0.61	0.97	0.98	0.91	0.99	1.00	0.99	0.93	1.00	0.99	1.00
LCE	0.11	-0.68	-1.46	0.75	0.43	0.60	0.43	0.21	-0.92	0.85	0.70	0.72
KSI	225.49	519.13	403.10	124.69	273.35	85.14	147.75	288.85	380.42	56.41	112.71	72.69
OVER	216.55	507.73	390.96	106.48	254.71	46.70	118.96	282.48	361.80	24.11	82.18	51.84
CPI	138.73	290.90	255.04	89.19	185.57	80.02	87.26	184.33	231.50	38.24	91.36	63.14
OM	496.96	481.98	184.71	224.37	264.30	74.87	314.18	213.99	214.53	159.12	147.23	90.88
PM	337.53	194.71	142.82	160.78	104.40	88.77	337.53	194.71	142.82	160.78	104.40	88.77
GPI	0.003	-4.658	-5.468	2.452	-2.042	1.185	2.876	-0.923	-3.260	4.872	1.694	3.269



Table 1 illustrate the statistical indicators representing the model accuracy of the estimated hourly GHI, DNI and DIF values at Wuhan and Xianghe station. The result indicated that the estimated hourly GHI, DNI and DIF values show high agreements with the hourly GHI, DNI and DIF measurements. The cloud has obvious effect on the accuracy of the estimated solar irradiance values. The modeling accuracy have been significantly improved after incorporating the cloud transmittances for solar irradiance. The GPI scores for  $GHI_{WHU}$ ,  $DNI_{WHU}$ ,  $DIF_{WHU}$ ,  $GHI_{XIA}$ ,  $DNI_{XIA}$  and  $DIF_{XIA}$  without cloud transmittances are 0.003, -4.658, -5.468, 2.452, -2.042 and 1.185, respectively; the GPI scores for  $GHI_{WHU}$ ,  $DNI_{WHU}$ ,  $DIF_{WHU}$ ,  $GHI_{XIA}$ ,  $DNI_{XIA}$  and  $DIF_{XIA}$  with cloud transmittances are 2.876, -0.923, -3.260, 4.872, 1.694 and 3.269, respectively. Table S1 show the validation results of the estimated GHI values in different CERN stations over mainland China. The estimated hourly GHI values show good agreements with the hourly GHI measurements, but with distinct spatial variations over mainland China. Relatively large model deviations are found in mountain and desert zones, due to the dramatic diurnal variations of the climate factors (water vapor, temperature, cloud and pressure etc.) there, for example the GPI scores for GSF, AKA, ALS, LAS and CHL are -9.869, -5.528, -2.685, -2.363 and -2.220, respectively.

Table S1. Validation results of the estimated hourly GHI values at CERN stations.

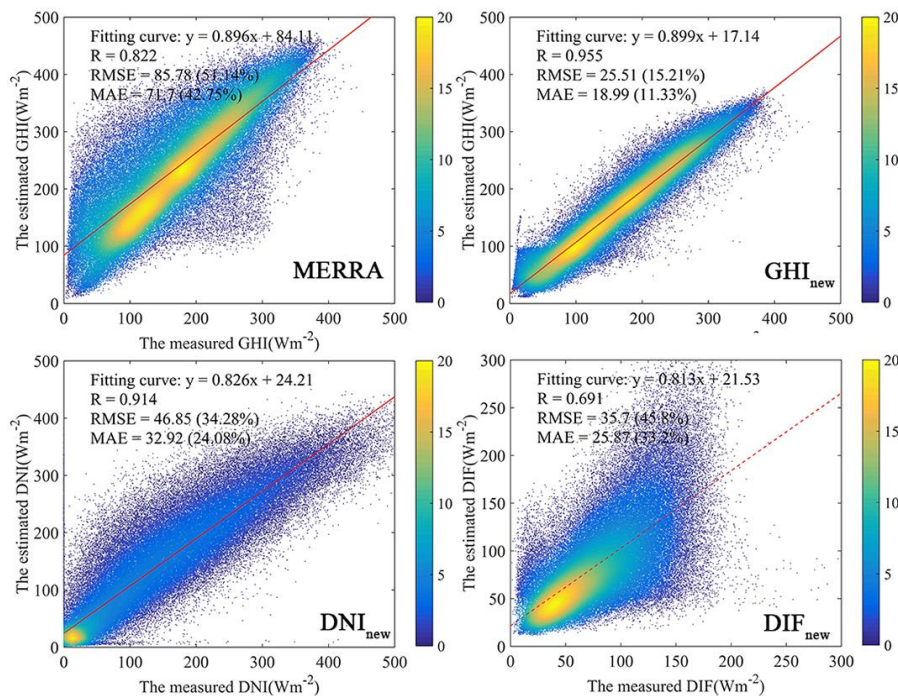


Figure 3. Validation of the daily mean GHI, DNI and DIF values at CMA stations.



Figure 3 is the scatter plot showing the model accuracy of MERRA-2 GHI records and the estimated GHI, DNI and DIF values from the REST model. Table 2 specifies the statistical indicators representing the model accuracy of the estimated GHI, DNI and DIF values. It is clear that the GHI estimations by the REST model (GHInew) show greater agreement with the measured GHI values than with the MERRA-2 GHI records. The RMSE, MAE, RMSD, MAD, MBD, SD, U95, TS, KSI, OVER and CPI (Group 1 indicators) for MERRA-2 GHI records are significantly larger than for GHInew, while the R, NSE, WIA, ICE (Group 2 indicators) for MERRA-2 GHI records are significantly lower than for GHInew. The RMSE, MAE and R for MERRA-2 GHI records are  $85.775 \text{ Wm}^{-2}$ ,  $71.696 \text{ Wm}^{-2}$  and 0.822, respectively. The RMSE, MAE and R for GHInew are  $25.505 \text{ Wm}^{-2}$ ,  $18.994 \text{ Wm}^{-2}$  and 0.955, respectively. The accuracy of GHI records is significantly improved. The DNI and DIF estimations by the REST model (DNInew and DIFnew) also show a high correlation with the ground DNI measurements. The RMSE, MAE and R for DNInew are  $46.853 \text{ Wm}^{-2}$ ,  $32.917 \text{ Wm}^{-2}$  and 0.914, respectively. The RMSE, MAE and R for DIFnew are  $35.700 \text{ Wm}^{-2}$ ,  $25.870 \text{ Wm}^{-2}$  and 0.690, respectively.

Table 2. The statistical indicators representing the model accuracy of the estimated daily GHI, DNI and DIF values.

Indicators	GHInew	DNInew	DIFnew	MERRA-2 GHI
RMSE	25.52	46.85	35.70	85.81
MAE	19.01	32.91	25.87	71.72
RMSD	15.18	34.17	42.06	36.60
MAD	11.30	24.00	30.49	30.59
MBD	-0.17	-0.32	-8.16	-28.41
SD	15.18	34.17	42.06	36.60
R	0.95	0.91	0.69	0.82
SBF	1.01	1.01	0.59	0.75
U95	42.07	94.71	116.59	101.44
TS	3.91	3.32	70.75	440.33
NSE	0.90	0.80	0.43	0.16
WIA	1.00	1.00	0.99	0.84
LCE	0.72	0.63	0.30	0.08
KSI	122.04	263.13	284.45	738.80
OVER	103.56	255.81	275.86	732.45
CPI	63.99	146.82	161.11	386.11

\* The units for RMSE and MAE are  $\text{Wm}^{-2}$ ; the units for RMSD, MAD, MBD, SD, U95, TS, KSI, OVER and CPI are %; R, NSE, WIA, ICE are dimensionless indexes.

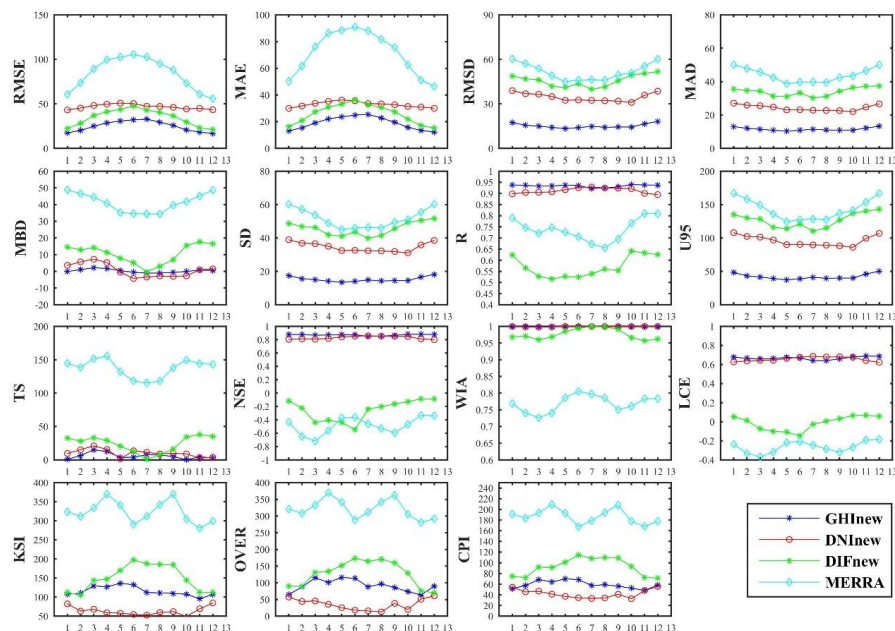


Figure 4. The statistical indicators representing the model accuracy of the estimated GHI, DNI and DIF values in different months (The units for RMSE and MAE is MJ m<sup>-2</sup>day<sup>-1</sup>; the units for RMSD, MAD, MBD, SD, U95, TS, KSI, OVER and CPI are %; R, NSE, WIA, ICE are dimensionless indexes.).

Figure 4 shows the statistical indicators representing the model accuracy of the estimated GHI, DNI and DIF values in different months. The MERRA-2 GHI records are not as accurate as GHInew in all months throughout the year. The values of Group 1 indicators for MERRA-2 GHI records are significantly larger than GHInew, while the values of Group 2 indicators for MERRA-2 GHI records are significantly lower than GHInew. The fluctuations in the values of Group 1 and Group 2 indicators for MERRA-2 GHI records in Figure 4 are also more obvious than that for GHInew, DNInew and DIFnew values, which further verified that the accuracy and robustness of GHI records are significantly improved in this study.



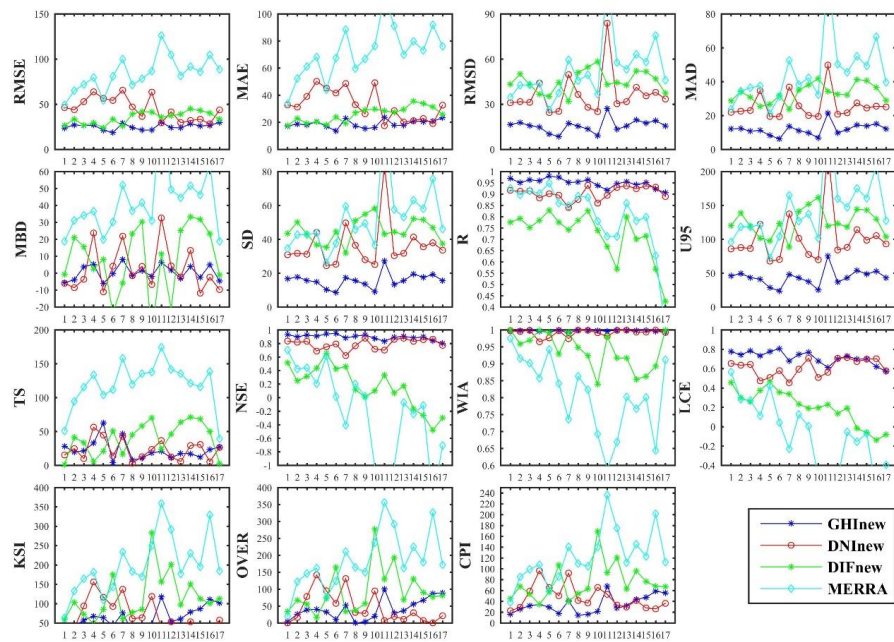


Figure 5. The statistical indicators representing the model accuracy of the estimated GHI, DNI and DIF values at 17 CMA stations (The number in x axis correspond to the ID in Table 1.).

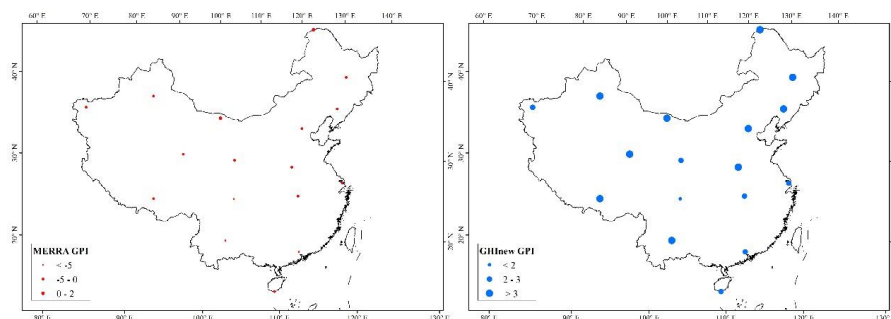
The model accuracy of the MERRA-2 GHI records and the estimated GHI, DNI and DIF values are closely correlated to local climate and terrain features. Figure 5 illustrated the statistical indicators representing the model accuracy of the estimated GHI, DNI and DIF values at 17 CMA stations. It was clear that the GHInew performance was superior to MERRA-2 GHI records with higher accuracy and robustness. The values of Group 1 indicators for GHInew were significantly lower than for MERRA-2 GHI products in all months throughout the year, while the values of Group 2 indicators of GHInew were significantly higher than those of MERRA-2 GHI products in all months throughout the year.

Furthermore, the GPI scores for the MERRA-2 GHI and GHInew values are also calculated to show the overall model estimation error at 17 CMA stations over mainland China. Figure 6 show the spatial distribution of GPI scores for MERRA-2 GHI and GHInew values at 17 CMA stations in China. The accuracy of GHInew values is obviously higher than MERRA-2 GHI records with higher GPI scores. The mean GPI scores for GHInew and MERRA-2 are 3.079 and -3.079, respectively. Relatively larger estimation errors are found in the Sichuan Basin, which may be due to the strong atmospheric radiation dumping processes there (frequent rainy and cloudy weather). The GPI scores for MERRA-2 GHI and GHInew at the Chengdu station were 1.911 and -10.329, respectively. The accuracy of MERRA-2 GHI





297 and GHInew values were higher in arid zones and plateau zones with high GPI scores due to the relative  
298 clear sky conditions there. The GPI scores for MERRA-2 GHI and GHInew at the Ejinaqi station were  
299 3.284 and 0.753, respectively. The GPI scores for MERRA-2 GHI and GHInew at the Germu station  
300 were 3.904 and -1.469, respectively.



301  
302 Figure 6. The GPI scores of the MERRA-2 GHI and GHInew values at 17 CMA stations. (The GPI is a  
303 dimensionless index.).

304 Figure 7 shows the validation results of the monthly mean MERRA-2 GHI, GHInew, DNInew and  
305 DIFnew values. It was obvious that the monthly mean GHInew, DNInew and DIFnew estimation results  
306 could meet the requirement of the potential solar energy estimations and the proper installations of solar  
307 power plants using CST with acceptable accuracy. The RMSE, MAE and R for the monthly mean  
308 GHInew estimations were  $14.745 \text{ Wm}^{-2}$ ,  $10.602 \text{ Wm}^{-2}$  and 0.973, respectively. The RMSE, MAE and R  
309 for the monthly mean DNInew estimations were  $27.778 \text{ Wm}^{-2}$ ,  $20.463 \text{ Wm}^{-2}$  and 0.922, respectively. The  
310 RMSE, MAE and R for the monthly mean DIFnew estimations were  $22.730 \text{ Wm}^{-2}$ ,  $17.690 \text{ Wm}^{-2}$  and  
311 0.798, respectively.

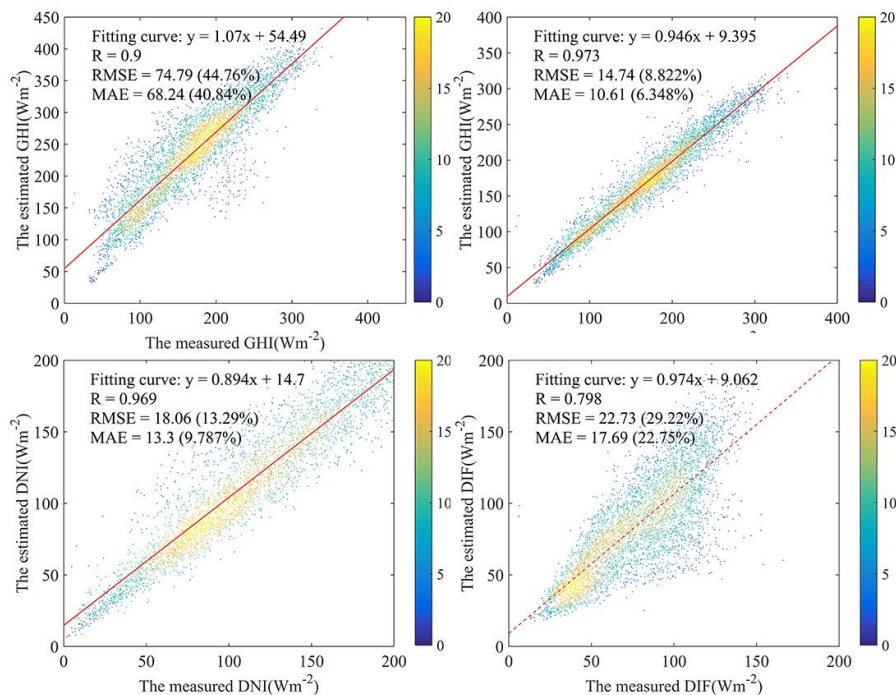


Figure 7. Validation of the monthly mean GHI, DNI and DIF values at CMA stations.

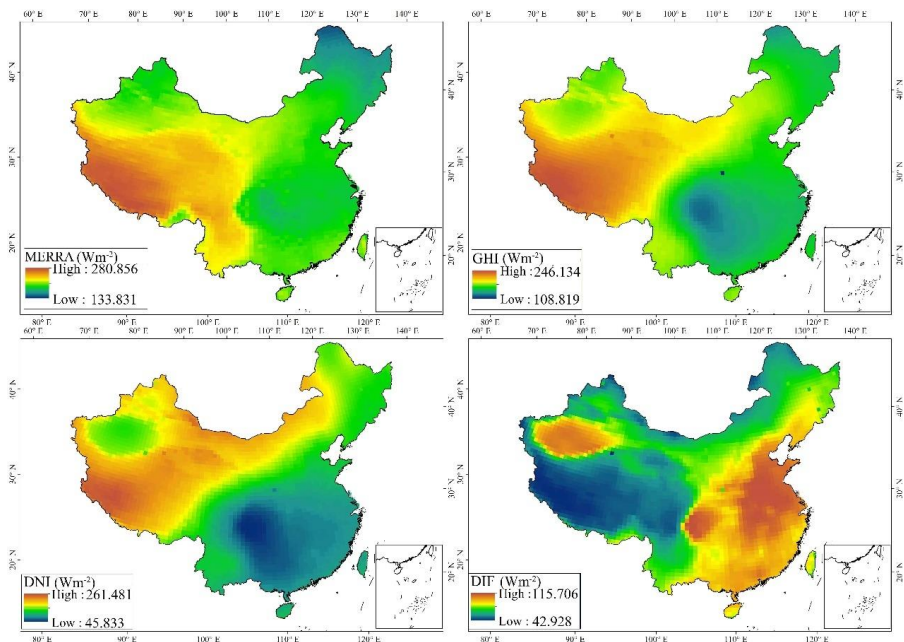
Overall, the MERRA-2 GHI products have been significantly improved in this study. Moreover, the DNI and DIF datasets during 1981-2014 were generated with acceptable accuracy, which can be used for solar energy research and applications.

### 3.3 Spatial and temporal variations in surface solar radiation

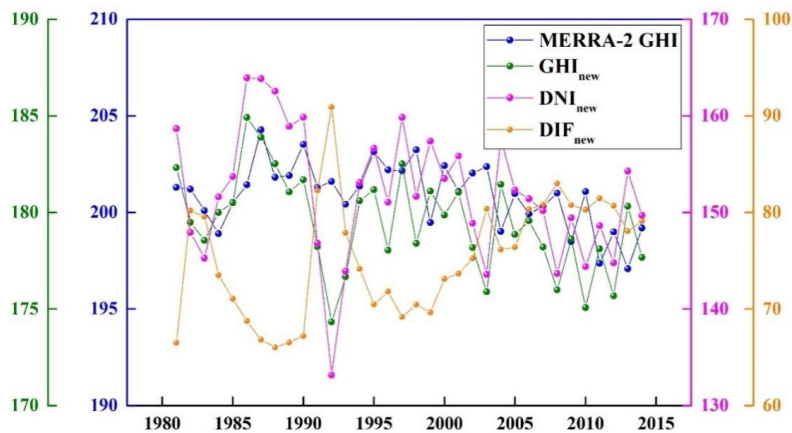
By applying the REST model, grid GHI, DNI and DIF datasets ( $0.50^\circ$  (lat)  $\times 0.625^\circ$  (lon)) during 1981-2014 throughout China are constructed. Figure 8 illustrates the spatial distributions of mean daily GHI, DNI and DIF values from 1981-2014 in China. The MERRA-2 GHI records may overestimate the GHI values in China, especially in the Sichuan Basin and Yungui Plateau, which may be due to ignoring the effect of sunshine duration and clouds. The ranges of GHI values for MERRA-2 are  $133.831 \text{ Wm}^{-2}$ - $280.856 \text{ Wm}^{-2}$ , while the ranges for GHI values by the REST models are  $108.819 \text{ Wm}^{-2}$ - $246.134 \text{ Wm}^{-2}$ . The DNI values are closely correlated to the GHI values with similar spatial distribution patterns. Generally, both GHI and DNI gradually decline from Northwestern China to Southeastern China. However, DIF show a distinct spatial distribution pattern from that of GHI and DNI. The DIF values are higher in Southeastern China and the Tarim Basin in Xinjiang Province. The Qinghai Tibetan Plateau is



328 always an area with the highest GHI, the highest DNI values and the lowest DIF values owing to the  
329 relatively weak radiation dumping effect there, while the Sichuan Basin is always an area with the highest  
330 DIF values and lowest GHI; the lowest DNI values are affected by the strong cloud cover effect.



331  
332 Figure 8. The spatial variation of GHI, DNI and DIF over mainland China (The units for GHI, DNI and  
333 DIF are  $\text{Wm}^{-2}$ ).

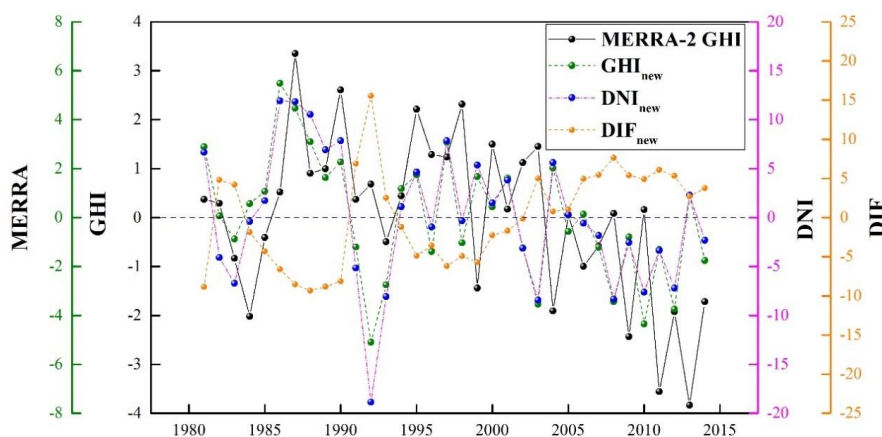


334  
335 Figure 9. The annually variations of GHI, DNI and DIF throughout China. (The units for GHI, DNI and  
336 DIF are  $\text{Wm}^{-2}$ .)

337 Figure 9 indicates the yearly variations in GHI, DNI and DIF values over mainland China during  
338 1981-2014. To better characterize the yearly variations in solar radiation in China, the anomaly GHI,



339 DNI and DIF values are calculated. Figure 10 illustrated the annual variations of the anomaly GHI, DNI  
340 and DIF values in China. The results show that the MERRA-2 GHI values obviously overestimate the  
341 GHI values in China during 1981-2014. Combined with the validation results of MERRA-2 GHI and  
342 GHInew values, we think GHInew estimations fit the measured GHI values better. It is clear from Figure  
343 9 and Figure 10 that GHI values have been gradually decreased from 1981 to 2014. The lowest annual  
344 mean GHI ( $174.329 \text{ Wm}^{-2}$ ) and anomaly GHI ( $-50.914 \text{ Wm}^{-2}$ ) value occurred in 1992, which was  
345 supposed to be caused by the strong aerosol radiative effect of volcanic eruption events in the Philippines  
346 in 1992. The DNI values are directly proportion to GHI values with similar temporal variations, because  
347 DNI is the main component of GHI values. The highest and lowest GHI values are found in 1981  
348 ( $158.657 \text{ Wm}^{-2}$ ) and 1992 ( $133.137 \text{ Wm}^{-2}$ ), respectively. In contrast, DIF values show an opposite  
349 temporal variation with GHI and DNI values. The DIF values have been gradually increasing from 1981  
350 to 2014, especially in 1992 ( $\text{DIF} = 90.914 \text{ Wm}^{-2}$ , anomaly DIF value =  $15.544 \text{ Wm}^{-2}$ ) with an explosive  
351 growth of DIF values. It is thought that DIF values are directly proportional to AOD value.



352  
353 Figure 10. The annually variations of anomaly GHI, DNI and DIF throughout China (The units for  
354 anomaly GHI, DNI and DIF are  $\text{Wm}^{-2}$ ).

#### 355 4 Discussion

356 The model accuracy of DNInew and DIFnew is relatively lower than that of GHInew, which may  
357 be caused by three factors. First, there are too many low DNI and DIF values, because low DNI values  
358 generally correspond to cloudy sky conditions, which may cause large uncertainties in DNI estimations.  
359 Second, although they were the only available sunshine duration datasets with the highest density in



China, the point density of sunshine duration measurements at 2474 CMA stations was still sparse. Finally, the spatial resolution of the input parameters derived from MERRA-2 products and the output values (DNI<sub>new</sub> and DIF<sub>new</sub>) were  $0.50^\circ$  (lat)  $\times 0.625^\circ$  (lon), which may also degrade the accuracy of the estimated DNI and DIF values.

Table 3. Validation of GHI estimations in China in previous studies.

Parameters	Models/ Products	Format	Spatial resolution	Spatio-temporal continuity	RMSE (Wm <sup>-2</sup> )	R
GHI	EPP	Grid	$1^\circ \times 1^\circ$	Few vacancies	34.028	0.930
GHI	ISCCP-FD	Grid	$1^\circ \times 1^\circ$	High continuity	35.764	0.910
GHI	GEWEX-SRB	Grid	$1^\circ \times 1^\circ$	High continuity	34.144	0.930
GHI	EPP-TANG	Grid	5km	Few vacancies	33.912	0.930
GHI	GLASS	Grid	$0.05^\circ \times 0.05^\circ$	High continuity	34.144	0.930
GHI	MERRA	Grid	$0.625^\circ \times 0.5^\circ$	High continuity	85.775	0.822
GHI	GHI <sub>new</sub>	Grid	$0.625^\circ \times 0.5^\circ$	High continuity	25.509	0.955

The estimated GHI values in this study were compared with other estimation results in previous studies. The validation of DNI and DIF estimations in previous studies is not listed and discussed, because the evaluation of DNI and DIF values over mainland China is not founded in previous studies. Table 3 shows the validation results of GHI estimations in China in this study and previous studies. Detailed descriptions of these models and products have been described in the Introduction Section. It was clear that the surface GHI estimation results in this study show higher agreement with surface solar radiation measurements at CMA stations than those of other estimation results in previous studies, which may be due to the consideration of cloud effects on GHI. Although the spatial resolution of GHI estimations by EPP-TANG is higher than GHI<sub>new</sub>, high spatiotemporal continuity and long temporal ranges of the GHI<sub>new</sub> estimation results could remedy the defect of relatively lower spatial resolutions.

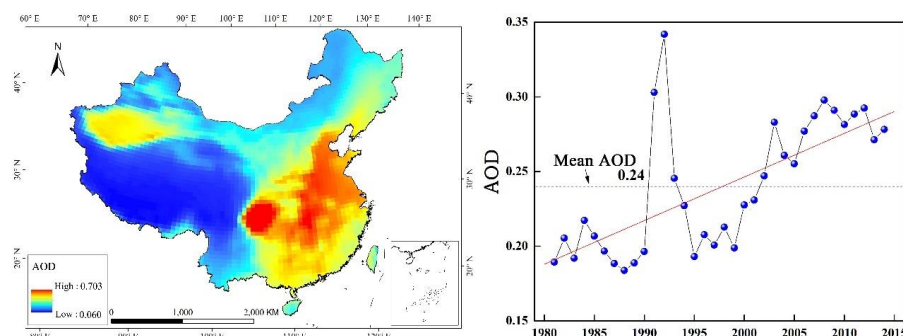


Figure 11. The spatial and temporal variations of AOD values during 1981-2014 throughout mainland China.

It was supposed that the DIF values were strongly correlated to aerosol optical depth (AOD) in China. Thus, we introduced the MERRA-2 AOD product to analyze the correlation between DIF values and AOD values in China. Figure 11 show the spatial distribution of AOD values over mainland China. As seen from Figure 8, Figure 9 and Figure 11, DIF values show similar spatial distribution patterns with AOD values in China. The correlation coefficient between the annual mean DIF and annual mean AOD values is 0.890. The Sichuan Basin is the area with the highest AOD (0.703) and the highest DIF ( $115.706 \text{ Wm}^{-2}$ ) values, while the Qinghai Tibetan Plateau is the area with the lowest AOD (0.060) and the lowest DIF ( $42.928 \text{ Wm}^{-2}$ ) values. It is certain that AOD is an import factor in the DIF variations in China.

## 5 Data availability

The MERRA-2 Reanalysis data are available at GES DISC by NASA (<https://disc.gsfc.nasa.gov/daac-bin/FTPSubset2.pl>). We provide these solar irradiance data in publicly available repository: <https://doi.org/10.6084/m9.figshare.10026563> (Qin, W. et al., 2019). The corresponding author can be contacted for access meteorological data at CMA stations and solar irradiance dataset generated in this study during 1981-2014 as well as ancillary data.

## 6 Summary

The applicability of REST2 in modeling GHI, DNI and DIF values using MERRA-2 reanalysis products ( $AOD_{550}$ ,  $p$ ,  $rog$  and  $w$ ) and sunshine duration measurements at 2474 CMA stations throughout China was tested in this study. Long-term grid GHI, DNI and DIF datasets ( $0.50^\circ$  (lat)  $\times 0.625^\circ$  (lon)) throughout China were then constructed. Finally, the spatiotemporal characteristics of GHI, DNI and DIF



397 in China were investigated.

398 The estimated SI values show high agreements with SI measurements at 17 CMA stations with  
 399 radiation measurements. Eighteen indicators including RMSE, MAE, RMSD, MAD, and MBD were  
 400 used to represent the model accuracy of the MERRA-2 GHI, GHInew, DNInew and DIFnew values. The  
 401 RMSE for MERRA-2 GHI, GHInew, DNInew and DIFnew are 85.775, 25.509, 46.852 and 35.700 Wm<sup>-2</sup>,  
 402 respectively; the MAE are 71.701, 18.993, 32.917 and 25.870 Wm<sup>-2</sup>, respectively; and the R are 0.822,  
 403 0.955, 0.914 and 0.691, respectively. It could be concluded that the accuracy of MERRA-2 GHI values  
 404 has been significantly improved in this study. Relatively large estimation errors for MERRA-2 GHI and  
 405 GHInew values are found CHD in Sichuan Basin with GPI scores of 1.911 and -10.329, respectively,  
 406 because of the cloud air conditions there.

407 The spatiotemporal characteristics of GHI, DNI and DIF values from 1981-2014 over mainland  
 408 China were discussed using the generated grid GHI, DNI and DIF datasets in this study. The results show  
 409 that the MERRA-2 GHI records may overestimate the GHI values over mainland China. Generally, the  
 410 GHI and DNI values have gradually decreased from 1981-2014. However, DIF values have gradually  
 411 increased from 1981 to 2014, especially in 1992 (DIF = 90.914 Wm<sup>-2</sup>, anomaly DIF value = 15.544 Wm<sup>-2</sup>,  
 412 which may be caused by the increasingly strong aerosol radiative forcing effects throughout China  
 413 during 1981-2014. The Qinghai Tibetan Plateau has always been the area with the highest GHI, highest  
 414 DNI and lowest DIF values (clear sky condition), while the Sichuan Basin has always been the area with  
 415 the lowest GHI, lowest DNI and highest DIF values (cloudy and rainy sky condition). It was validated  
 416 that the DIF values are strongly correlated with aerosol optical depth (AOD) in China.

417 Certainly, the REST2 model should be further validated in other climate zones around the world. As  
 418 discussed above, the GHI, DNI and DIF estimations are subject to input data quality, the interpolated  
 419 method and the relatively coarse resolution of MERRA-2 products. Further work should be conducted  
 420 to improve the accuracy of the GHI, DNI and DIF datasets generated in this study. Moreover, significant  
 421 relations between DIF and the AOD values are validated in this study, and further studies should be  
 422 undertaken to reveal the main driving factors for the spatio-temporal variations in GHI, DNI and DIF  
 423 values.

## 424 Acknowledgements

425 The surface solar radiation (GHI, DNI and DIF values) data records at 17 sites of China were





obtained from National Meteorological Information Center of China <http://data.cma.cn>; MERRA-2 (Modern-Era Retrospective Analysis for Research and Applications, version 2) products were available from Goddard Space Flight Center Level 2 and Atmosphere Archive Distribution System <http://ladsweb.nascom.nasa.gov>. This work was financially supported by the National Natural Science Foundation of China (No.41601044), the Special Fund for Basic Scientific Research of Central Colleges, China University of Geosciences, Wuhan (CUGCJ1704, 007-G1323519253 and 111-162301182738), and the 111 Project (grant No. B08030). We would like to thank Professor Christian A. Gueymard for providing the code of the REST2 software. We would like to thank the China Meteorological Administration (CMA) for providing the meteorological and radiation data.

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